



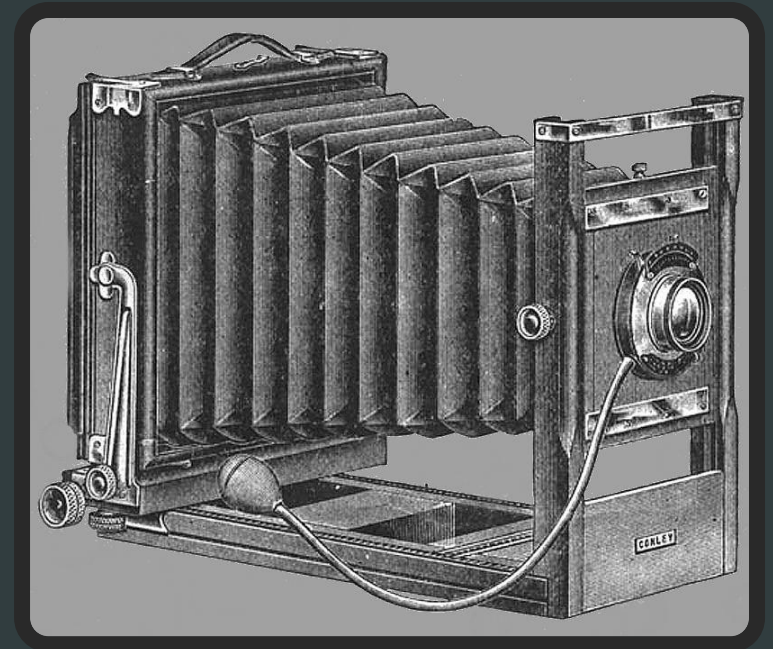
Large Synoptic Survey Telescope: Building the biggest digital camera at BNL



Paul O'Connor
BNL Instrumentation Division

Anže Slosar
Physics Department

- Taking pictures of the night sky
- Discovering Dark Energy
- Building the largest digital camera

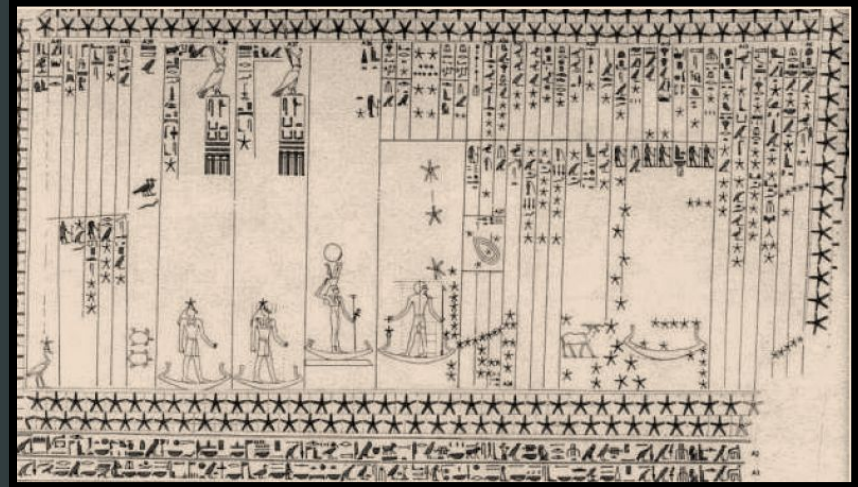


Capturing the night sky

N. Europe
1600BC



Egypt
1460BC



Mesoamerica
1050AD



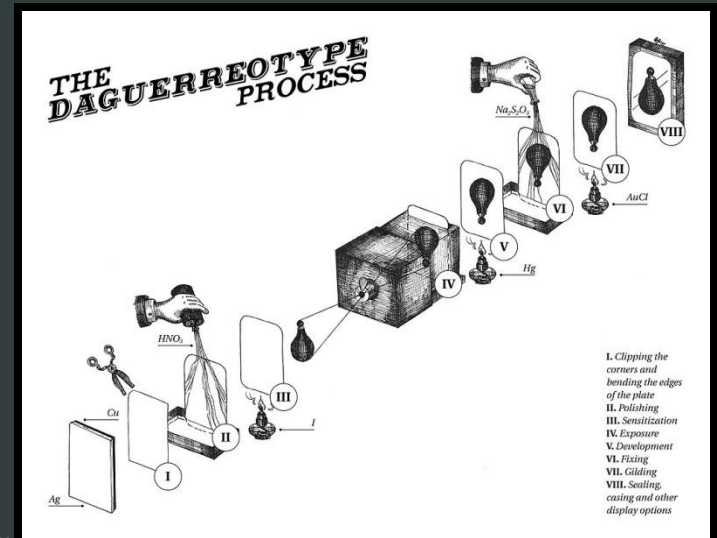
China
185BC



Photochemical imaging: the Daguerreotype



Louis Daguerre
1839



First astrophotograph



Moon
Draper 1840

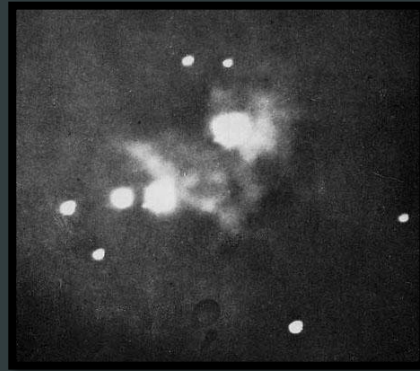


Bond
1852

Early astrophotography



Vega
Bond 1850



Orion nebula
Draper 1880

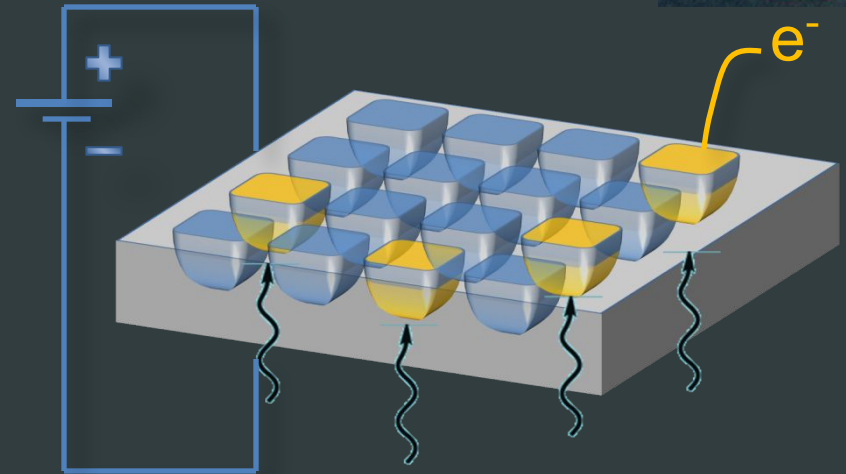
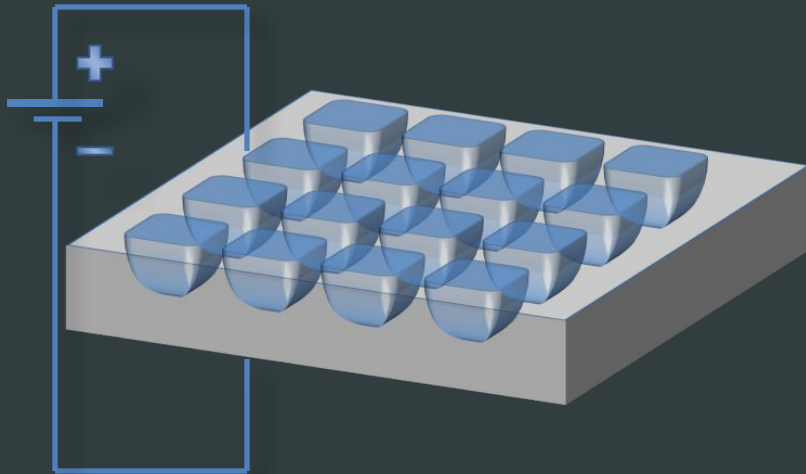


M31
Roberts 1887

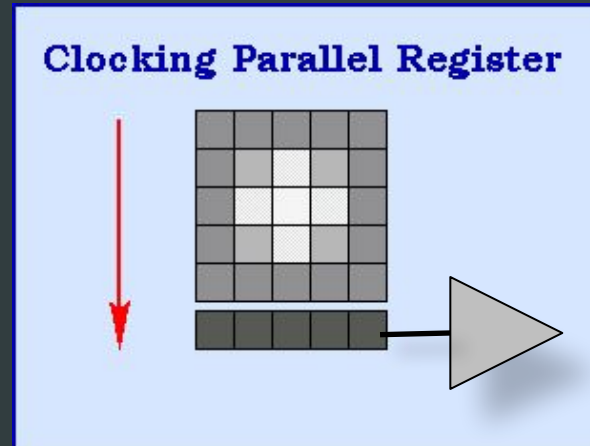
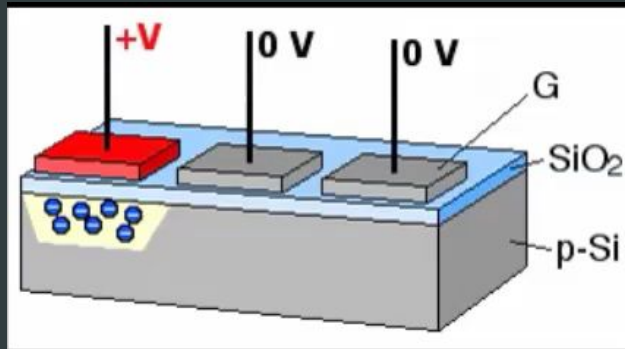


