

(a) Title: Numerical Methods for the Determination of Mixing

(b) Names and Affiliation of Authors

S. Dutta E. George J. Glimm X. L. Li A. Marchese Z.-L. Xu Y.-M. Zhang

Department of Applied Mathematics and Statistics

Stony Brook University, Stony Brook, NY 11794-3600

J. W. Grove

Computer and Computational Science Division

Los Alamos National Laboratory Los Alamos, New Mexico 87545

David H. Sharp

Theoretical Division, Los Alamos National Laboratory

Los Alamos, New Mexico 87545

© Corresponding Author:

XiaoLin. Li

Email address: linli@ams.sunysb.edu

Phone: (631)-632-8353, Fax: (631)-632-8490

Department of Applied Mathematics and Statistics

Stony Brook University, Stony Brook, NY 11794-3600

(d) Address for reprints: same as above

(e) Short Title: Numerical Methods for Mix

(f) Number of manuscript pages: 24

Number of tables: 1

Number of figures: 7

Abstract:

Numerical Methods for the Determination of Mixing

Rayleigh-Taylor mixing rate is studied through the front tracking simulation; the result shows that the acceleration rate falls within the range of experiments. The front tracking method prevent interfacial mass diffusion. An analysis is presented to support the assertion that the lower acceleration rate found in untracked simulations is caused, at least to a large extent, by a reduced buoyancy force due to numerical mass diffusion across the interface. Quantitative evidence includes results from a time dependent Atwood number analysis of the diffusive simulation, which yields a renormalized mixing rate coefficient for the diffusive simulation in agreement with experiment. The Richtmyer-Meshkov instability is also studied through the front tracking method in spherical geometry. The growth rate of fingers at an unstable shell driven by an imploding spherical shock is studied through simulations; a qualitative understanding of this system has been achieved.

Keyword: Front Tracking, Rayleigh-Taylor instability, Richtmyer-Meshkov instability, numerical diffusion.

1. Introduction

Since Read and Youngs [12,14] published the first experimental study of Rayleigh-Taylor instability with a randomly perturbed fluid interface, attention has been drawn to the non-dimensional acceleration rate of the bubble envelope. Assume that two fluids are separated by a randomly perturbed interface and that the gravitational field points from the heavy fluid ρ_H to the light fluid ρ_L . Read and Youngs confirmed the Sharp-Wheeler theoretical prediction [13] that the average bubble front moves with acceleration scaling

$$h(t) = \alpha A g t^2 \quad [1]$$

where h is the height of the bubble envelope, $A = (\rho_H - \rho_L)/(\rho_H + \rho_L)$ is the Atwood number, g is the gravity and t is time. Read and Youngs show that the acceleration rate α is almost a constant, with $\alpha \approx 0.063 - 0.077$ in 3-D experiments. The experiments have been repeated by various authors with different apparatus, and similar values of α have been obtained; we mention the experiments of Dimonte and Schneider [4,5,6], giving $\alpha = 0.05 \pm 0.01$. The theoretically determined value, $\alpha \approx 0.05 - 0.06$, is obtained from a bubble merger renormalization group model. See [3] and references cited there. Computation of the center of mass of the mixing zone introduces a coupling between its two edges. Therefore, a characterization of the center of mass (nearly stationary for $A \leq 0.8$, for example) determines the total mixing zone size in terms of α alone [1,2].

The coefficient α is thus important. It characterizes the size of the mixing zone, and thus largely determines the amount of material which is mixed. It has been reported

by experimentalists as being approximately universal, in the sense that it is nearly independent of the random initial conditions of an experiment.

Researchers in several laboratories have tried to reproduce the Sharp-Wheeler theoretical scaling law with the experimental value for α through numerical simulation. Most researchers report a time dependent, decreasing value for α , ranging from 0.015 to 0.03.

These simulations are from computational codes using numerical schemes with interfacial mass diffusion. We have compared numerical simulations using a high resolution front tracking code *FronTier* with zero interfacial mass diffusion to our own simulations using an untracked TVD-level set code with interfacial mass diffusion similar to the others. We also introduce an analytic study of the effects of mass diffusion on buoyancy reduction and we predict the numerically observed reduction in α for untracked simulations. Our main result is that all values of α (theory, experiment, simulation) are consistent if the diffusive calculation of α is renormalized to account for mass diffusion.

We also performed simulations of the Richtmyer-Meshkov instability in spherical geometry using a 2-D cylindrical front tracking code. In addition to the usual difficulties for RM instability in planar geometry, spherical geometry complicates the RM instability due to the curved geometry in three dimensions. In order to achieve a stable spike and bubble velocity, we used the following methods. For the boundary conditions, we use a reflecting boundary condition at $r = r_{\min} > 0$, and Dirichlet flow through conditions on the other three boundaries. The idea of a flow through

boundary condition can be briefly described as follows. In order to update the boundary state, we imagine there exists a far field state which is defined by the extrapolation of the states near the boundary point. Then we solve the Riemann problem using the interior state near the boundary and the extrapolated far field state and retain only the incoming waves. A flow through boundary is thus time-dependent. It is equivalent to an open boundary and an infinitely large domain. All waves will propagate out of the domain and no boundary signal will be reflected back at least theoretically.

