

# Broadband microwave emissions from LWIR picosecond laser ablation with pre-ionization

#### **Proposal #312114**

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Funding: ONR (JHU/APL, received) AFRL Program Funds (AFRL, received)

#### **Broadband emissions from pulsed laser ablation**

Extreme nonlinear optics of ultra-short pulse lasers enables research and applications across the entire electromagnetic spectrum...

Radio	Microwave	Infrared	<mark>∨is</mark>	UV	X-Ray	Gamma
RF	EMP	Parametric Conversion		Hig G	h Harmonic eneration	
60	Terahert Sources	z S Su	Optica percontir	nuum	Acc li	elerator-based
40 20 0 -40 -40 -1 -0.5 0	0.5 1 1.5 2 2.5 time [ns]				OPN News, Laser Wak Generation Light Source	Kefield Accelerators: Next- ces," Albert 2018.

# **RF Emissions From Short Pulse Laser Ablation**

- Low-frequency (RF) emissions measured from ablation as early as 1970's
- Typically found in large laser facilities with lasers reaching relativistic intensities
- Recent work has investigated "atmospheric" and relatively low intensity sources of EMP from laserplasma interactions
- Three main sources of low frequency emission
  - "Rectification" physics
  - Long scale-length plasma currents
  - Target charging and neutralization



# Wavelength Scaling into the LWIR



λ (microns)	Energy (joules)	Duration (fs)	Laser Parameters: $\lambda$ ( $\mu$ m), Trradiance (x10 <sup>11</sup> W/cm <sup>2</sup> )
0.8	.04	50	-0.8, 236.1
1.0	.03	1300	-1.0, 21.1
9.2	3	2000	<b>—9.2</b> , <b>48.5</b>
10.6	.01	3000	—10.6, 3.1

• FY22 results exceeded expectations!

 Orders of magnitude improvement across all frequencies achieved with BNL laser



### **Target pre-ionization**

- Previous work at JHU/APL showed orders of magnitude increase in RF emissions from pre-ionized dielectric targets with NIR drive lasers
- Hypothesize improved laser absorption at the critical density layer in the expanding plasma causes a higher electron temperature



S. Varma, J. Spicer, B. Brawley, and J. Miragliotta, "Plasma enhancement of femtosecond laser-induced electromagnetic pulses at metal and dielectric surfaces," *Opt. Eng.*, vol. 53, no. 5, p. 051515, 2014.



# **Key experimental questions**

- How does RF emission change with laser incidence angle and intensity? Can we match theoretical models?
  - How does emissions strength compare between dielectric and metal targets?
- Does pre-ionization increase the RF emission from solid targets similar to what we see in the NIR?
  - What pulse separation time scale improves RF emission?
  - Does pre-ionization improve LWIR absorption by the plasma?
- Can we detect a temperature electron population through keV bremsstrahlung? Does the electron temperature correlate with RF emission?
- What effects does the emitted ultrawideband RF field have on electronic devices?



X-ray emission spectrum measured from USPL ablation of copper wire in air

#### **Diagnostics**

- RF collection equipment
  - High speed oscilloscope(s) and broadband horn antenna(s)
- Shadowgraph imaging of shock dynamics
  - mW-class laser diode and gated ICCD
  - Shock evolution allows estimate of laser absorption
- Pump reflectometer
  - Power meter based reflectance measurement
- X-ray spectrometer (preferably filter based)
  - Detects bremsstrahlung emission from hypothesized high temperature electrons
- "DUT" electronics
  - Field effects on transistors, etc.









# **Funding / Timeline**

- AFRL and JHU/APL programs are funded through FY23
- Lead times for experimental components push likely timeline to summer 2023
- Proposed 1 year effort 2 to 3 weeks of beam time
  - Week 1 Set up dual pulse pre-ionization formats. Check initial RF signal levels and begin setup of diagnostics.
  - Week 2 Scan pulse delay with spectrometer and probe pulse diagnostics to verify optimal beam timing and overlap. Study RF emission as a function of incident laser properties.
  - Week 3 Test electronic device response to high field RF.
- Potential FY24 proposal for more advanced plasma diagnostics (e.g. spectral interferometry on reflected and/or transmitted femtosecond probe beam)

Positive results likely to drive continued interest in ultrafast CO2 laser development from DoD sponsors



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#### **Electron Beam Requirements**