Transit of Venus
Janssen 1874

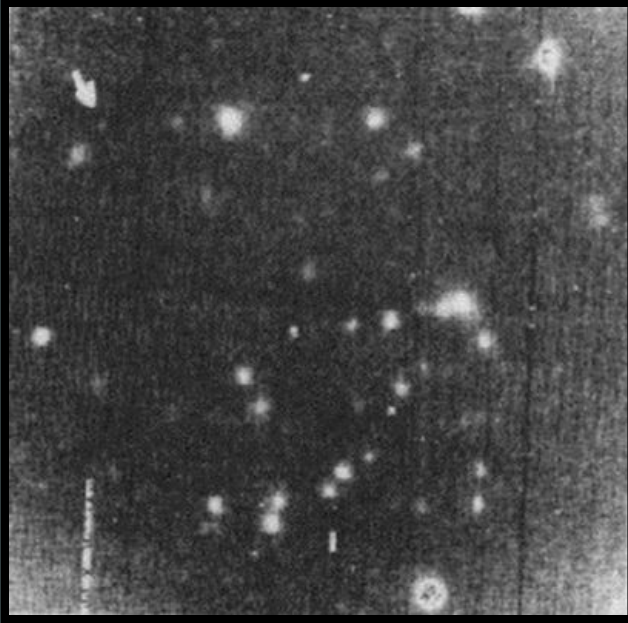
Electronic imaging: the Charge Coupled Device



William Boyle and George Smith received 2006 Nobel prize for invention of CCDs



CCD astrophotography



Cl 13K α
Gunn 1981
500x500

“The CCD is a nearly perfect device, with problems that appear only because the device performance is so high compared to other detectors...”

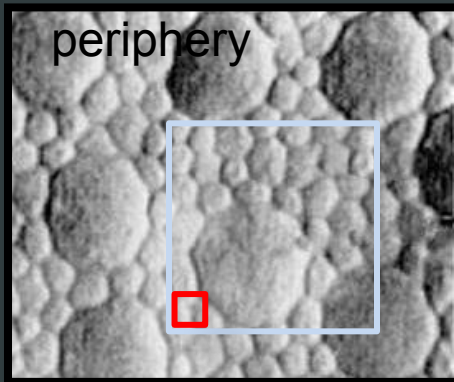


NASA/ESA/Hubble Heritage Team
(STScI/AURA)/J. Hester, P. Scowen
(Arizona State U.)

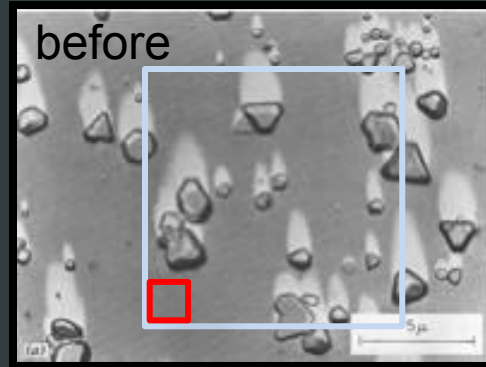
M16
2014

Three types of photosensor compared

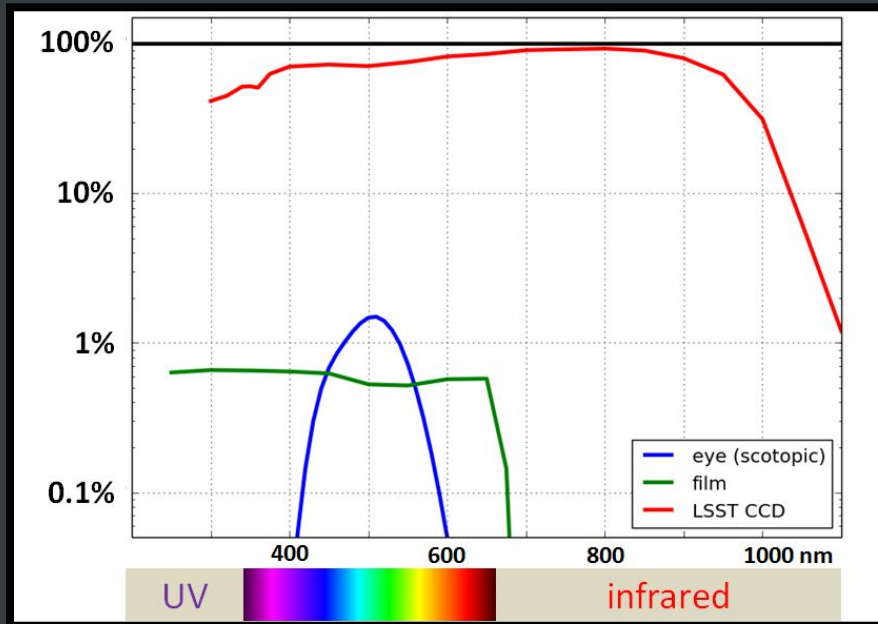
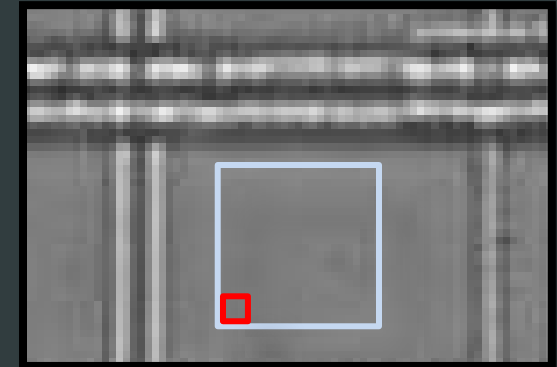
retina