2. Diffusive and Nondiffusive Simulations

An earlier comparison shows that *FronTier* simulations produce values for α close to agreement with experiment while untracked TVD simulations produce low values for α [8]. These comparisons were limited in the simulation time and in the penetration depth of mixing achieved. Here we extend the comparison to a later time, comparable to most other simulations. Fig.1 shows the evolution of the fluid interface in the *FronTier* simulation. The color coding displays the height through the mixing zone, and the cut plane near the bubble surface at the top of the right frame shows the location of the 5% volume fraction contour for the light fluid. Note that there are a number of light fluid bubbles at the later time. The dynamics is multimode, not dominated by a single large space filling bubble up to this time. Such a large bubble would indicate the end of any possible self similar flow regime, as the acceleration scaling depends upon a continued growth in the transverse scale of the mixing structures. We expand on this idea. The dynamics of continued acceleration of the mixing zone edge, as expressed in the t^2 scaling in Eq. 1, depends on a continued

growth of the large scale structures (the bubbles). See for example [3,13]. Bubbles grow through a process of bubble competition and merger. Thus the t^2 scaling and the determination of α requires a simulation which is still in the multimode regime, where bubble competition and merger can occur.

The $t = 0$ interface is constructed out of Fourier modes with random amplitudes and frequencies in the range of 8 to 16 modes per computational domain width. See [8] for further information concerning these simulations. The $2 \times 2 \times 8$ computational domain used here allows computationally efficient late time, deep penetration simulations.

Within this computational aspect ratio, the Fourier mode numbers represent a balance between the conflicting requirements of spatial resolution, favoring low numbers of modes, and late time statistical validity, favoring large numbers of modes. Except as noted, the simulations used a $128^2 \times 512$ grid. Our simulations, thus balanced, have about $12^2 = 144$ initial bubbles and a grid resolution of about ten cells in each dimension per initial bubble. The final time considered here has about five bubbles (see Figure 1).

[Figure 1 goes here]

A comparison of the mixing rates for the two simulations is shown in Figure 2 (left), plotting bubble height h vs. $Ag t^2$. *FronTier* has a distinctly higher growth rate than does the interface mass diffusive TVD simulation. The value of $h(t)$ is the difference between the $t=0$ bubble height and the time t bubble height. The latter quantity is defined in terms of a 1% volume fraction, i.e., the greatest height at which the

fluid is 99% heavy and 1% light according to the front tracking front or the TVD level set. This definition is somewhat unstable statistically, and a few spurious oscillations associated with the definition were removed in the plots of Figure 2.

[Figure 2 goes here]

Mass diffusion is a common feature of most untracked simulation codes. Due to the interpolation constraint, numerical schemes (finite difference, finite volume) can have only first order accuracy in their spatial derivatives near a discontinuity. For a contact discontinuity, the corresponding characteristic is linear for the wave equation of the Riemann invariant

$$\frac{\partial w}{\partial t} + u \frac{\partial w}{\partial x} = 0, \quad \text{where } w = \rho - \frac{P}{c^2}, \quad [2]$$

and so the truncation error will spread to the interior region. Assuming that a finite difference scheme is second order in time and first order in space at a contact discontinuity, we have the equivalent equation

$$w_t + uw_x = \frac{1}{2} \Delta x w_{xx} \quad [3]$$

so that the width ΔL of the numerically diffused density profile satisfies $\Delta L \sim \sqrt{\Delta x t}$.

In order to understand the difference between the two simulations, we compare the cross sectional density plots in a series of horizontal slices from the bubble (upper) portion of the mixing region. Figure 3 shows the cross sectional density plots in these simulations. Observe that there is a substantial smearing-out of the density across the boundary between the two fluids in the untracked TVD simulation, while the *FronTier* simulation maintains a sharp boundary with a discontinuous density profile

throughout the simulation. As a further difference, we note the fine scale structure size in the *FronTier* simulation in comparison to the TVD simulation.

We compute an effective Atwood number $A(t)$ as a function of time for the TVD simulations. This is determined from the highest and lowest densities in a horizontal slice, with the resulting time and space dependent Atwood number averaged over heights in the upper third of the mixing zone at a fixed time to get an Atwood number dependent on time alone. In Figure 4, we plot $A(t)$ vs. t for three simulations (fine and coarse grid TVD and fine grid *FronTier*). The time dependence of $A(t)$ in the *FronTier* simulation is caused purely by (small) compressibility effects. For the mass diffusive TVD simulation, the initial density contrast, $A(t = 0) = 0.5$, is almost completely washed out; the earliest time displayed shows $A(t = 2) \approx 0.15$. As new pure (heavy and light) fluid is injected into the mixing region, the effective Atwood number increases, but it is still reduced to about $A \approx 0.3$ on a time averaged basis, or nearly a 50% reduction relative to its initial value.