No electron beam time requested

#### CO<sub>2</sub> Laser Requirements

Configuration	Parameter	Unit s	Typical Values	Comments	Requested Values
CO <sub>2</sub> Regenerative Amplifier Beam	Wavelength	μm	9.2	Wavelength determined by mixed isotope gain media	
	Peak Power	GW	~3		
	Pulse Mode		Single	<ul> <li>Regen beam needed for beam particular</li> </ul>	th alignment.
	Pulse Length	ps	2	<ul> <li>Request access to regen output to</li> </ul>	o install
	Pulse Energy	mJ	6	Mach-Zehnder style pulse splitter	(2 pulses,
	M <sup>2</sup>		~1.5	provided by JHU/APL)	
	Repetition Rate	Hz	1.5	3 Hz also available if needed	
	Polarization		Linear	Circular polarization available at slightly reduced power	
CO <sub>2</sub> CPA Beam	Wavelength	μm	9.2	Wavelength determined by mixed isotope gain media	9.2
Note that delivery of full power pulses to the Experimental Hall is presently limited to Beamline #1 only.	Peak Power	TW	5	~5 TW operation will become available shortly into this year's experimental run period. A 3-year development effort to achieve >10 TW and deliver to users is in progress.	0.5 – 2
	Pulse Mode		Single	JHU/APL to install a pulse splitter for some experiments	Single / Double
	Pulse Length	ps	2		2ps
AL	Pulse Energy	J	~5	Maximum pulse energies of >10 J will become available within the next year	1 – 4

#### Other Experimental Laser Requirements

Ti:Sapphire Laser System	Units	Stage I Values	Stage II Values	Comments	<b>Requested Values</b>
Central Wavelength	nm	800	800	Stage I parameters are presently available and setup to deliver Stage II parameters should be complete during FY22	800nm
FWHM Bandwidth	nm	20	13		13
Compressed FWHM Pulse Width	fs	<50	<75	Transport of compressed pulses will initially include a very limited number of experimental interaction points. Please consult with the ATF Team if you need this capability.	75
Chirped FWHM Pulse Width	ps	≥50	≥50		
Chirped Energy	mJ	10	200		
Compressed Energy	mJ	7	~20	20 mJ is presently operational with work underway this year to achieve our 100 mJ goal.	20+
Energy to Experiments	mJ	>4.9	>80		10+
Power to Experiments	GW	>98	>1067		
System	Units	Typical Values		Comments	<b>Requested Values</b>
Wavelength	nm	1064	Single p	oulse	
Energy	mJ	5			
Pulse Width	ps	14			
Wavelength	nm	532	Freque	ncy doubled	
Energy	mJ	0.5			
Pulse Width	ps	10			

#### **Special Equipment Requirements and Hazards**

- Electron Beam
  - N/A
- CO<sub>2</sub> Laser
  - 'Single beam' pre-ionization scheme will require installation of a pulse stacker between the CO2 regenerative amplifier and the main power amp.
  - Experiments performed in air will require disassembly of part of the beam routing vacuum tube. We anticipate using the existing 80/20 stands to help support our experimental diagnostics and alleviate space constraints. Setup will require at least 4' x 1.5' space, preferably a little more.
- Ti:Sapphire and Nd:YAG Lasers
  - Requesting use of the Ti:Sapphire beam for probe and/or target pre-ionization. Beam will need to be routed to our experimental area. Experiments at JHU/APL ongoing to determine if compressed vs uncompressed beam is needed.
- Hazards & Special Installation Requirements
  - Large installation (chamber, insertion device, etc.): Disassembly of vacuum beam line to accommodate experiment in ambient environment
  - Cryogens: N/A
  - Introducing new magnetic elements: N/A
  - Introducing new materials into the beam path: CO2 pulse stacker between regen and main amp (ZnSe optics)
  - Any other foreseeable beam line modifications: No

#### **Experimental Time Request**

#### CY2023 Time Request

Capability	Setup Hours	Running Hours
Electron Beam Only		
Laser* Only (in Laser Areas)	0.5 – 1 week	2 weeks
Laser* + Electron Beam		

#### Total Time Request for the 3-year Experiment (including CY2023-25)

Capability	Setup Hours	Running Hours
Electron Beam Only		
Laser* Only (in Laser Areas)	0.5 – 1 week	2 weeks
Laser* + Electron Beam		

\* Results may warrant follow-on FY24/25 proposals if funding is secured

\* Laser = Near-IR or LWIR ( $CO_2$ ) Laser



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