film

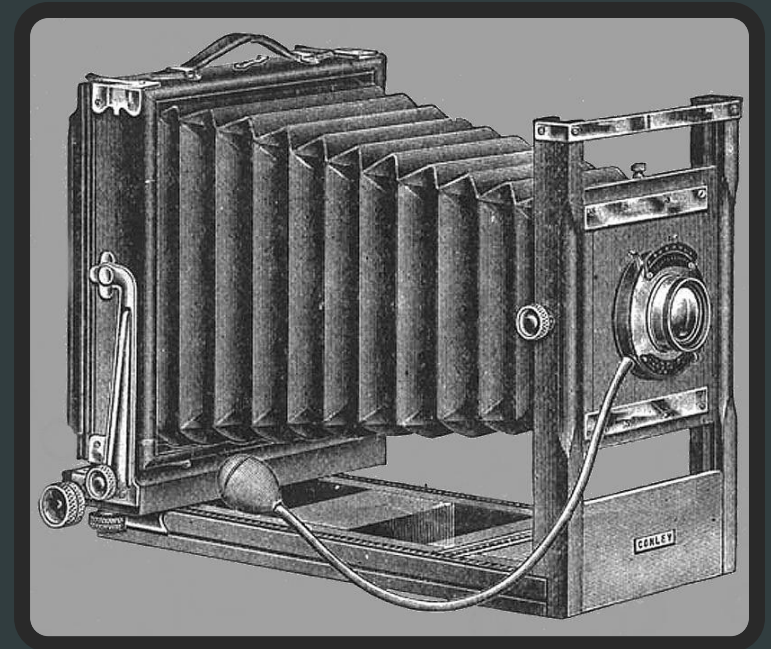


CCD

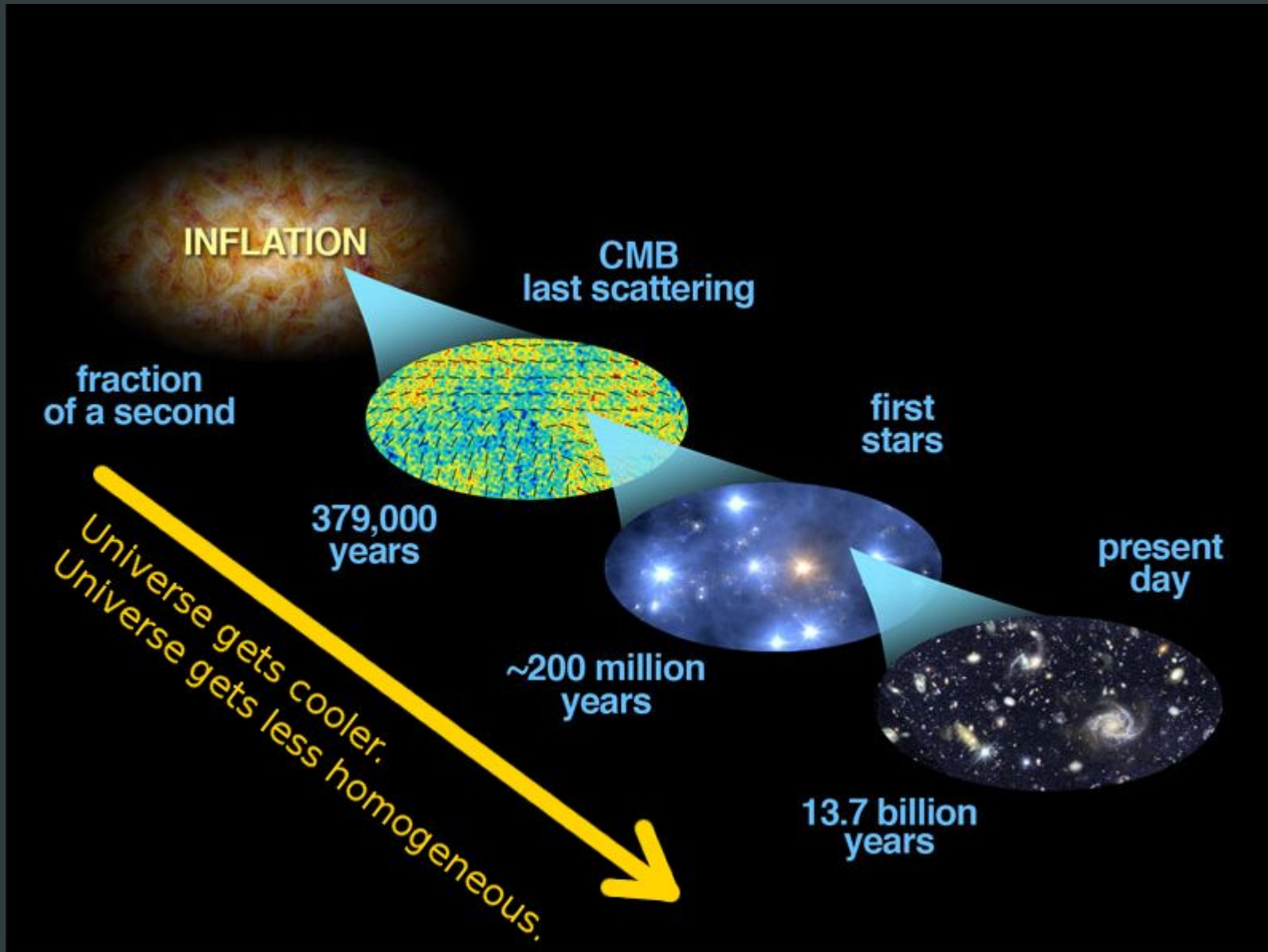


	Dynamic range	Retention time
Retina	5000	30ms
Film	16000	100ms
CCD	60000	∞

- Taking pictures of the night sky
- **Discovering Dark Energy**
- Building the largest digital camera



Cosmic History



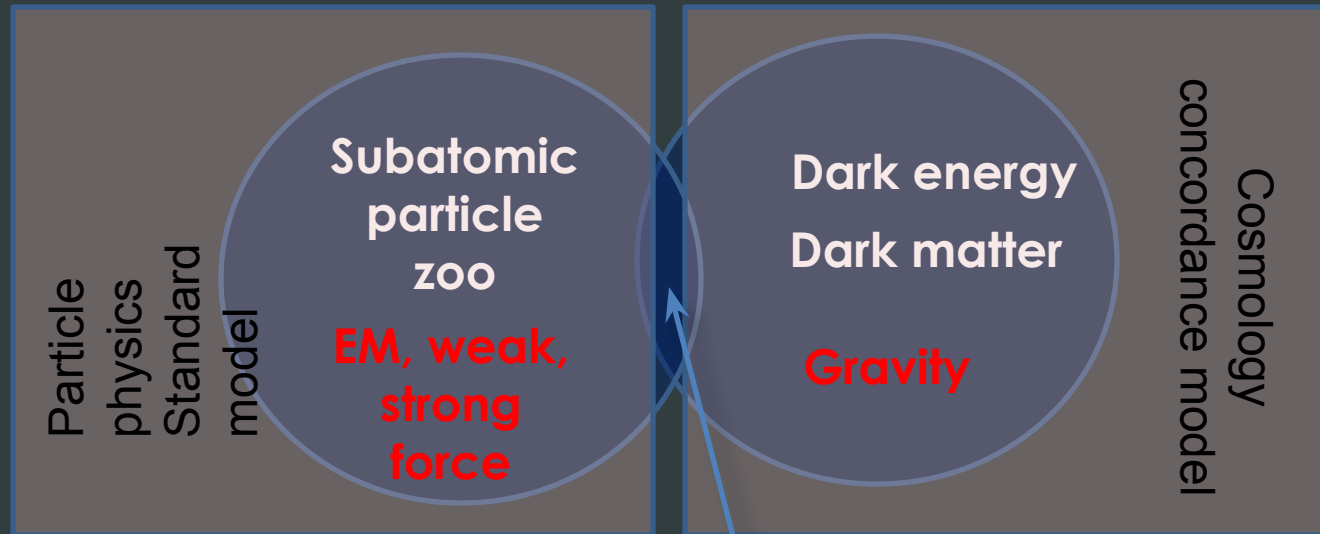
Cosmic Pizza

- 95% of the Universe made of stuff we don't know
- dark components well understood macroscopically, but their physics remains a mystery
- **Dark Matter** is cold, non-interacting “stuff” collapsing under its own gravity



- **Dark Energy** drives an accelerate expansion of the Universe - one of the most unexpected turns of late 20th century

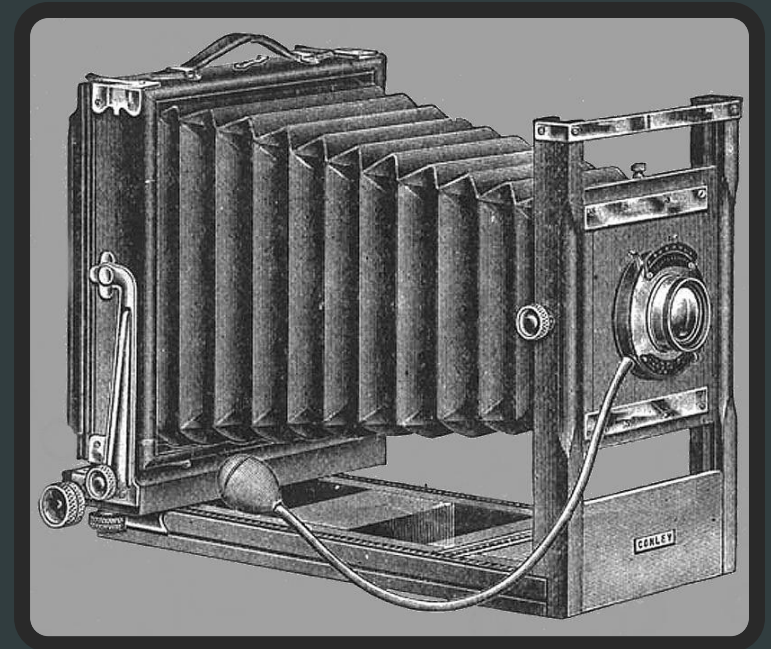
Clash of the “standard models”



Atoms
Plasma
Radiation

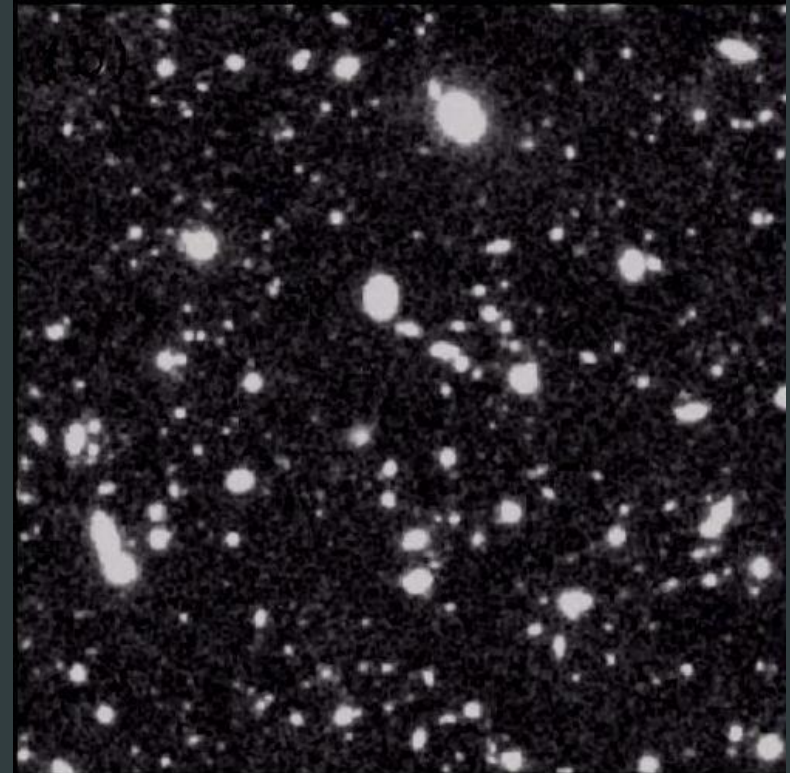
“The acceleration of the universe is, along with dark matter, the observed phenomenon that most directly demonstrates that our theories of fundamental particles and gravity are either incomplete or incorrect”.
Albrecht et al., Dark Energy Task Force report

- Taking pictures of the night sky
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The ideal survey for cosmology

- Detect as many galaxies as possible
- Over as wide an area of sky as possible
- Measure their positions, brightness, colors, and shapes as accurately as possible