To compensate for the time dependent Atwood number $A(t)$, we define an effective alpha, $\alpha_{eff} \approx h/2 \iint A(s)gdsdt$ (see Figure 2, right). Specifically, α or α_{eff} is defined here as the slope of the straight line joining the beginning and end of the $h(t)$ curve in Figure 2. This definition, although somewhat arbitrary, is conventional, and thus allows comparison to the results of others. We observe an improved comparison between *FronTier* and TVD and between TVD and experiment. Note that α_{eff} lies within the range of experimental values; see Table 1. On this basis, we can state that the diffusive buoyancy renormalization of α is capable of resolving existing

discrepancies among simulations, between diffusive simulations and nondiffusive experiments, and with theory.

[Table 1 goes here]

[Figure 3 goes here]

[Figure 4 goes here]

3. Diffusion Induced Buoyancy Reduction

The reduced mixing rate due to unphysical numerical diffusion can be understood from Figure 5. The left frame represents an immiscible bubble of radius r . The central and the right frame assume that this bubble is smeared out numerically to a radius R while the total mass inside the sphere of radius R is conserved. The buoyancy forces

$$f_1 = f_2 = \frac{4\pi r^3}{3}(\rho_H - \rho_L)g \quad [4]$$

for the bubbles in frames (a) and (c) are the same. However, due to the difference between the mass in the nondiffused bubble (a) and the diffused bubble (c), the two acceleration rates

$$a_1 = \frac{\rho_H - \rho_L}{\rho_L} g > a_2 = \frac{\rho_H - \rho_L}{\rho_L + \left(\frac{R^3}{r^3} - 1\right)\rho_H} g \quad [5]$$

are different.

As a result of the mass diffusion, the buoyancy force is distributed to a larger amount of mass, thus reducing the acceleration of the bubble.

[Figure 5 goes here]

4. Richtmyer-Meshkov Instability

We have performed verification and validation studies for axisymmetric simulations using *FronTier*. The validation was through comparison to laser driven hemispherical targets [7], and will not be repeated here. In Fig. 6, we present the results of a mesh refinement study for an axisymmetrically perturbed spherical Richtmyer-Meshkov problem, comparing the growth rates as a function of time for a 200×200 and a 400×400 mesh. The influence of the symmetry axis (the "North Pole" effect) on the statistical characterization of the instability evolution was studied [9,10,11], with the main conclusions that (1) the effect was a real consequence of axisymmetrically perturbed flows, *i.e.* it was not due to numerical effects, (2) it was independent of spherical flow geometry, and arises in cylindrically shaped flows, (3) that the effect occurs late in time and for spherical flows, is pronounced after reshock, and (4) that the effect is not eliminated through use of spherical harmonic (Legendre polynomial) perturbations.

We present in Fig.7 the results of a strong shock scaling law analysis of the mixing zone growth rate for a spherically perturbed Richtmyer-Meshkov problem. The perturbed spherical surface separates a heavy gas (on the interior) from a light gas (on the exterior), with the initial shock location in the heavy fluid, facing outwards (explosion). Following [15], where a similar scaling law was introduced for cylindrical implosions, we scale the velocity and times by the incident shock Mach number M , introducing a scaled velocity $v' = v/M$ and time $t' = Mt$. The results of the scaling show a very close near identity of growth rate curves, which is remarkable

in view of the large amount of structure in the curves themselves. Again the configuration is heavy exploding light.

[Figure 6 goes here]

[Figure 7 goes here]

5. Conclusion

We present a *FronTier* simulation run to late time and deep penetration. The simulation is terminated while still in a multimode regime. It has no interfacial mass diffusion, and the overall bubble mixing rate lies within the experimental range. We recalibrate the buoyancy force for mass diffusive TVD simulations, and obtain a renormalized α_{eff} which is also in agreement with experiment. On this basis, the nondiffusive simulation and the theory of mass diffused buoyancy reduction presented here are capable of resolving the principal differences between simulation and experiment for Rayleigh-Taylor mixing. Our results confirm the earlier agreement between theory and experiment [3]. Finally, we observe that our results open a door to further research, and do not close inquiry related to the determination of the mixing rate, as the uncertainties in the experimental, theoretical and simulation determination of α deserve further investigation. Concerning simulation, which is the main thrust of this paper, we mention the importance of improved resolution. The needs for resolution are numerical accuracy, governed by mesh cells per bubble, statistical accuracy, governed by the number of bubbles, especially at the end of the simulation, and convergence to self similar flow, governed by the length of time of the simulation.