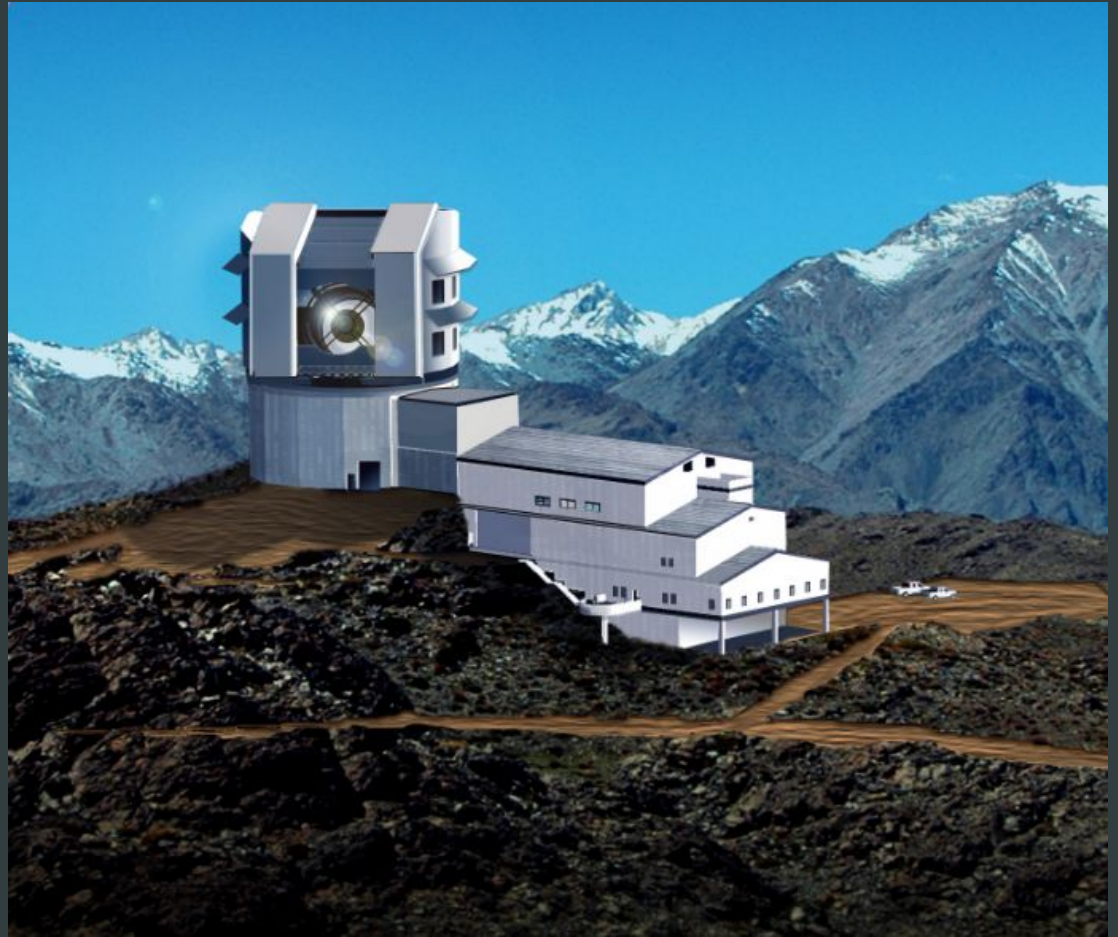
Miyazaki et al. 2011

Microscope vs Telescope



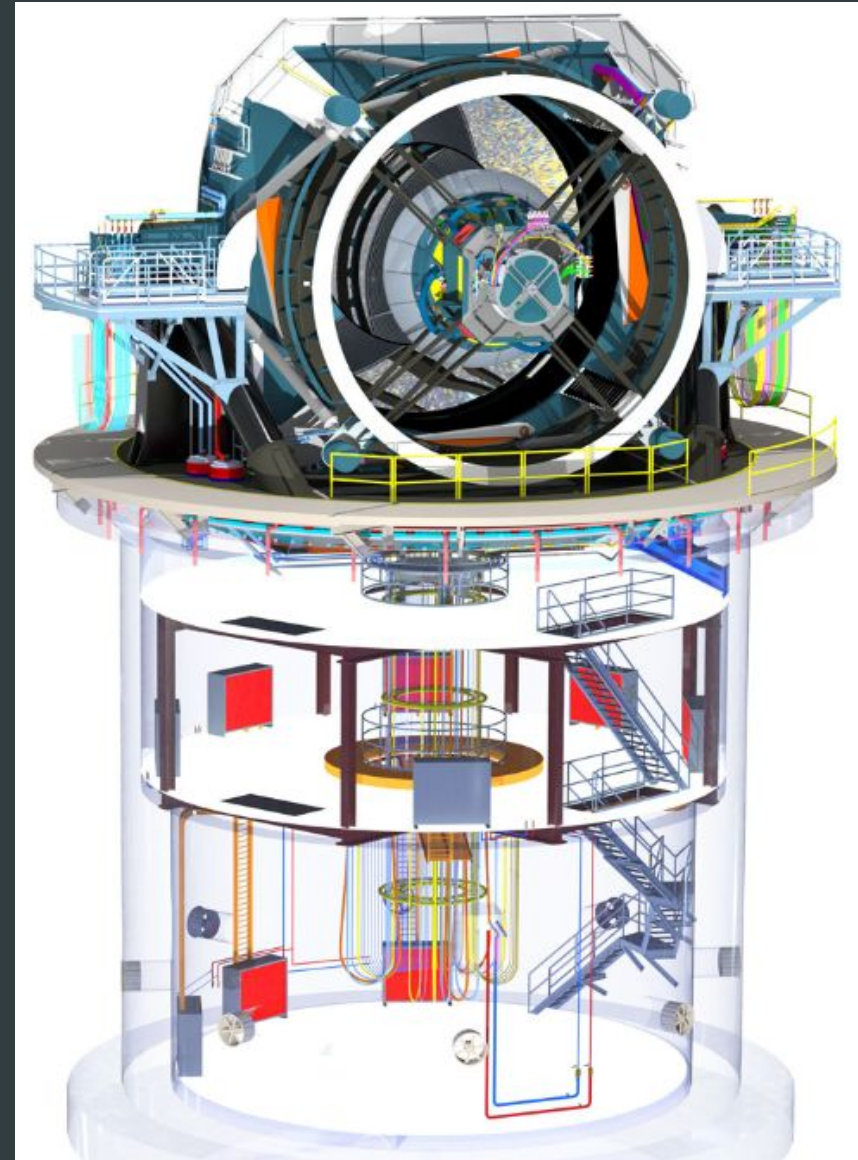
Large Synoptic Survey Telescope

- Leadership survey telescope
- Collaboration between NSF and DOE
- Will do all kinds of science: from asteroid searches to the largest observable structures in the universe
- We are in the game for cosmology



LSST in a Nutshell

- Decade long survey 2022-2032
- 8m mirror, 3200 megapixel camera
- takes a picture every 15s - movie of the sky
- it will detect 37 billion sources
- Cost and schedule:
 - Telescope, site, DM: \$473M
 - Camera: \$165M
 - Operations: \$37M/yr
 - First light: 2019
 - Survey operations start: 2022
- BNL responsible for detectors and analysis

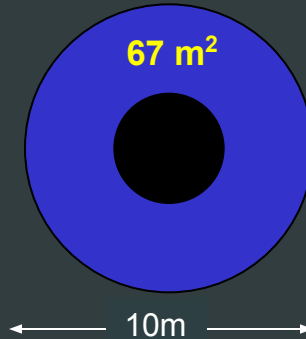


Collecting area \times Field of View = Survey Power

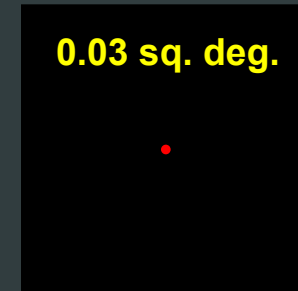
Keck



Primary mirror effective area



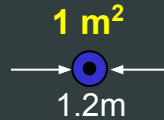
Field of view



Product

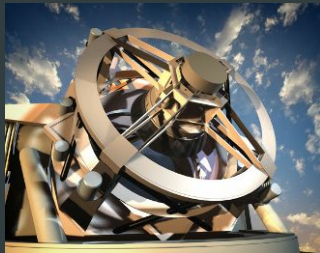
2 m²-deg²

Oschin Schmidt



20 m²-deg²

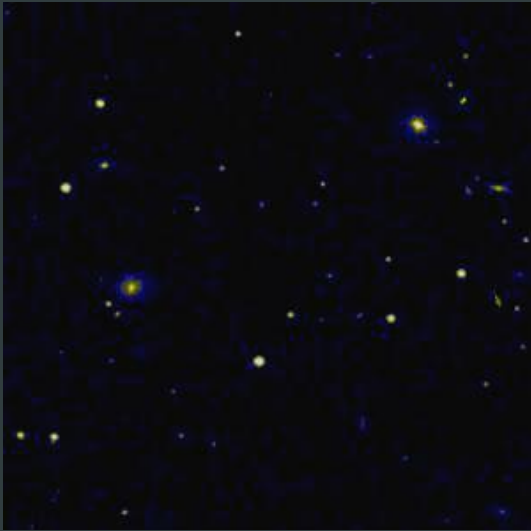
LSST



320 m²-deg²

LSST combines large aperture and field of view to go WIDE and DEEP.

LSST's survey will be DEEP



Largest all-sky
survey before LSST
(photographic)



Sloan digital sky
survey
(CCD)



LSST
(simulated)

LSST's survey will be FAST

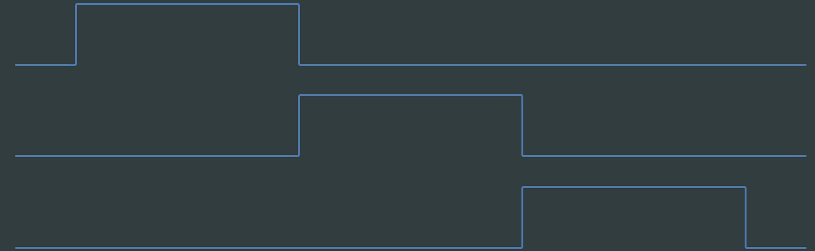


Field1

Field2

Field3

...

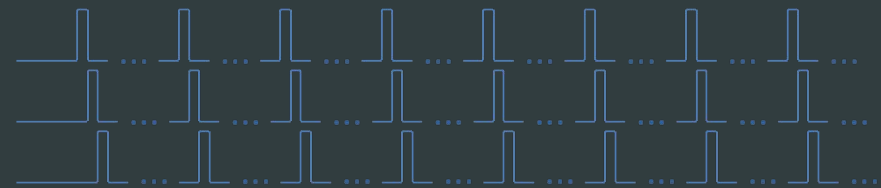


Field1

Field2

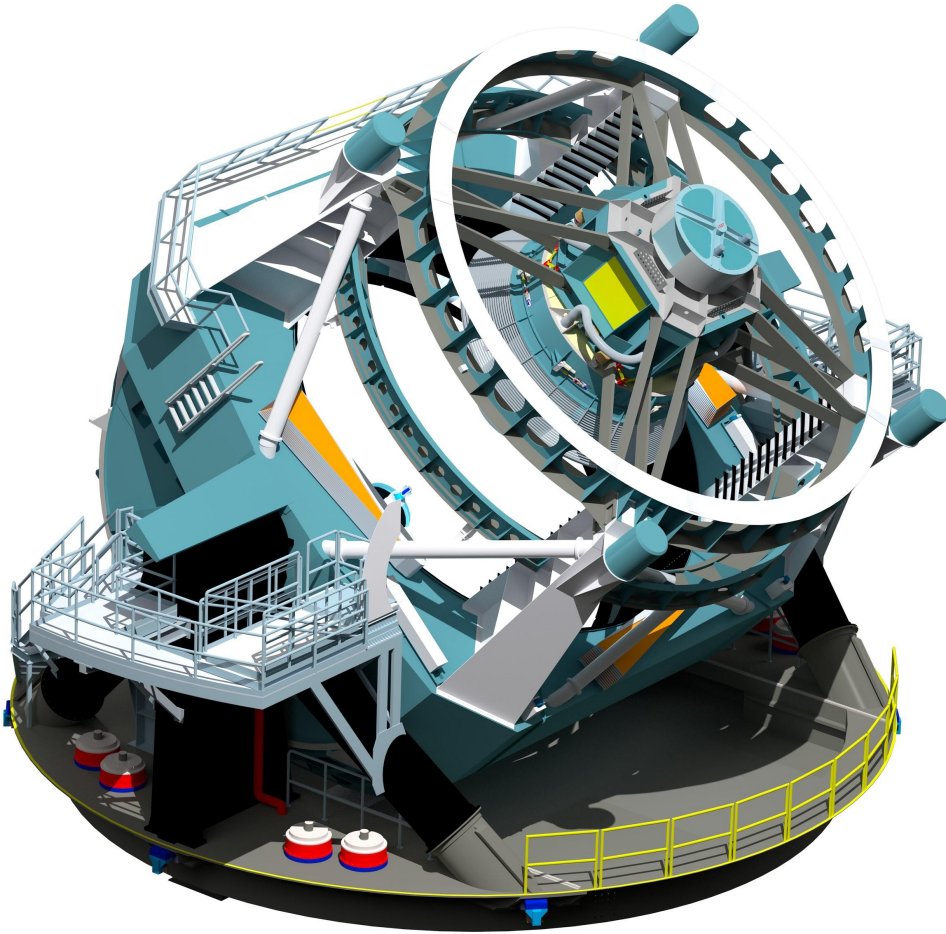
Field3

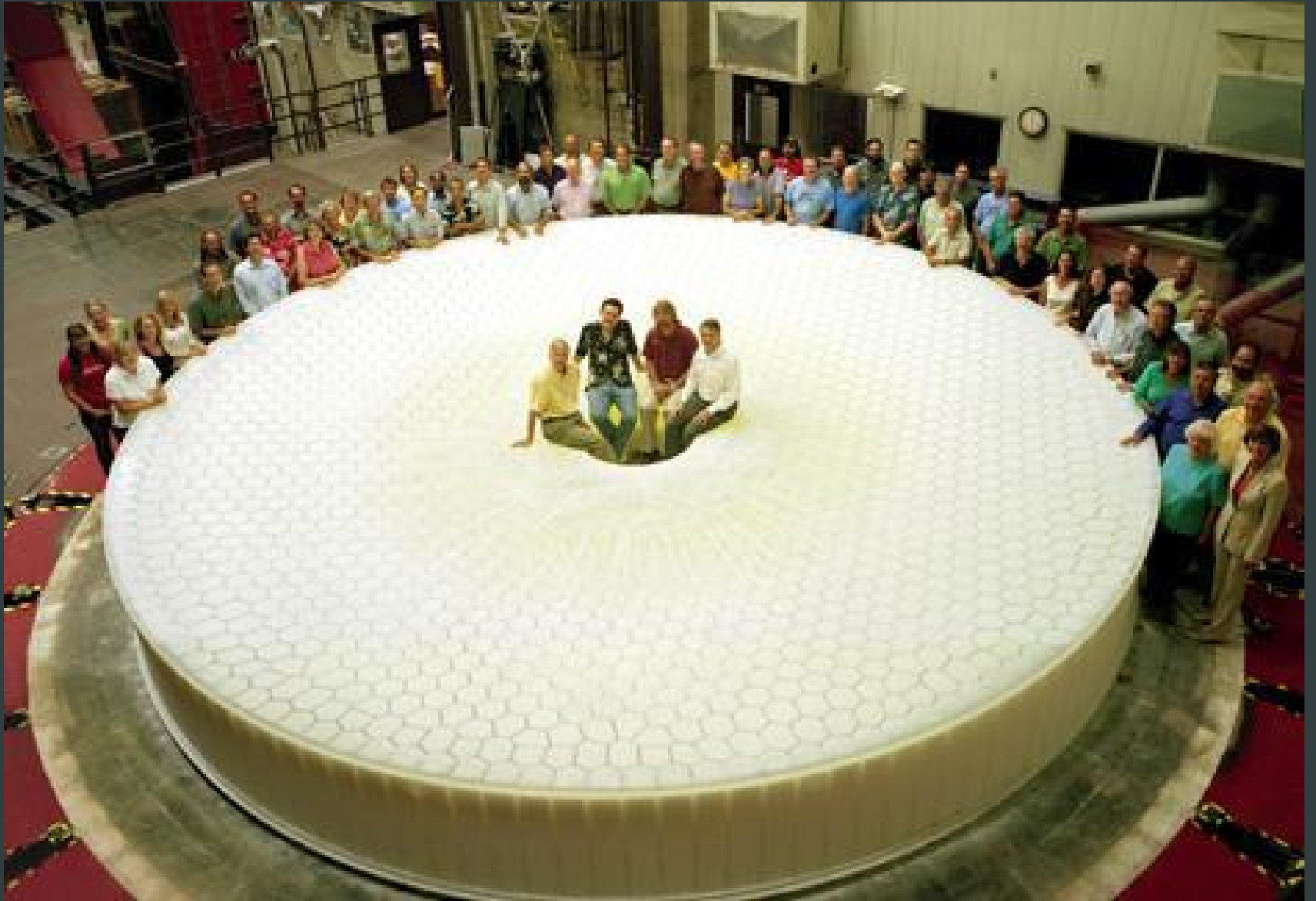
...