6. References.

1. B. Cheng, J. Glimm, D. Saltz, and D.H.Sharp. Boundary conditions for a two pressure two phase flow model. *Physica D*. 133:84-105, 1999.
2. B. Cheng, J. Glimm, and D.H.Sharp. Density dependence of Rayleigh-Taylor and Richtmyer-Meshkov mixing fronts. *Phys. Lett. A*, 268: 366-374, 2000.
3. B. Cheng, J. Glimm, and D.H.Sharp. A 3-D RNG bubble merger model for Rayleigh-Taylor mixing. *Chaos*, to appear, 2001. SUNYSB-AMS-01-04 Stat Univ. of New York at Stony Brook.
4. G. Dimonte. Nonlinear evolution of the Rayleigh-Taylor and Richtmyer-Meshkov instabilities. *Phys. Plasmas*, 6(5): 2009-2015, 1999.
5. G. Dimonte and M. Schneider. Turbulent Rayleigh-Taylor instability experiments with variable acceleration. *Phys. Rev. E*, 54:3740-3734, 1996.
6. G. Dimonte and M. Schneider. Density ratio dependence of Rayleigh-Taylor mixing for sustained and impulsive acceleration histories. *Phys. Fluids*, 12:304-321, 2000.
7. R.P. Drake, H.F. Robey, O.A. Hurricane, B.A. Remington, J. Knauer, J. Glimm, Y. Zhang, D. Arnett, D.D. Ryutov, J.O. Kane, K.S. Budil, and J.W. Grove. Experiments to produce a hydrodynamic unstable spherically diverging system of relevance to instabilities in supernovae. *Astrophysical Journal*, 564:869-906, 2002.
8. J. Glimm, J.W. Grove, X.L.Li, W.Oh, and D.H. Sharp. A critical analysis of Rayleigh-Taylor growth rates. *J. Comp. Phys.* 169: 652-677, 2001.
9. J. Glimm, J.W. Grove, and Y. Zhang. Three dimensional axisymmetric simulations of fluid instabilities in curved geometry. In M. Rahman and C. A.

Brebbia, editors, *Advances in Fluid Mechanics III*, pages 643-652. WIT Press, Southampton, UK, 2000.

10. J. Glimm, J. W. Grove, and Y. Zhang. Interface tracking for axisymmetric flows. *SIAM J. SciComp*, 2002 (Accepted for publication). LANL report NO. LA-UR-01-448.
11. J. Glimm, J. W. Grove, Y. Zhang, and S. Dutta. Numerical study of axisymmetric Richtmyer-Meshkov instability and azimuthal effect on spherical mixing. *J. Stat. Physics*, 2001, (accepted for publication). LANL report NO. LA-UR-01-2299.
12. K. I. Read. Experimental investigation of turbulent mixing by Rayleigh-Taylor instability. *Physica D*, 12:45-58, 1984.
13. D. H. Sharp and J. A. Wheeler, Late stage of Rayleigh-Taylor instability. Technical report, Institute of Defense Analyses, 1961. Unpublished Technical Report.
14. D.L.Youngs. Numerical simulation of turbulent mixing by Rayleigh-Taylor instability. *Physica D*, 12:32-44, 1984.
15. Q. Zhang and M. J. Graham. Scaling laws for unstable interfaces driven by strong shocks in cylindrical geometry. *Phys. Rev. Lett.* 79:2674-2677, 1997.

Table 1. Values of α determined from experiment, theory, and simulation

Method	α	Reference
Experiment	0.05-0.077	[12,14,4,5,6]
Theory	0.05-0.06	[3]
<i>FronTier</i> simulation (unrenormalized)	0.07	This paper (Fig. 2 Left) ref 8
<i>FronTier</i> simulation (renormalized)	0.07	This paper (Fig. 2 Right)
TVD simulation (unrenormalized)	0.035	This paper (Fig. 2 Left)
TVD simulation (renormalized)	0.06	This paper (Fig. 2 Right)

All values are consistent except the unrenormalized TVD value (with α determined from a time independent $t = 0$ Atwood number).

FIGURE CAPTIONS

Fig.1. Front plot for a *FronTier* simulation of RT mixing, with $A = 0.5$. Left frame: early time. Right frame: late time. Color coding represents vertical height. The initial mean height of the interface is 4, and the height scale on the color bar applies only to the later time.

Fig.2. Mixing growth comparison of a *FronTier* (nondiffusive) with a TVD (diffusive) simulation. For the TVD simulation, two grid levels are shown, the coarser being $64^2 \times 128$. In all cases, h is the height of the 1% volume fraction contour, and the initial mean height of the interface is 4. Left: h vs. $A(t=0)gt^2$ for *FronTier* and TVD. Right: h vs. $2 \int_0^t \int_0^{s_1} A(s)gdsds_1$ for *FronTier* and TVD. The solid line represents the *FronTier* simulation, the dashed line is the finer grid TVD simulation and the dotted line is the coarser grid TVD simulation.

Fig.3. Cross sectional plots showing density on a common rainbow color scale. The pure light fluid is colored blue and the pure heavy fluid is red. Yellow and green represent various levels of microscopic mixing. The ratio of extreme density values is 3.3:1. The right frames show a higher slice in the z direction. Top: *FronTier*, bottom: TVD. The simulations are shown at comparable penetration distances, but at different times ($Ag t^2 = 23$ for *FronTier*, $Ag t^2 = 66$ for TVD). It is evident that the density contrast for the TVD simulation has been reduced by about 50% due to mass diffusion. See also Fig.4.