Telescope

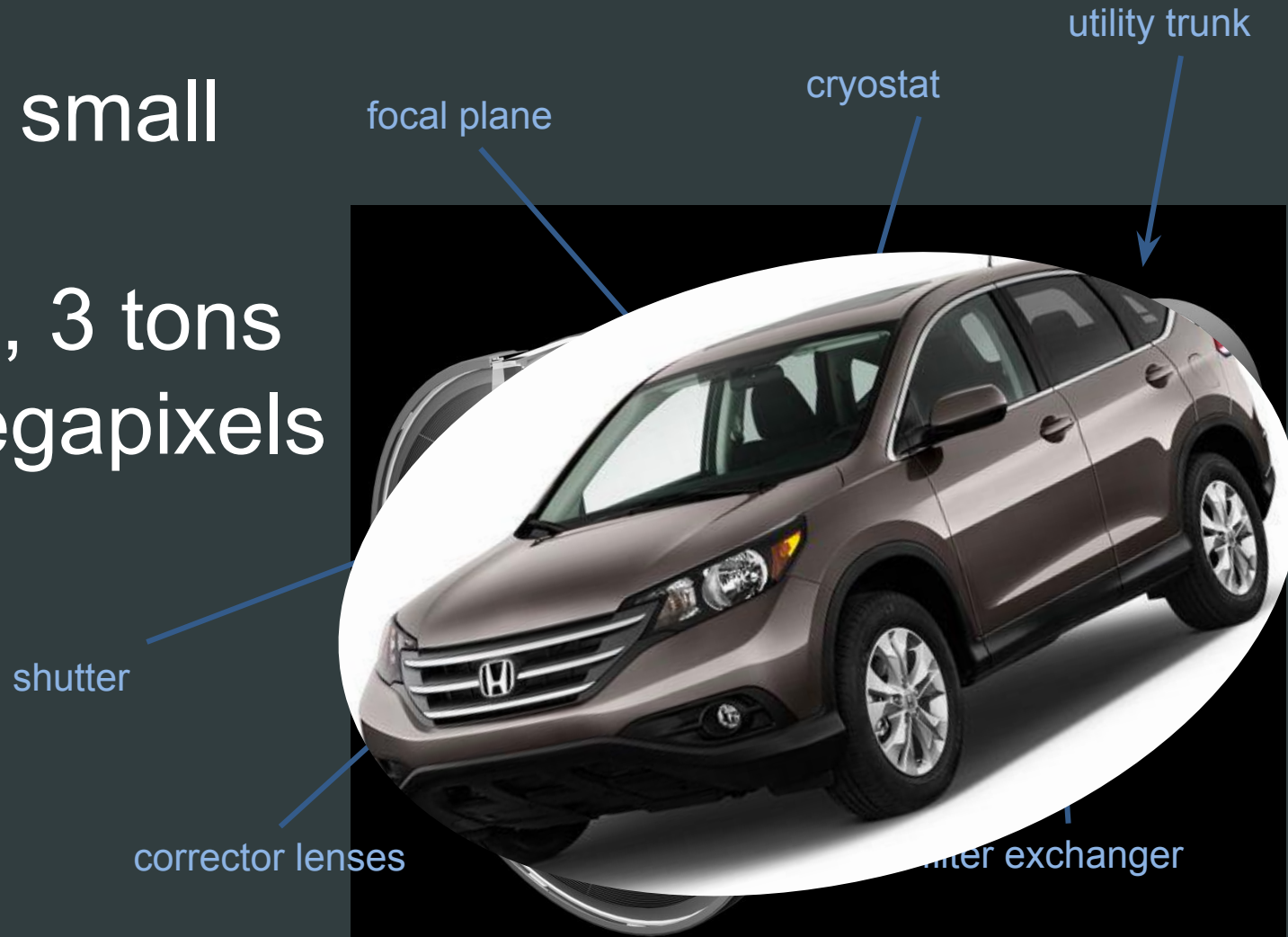
- 3-mirror design
- Fast $f/1.2$ beam
- Moving structure 350t
- 5s slew-and-settle time to $0.25''$





LSST camera – a different kind of pixel detector

- size of a small car
- 6200 lbs, 3 tons
- 3200 megapixels



LSST's new generation CCD sensors

Thick, fully depleted, high resistivity silicon substrate

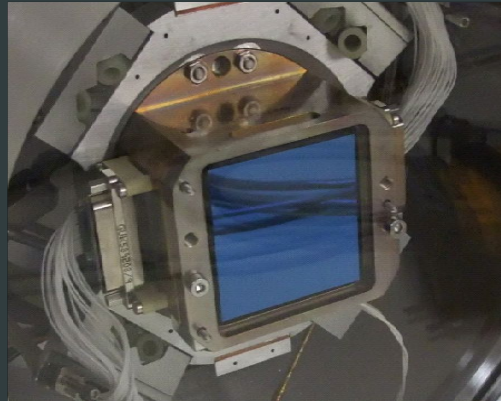
4K x 4K, 10 μ m pixels

16 independent amplifier segments

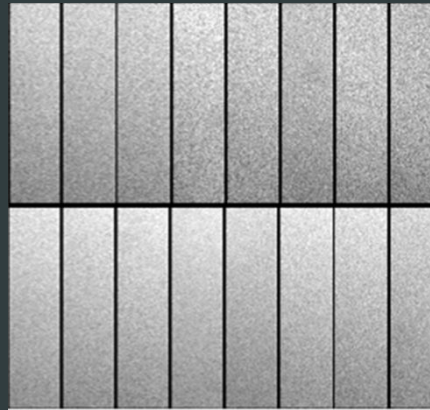
Plug-compatible prototypes obtained from 2 vendors, fully operational

E2V CCD250

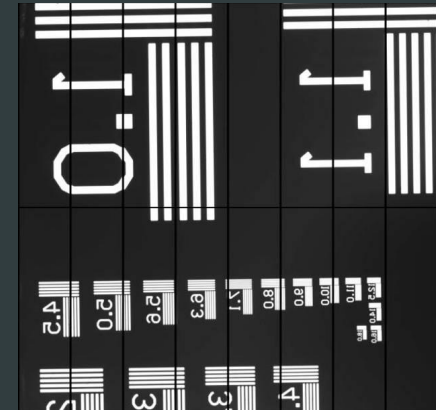
Packaged CCD



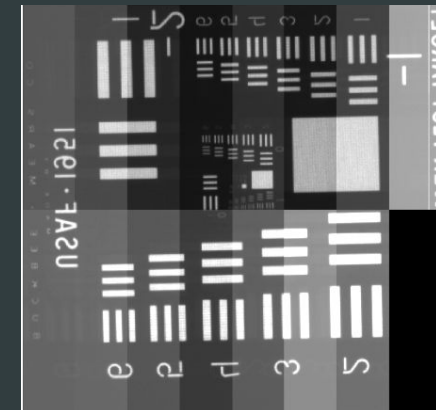
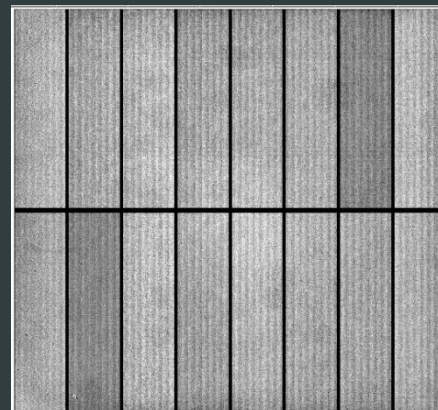
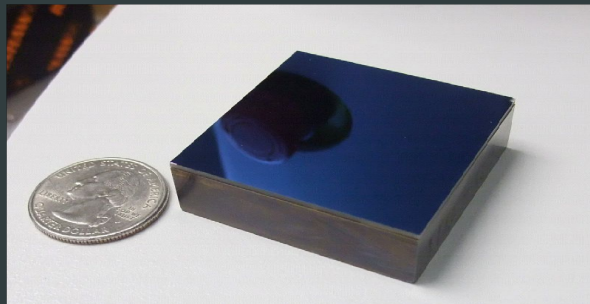
Uniform illumination



Test target image



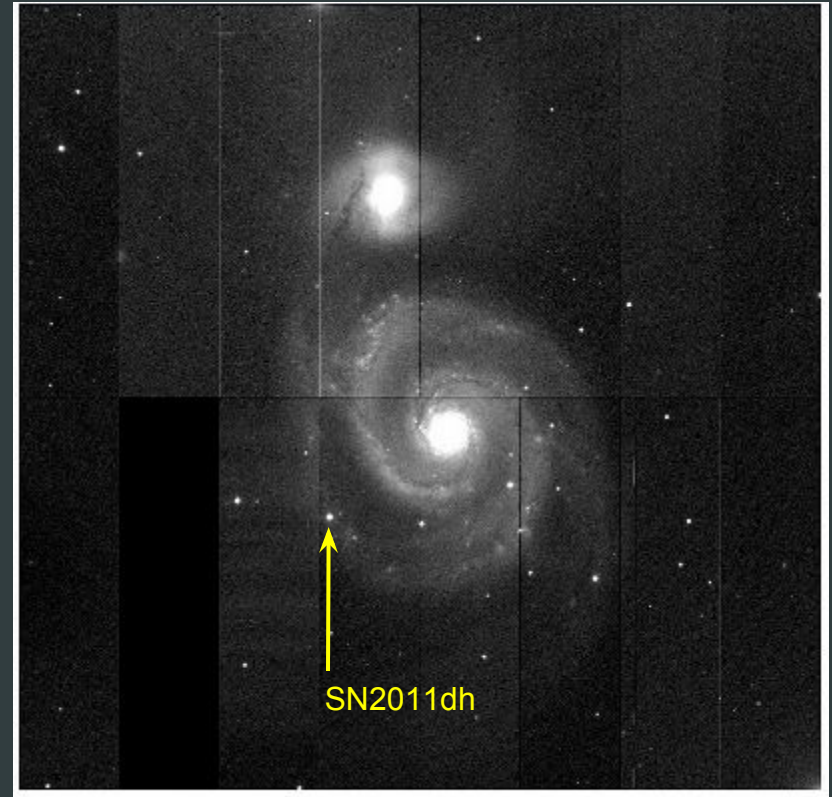
ITL STA3800B



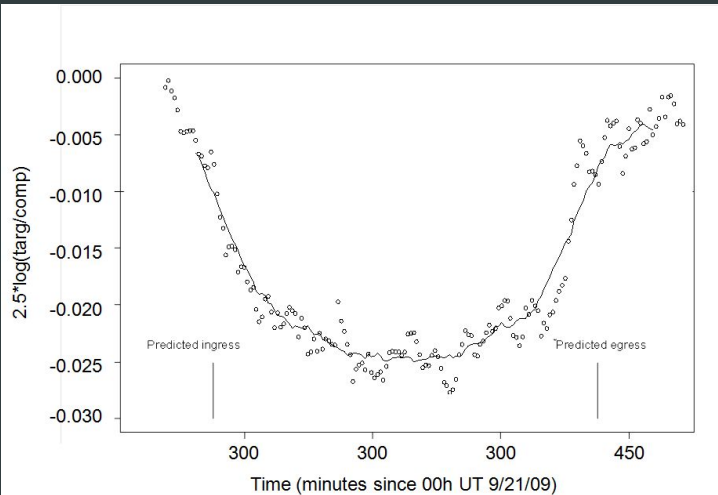
LSST CCD telescope test



Interfacing to 2.4m McGraw-Hill



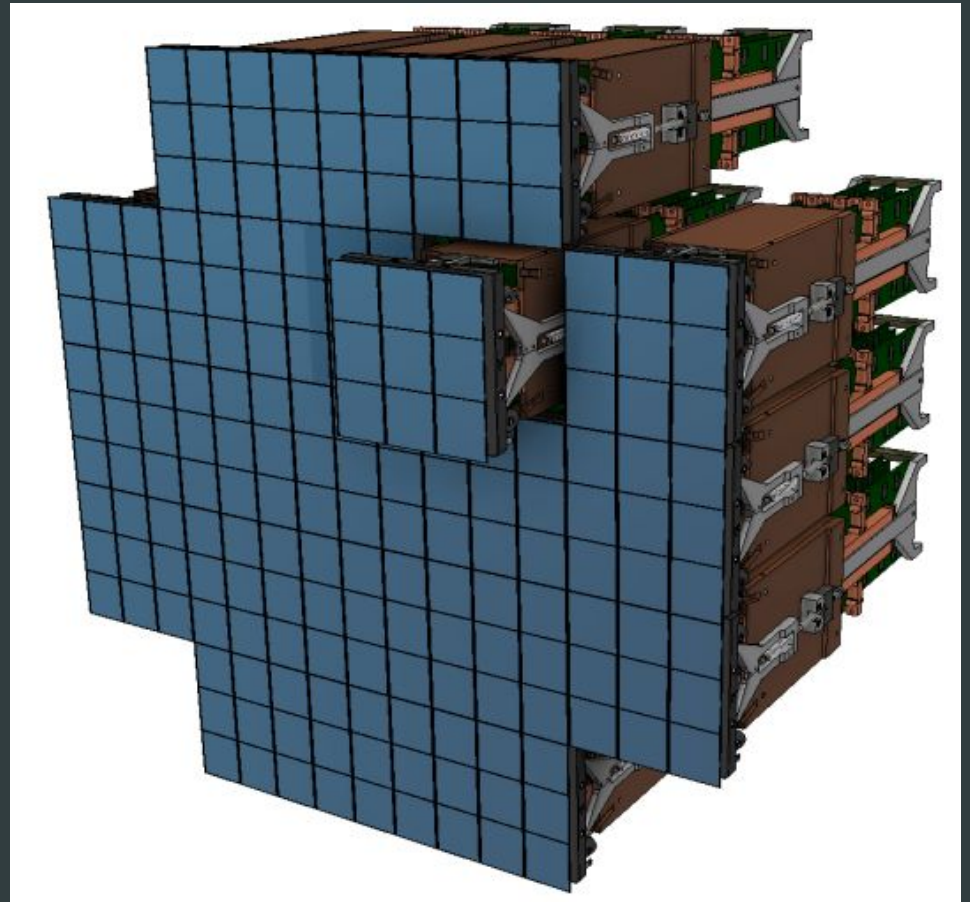
M51



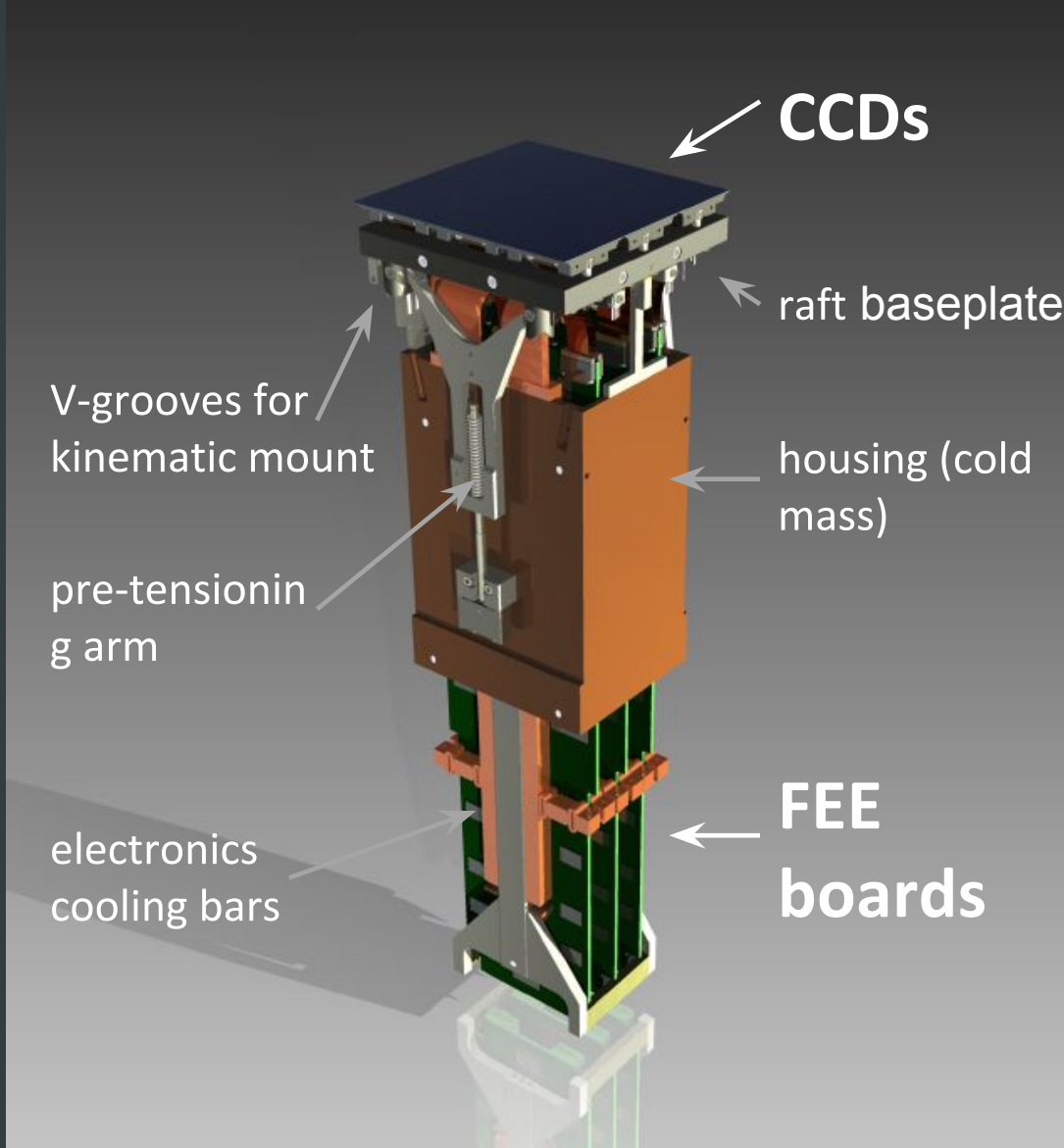
Exoplanet transit of TRES-1B

LSST science focal plane

- 21 “rafts”
- 9 CCDs each
- Integrated readout electronics
- Inside vacuum cryostat
 - CCD temperature -100°C
 - Electronics -20°C



Raft Tower Module



Complete 144-Mpixel imager

Provide bias, timing, and control signals for CCD operation

Low noise analog signal processing

Digitizing, multiplexing of pixel data

Diagnostics

Support sensors mechanically to meet strict coplanarity

Thermal management of sensors and electronics

12.7 x 12.7 x 42cm³; 12kg

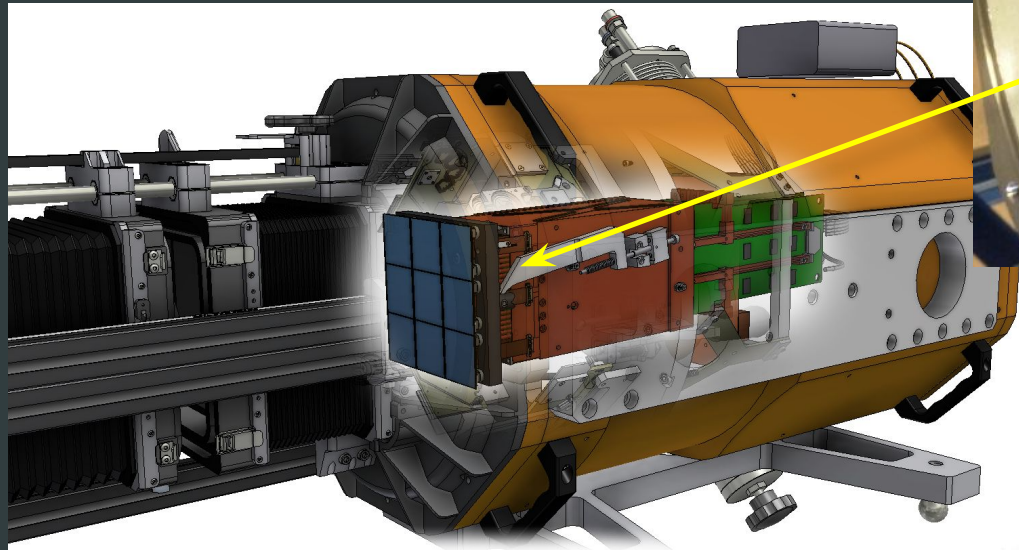
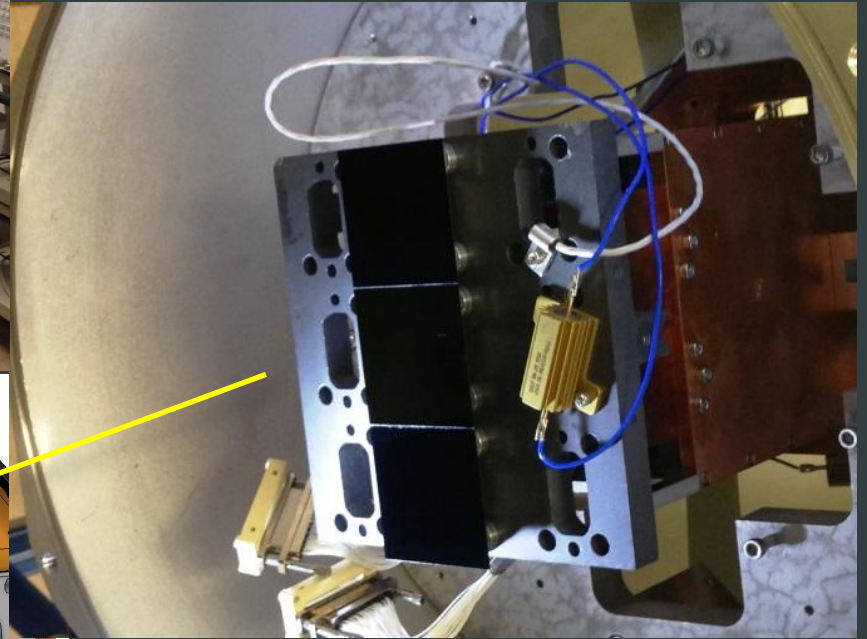
Photons in – bits out