Fig.4. Time dependent A (Atwood number) for fine grid *FronTier*, fine grid TVD, and coarse grid $64^2 \times 128$ TVD. At time $t = 0$, all three simulations have $A(t = 0) = 0.5$. This plot displays the reduced buoyancy of the diffusive TVD simulations as a function of time.

Fig.5. Left: Unmixed bubble of light fluid. Center: Unmixed bubble and heavy fluid mass which will be mixed with it. Right: Mixed bubble.

Fig. 6. Convergence under mesh refinement is demonstrated for the rate of growth of the mixing layer for an axisymmetrically perturbed spherical Richtmyer-Meshkov instability simulation, in the case of heavy fluid exploding light.

Fig. 7. Asyptotic scaling in the strong shock limit is demonstrated for the same problem as presented in Fig.6. Scaled results for incident shock Mach numbers 10 to 200 are virtually identical.

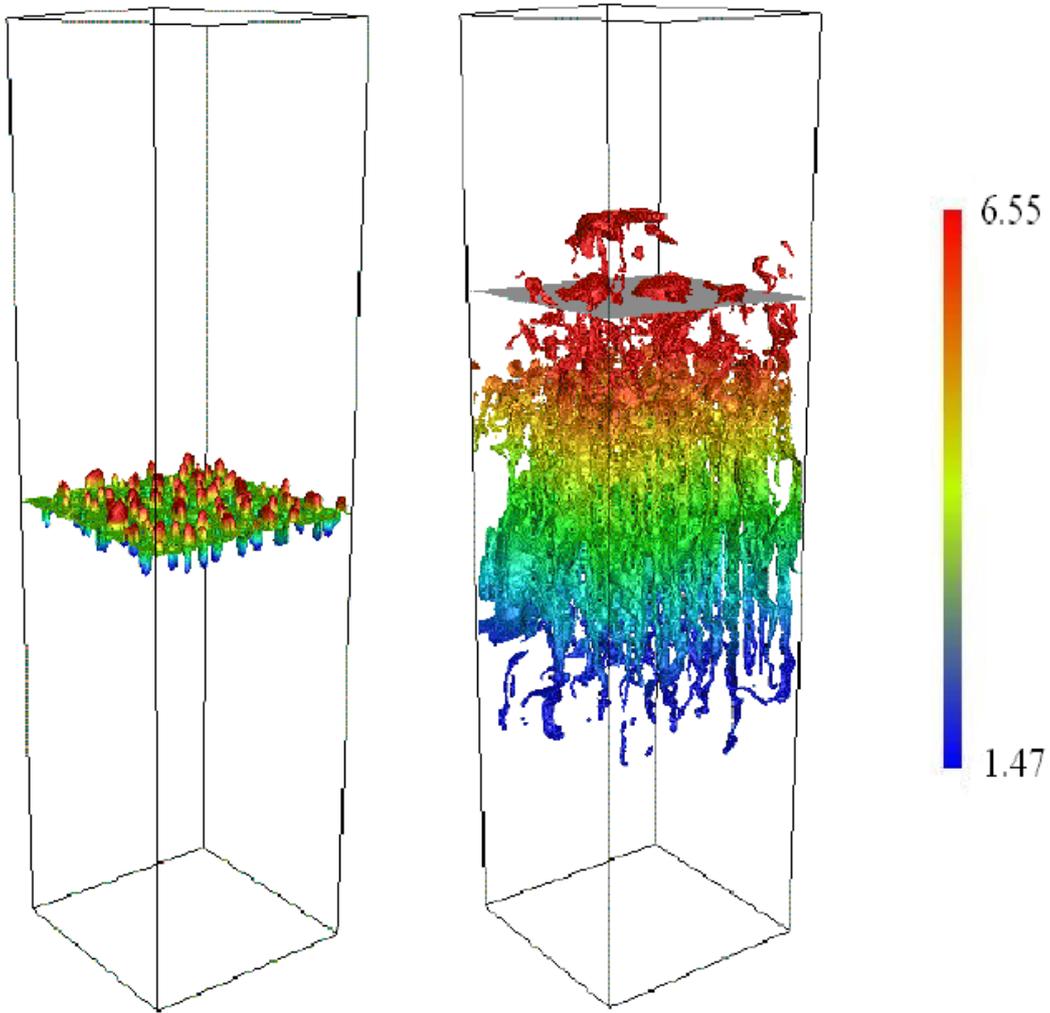


Figure 1.