BNL's deliverable

Raft tower module under construction



Testing in lab cryostat



Performance: photons in to images out

SAOImage ds9

File Edit View Frame Bin Zoom Scale Color Region WCS Analysis Help

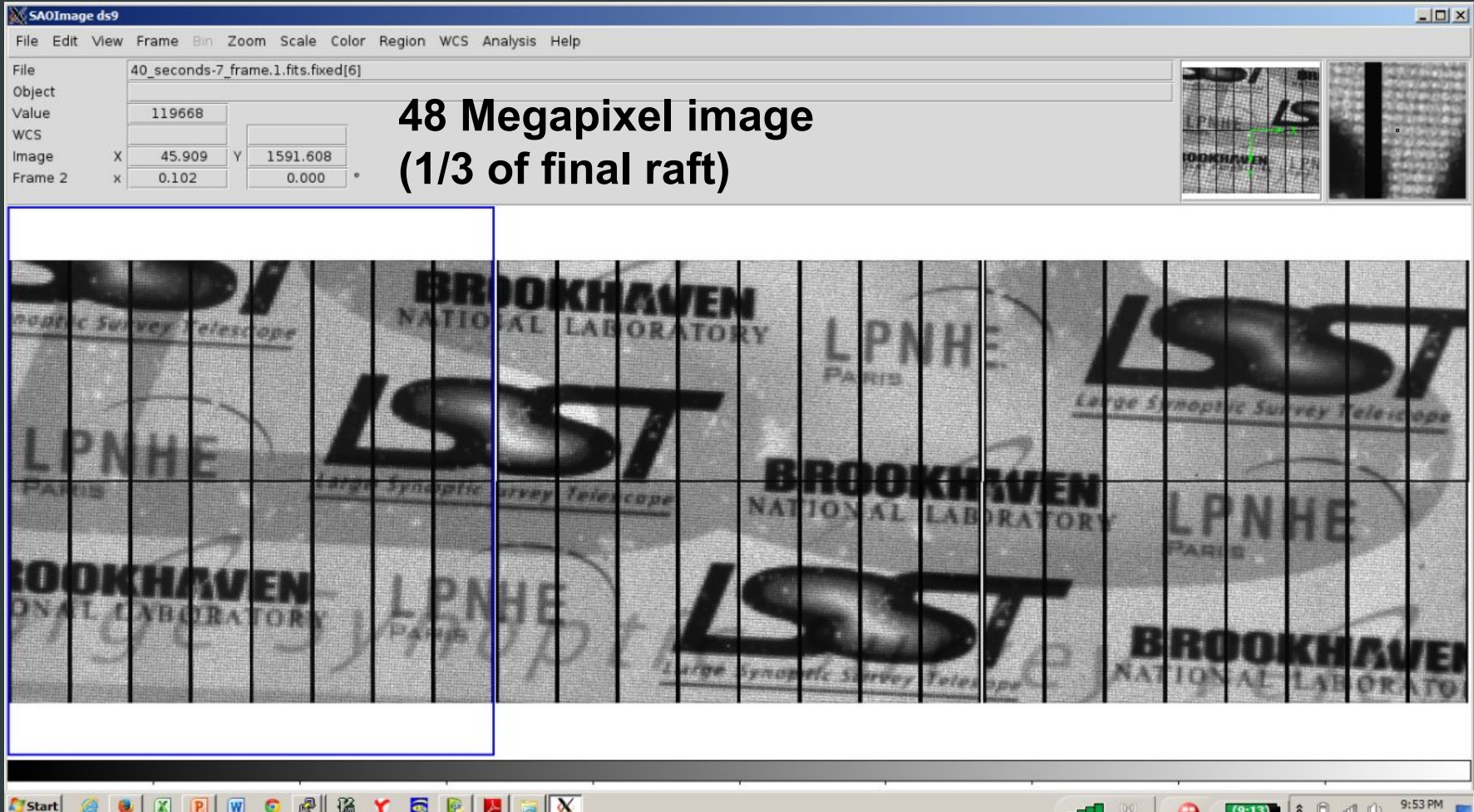
File: 40_seconds-7_frame.1.fits.fixed[6]

Object Value: 119668

WCS Image X: 45.909 Y: 1591.608

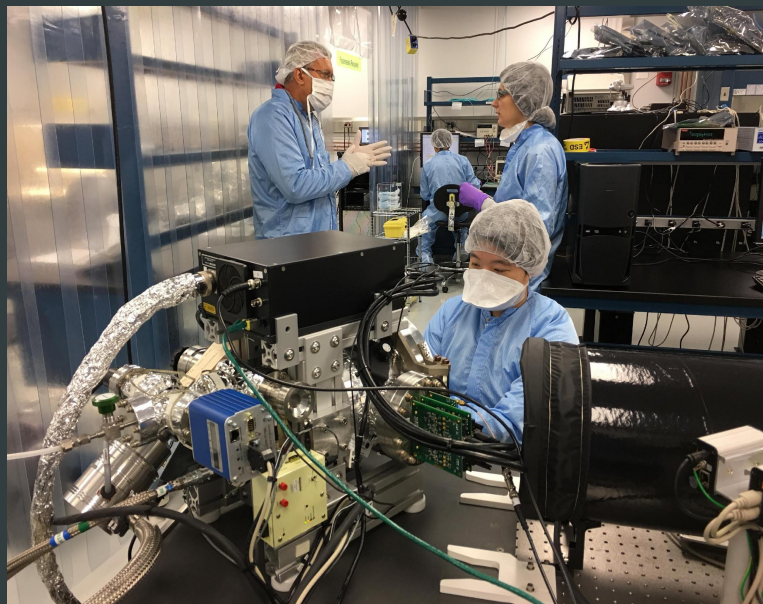
Frame 2 x: 0.102 y: 0.000

48 Megapixel image (1/3 of final raft)



Start [Taskbar icons] 9:53 PM

Production of rafts commenced this summer...



CCD Specifications for each sensor

Status	Spec. ID	Description	Specification	Measurement
×	CCD-007	Read Noise	$< 8 e^-_{\text{rms}}$	$7.79\text{--}8.78 e^-_{\text{rms}}$
✓	CCD-008	Blooming Full Well	$< 175000 e^-$	$161498\text{--}170522 e^-$
×	CCD-009	Nonlinearity	$< 2\%$	max. fractional deviation from linearity: 3.5×10^{-2}
✓	CCD-010	Serial CTE	$> 1 - 5 \times 10^{-6}$	$1 - 2.53 \times 10^{-6} \pm 1.20 \times 10^{-5}$ (min. value)
×	CCD-011	Parallel CTE	$> 1 - 3 \times 10^{-6}$	$1 - 5.44 \times 10^{-4} \pm 3.53 \times 10^{-6}$ (min. value)
×	CCD-012	Active Imaging Area and Cosmetic Quality	$< 0.5\%$ defective pixels	defective pixels: 306336 (1.8993%)
...	CCD-012a	Bright Pixels	...	1640
...	CCD-012b	Dark Pixels	...	3753
...	CCD-012c	Bright Columns	...	151
...	CCD-012d	Dark Columns	...	0
...	CCD-012e	Traps	...	0
...	CCD-013	Crosstalk	$< 0.19\%$...
✓	CCD-014	Dark Current 95th Percentile	$< 0.2 e^- s^{-1}$	$1.80 \times 10^{-1} e^- s^{-1}$
✓	CCD-021	u Band QE	$> 41\%$	62.4%
✓	CCD-022	g Band QE	$> 78\%$	88.6%
✓	CCD-023	r Band QE	$> 83\%$	90.6%
✓	CCD-024	i Band QE	$> 82\%$	94.9%
✓	CCD-025	z Band QE	$> 75\%$	88.3%
✓	CCD-026	y Band QE	$> 21\%$	31.3%
✓	CCD-027	PRNU	$< 5\%$	max. variation = 4.06% at 350 nm
✓	CCD-028	Point Spread Function	$\sigma < 5\mu$	4.84μ



Conclusions

- LSST will be the premier astronomical project of the 21st century
- BNL is taking a big part