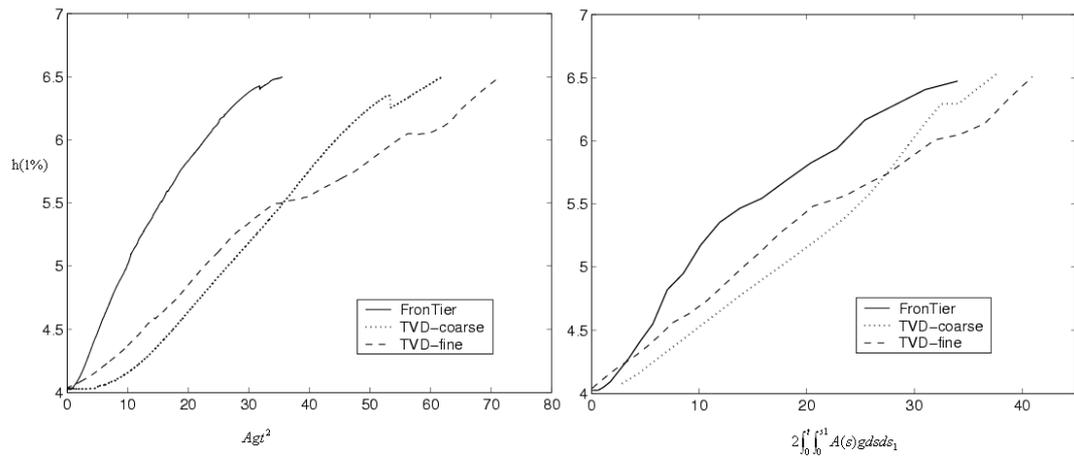


Figure 2

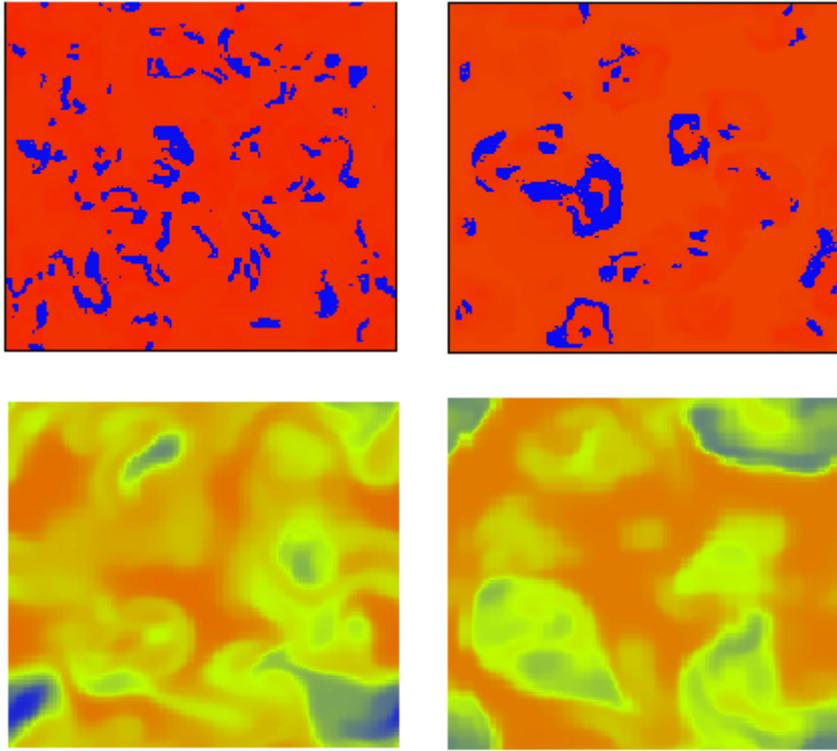


Figure 3.

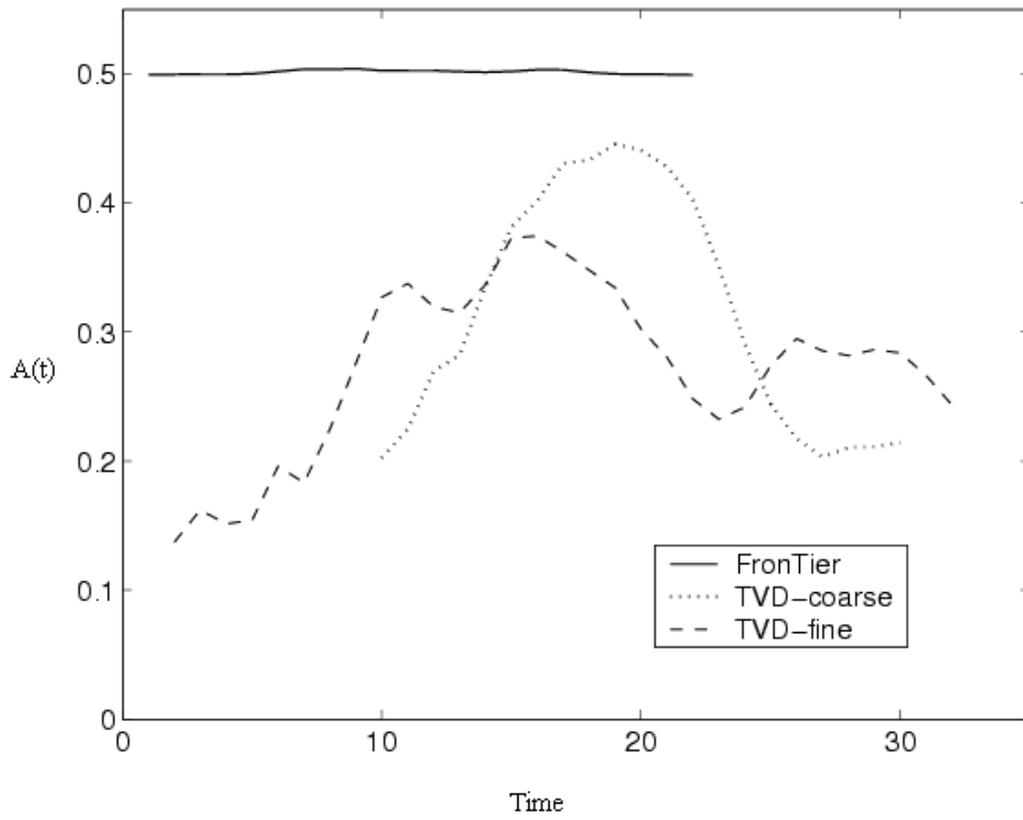


Figure 4.

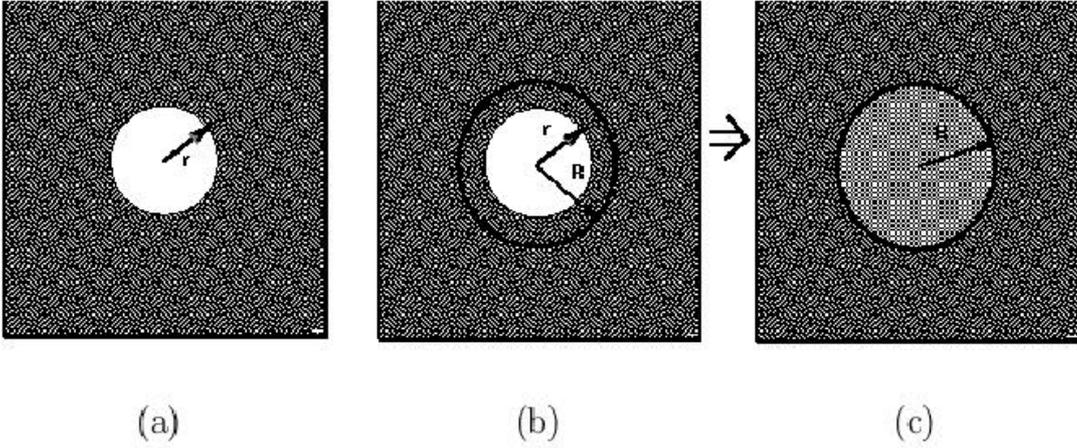


Figure 5.

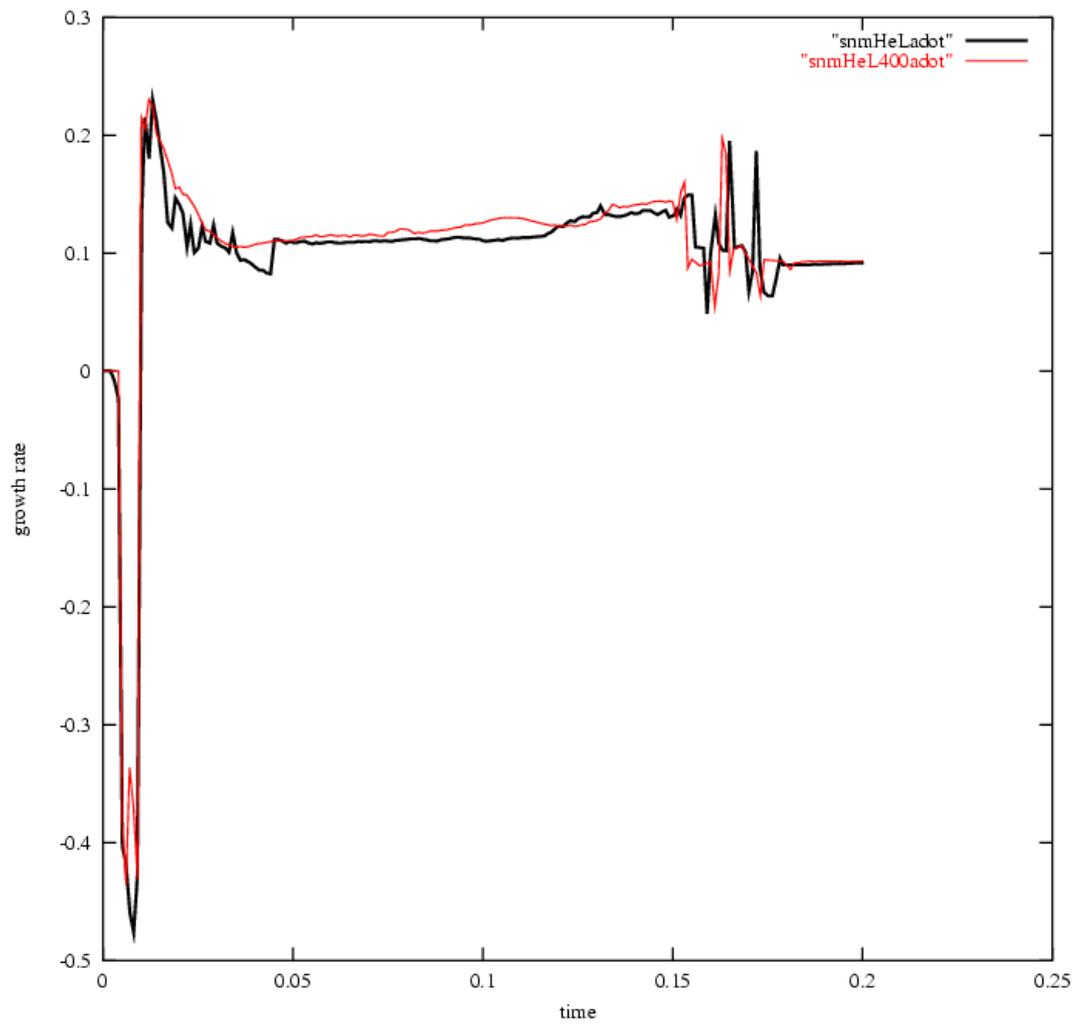


Figure 6.

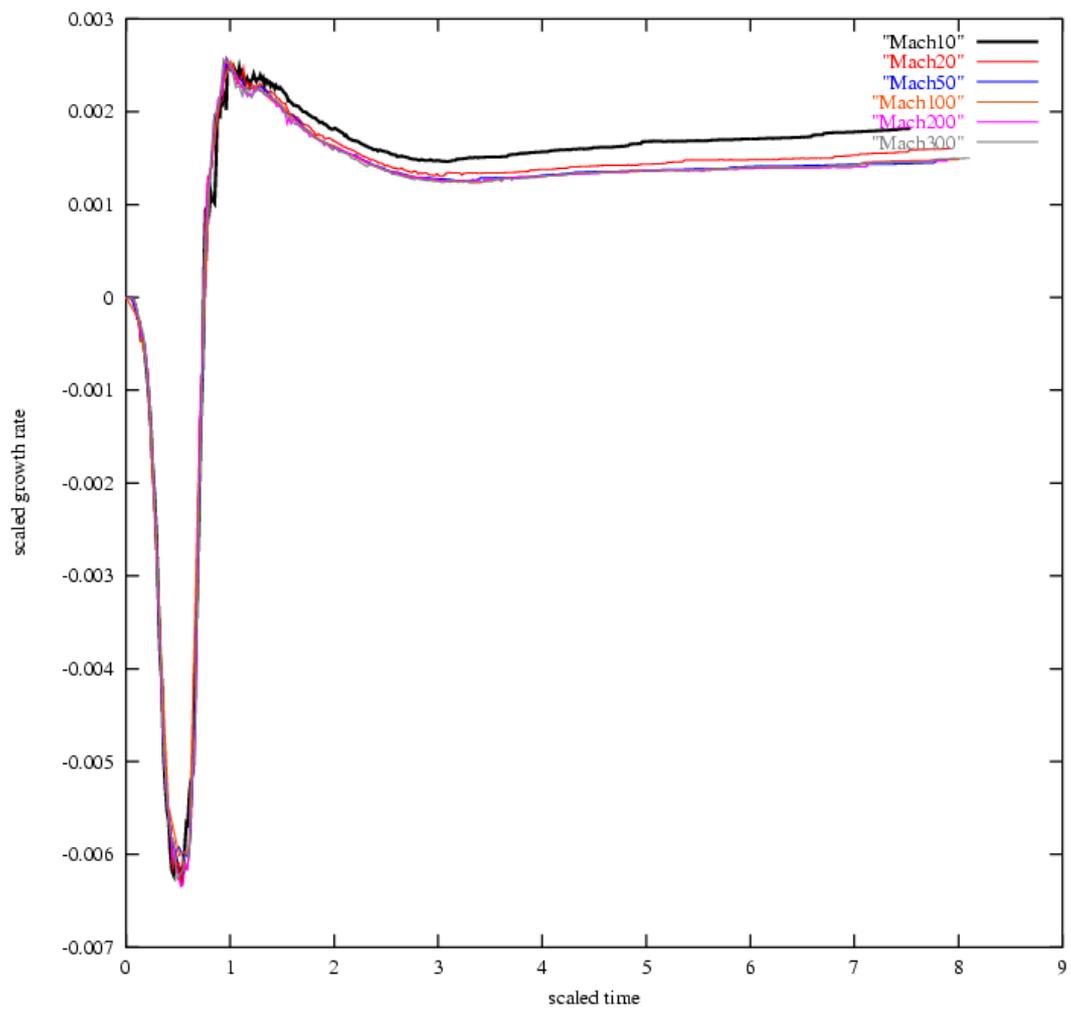


Figure 7.

