

The 44 years

of

Muon $g-2$

1956

Berestetskii, Krokhnin and Klebinkov,
Z.Eksp.Teor.Fiz. 30,788

$$\frac{\Delta a_{\mu}}{a_{\mu}} = \frac{2 m_{\mu}^2}{3 \Lambda^2}$$

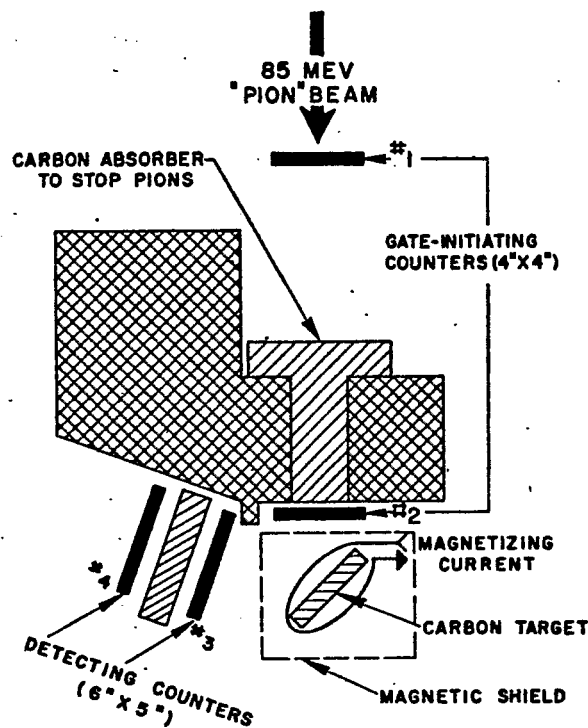
1957

Observations of the Failure of Conservation of Parity and Charge Conjugation in Meson Decays: the Magnetic Moment of the Free Muon*

RICHARD L. GARWIN,† LEON M. LEDERMAN,
AND MARCEL WEINRICH

*Physics Department, Nevis Cyclotron Laboratories,
Columbia University, Irvington-on-Hudson,
New York, New York*

(Received January 15, 1957)



Independent evidence that $\bar{g}=2$ (to $\sim 10\%$) comes from the coincidence of the polarization axis with the velocity vector of the stopped μ 's. This implies that the spin precession frequency is identical to the μ cyclotron frequency during the 90° net magnetic deflection of the muon beam in transit from the cyclotron to the 1-2 telescope.

1958 ANNUAL INTERNATIONAL CONFERENCE
ON HIGH ENERGY PHYSICS AT CERN

1958

Panofsky:

.....experiments were in progress at Columbia University and are being designed at the Joint Institute of Nuclear Research and in the University of Chicago

Marshak:

..... there may be another fundamental length coming in here since the μ meson is different from the electron. It may have a structure yielding deviations which are not quite so small.

Telegdi:

The fact is that everybody has been thinking about the (g-2) experiment

..... common impression that the first order term is well accepted and the real interest is to go beyond to the next order which is known to be different for electrons and muons.

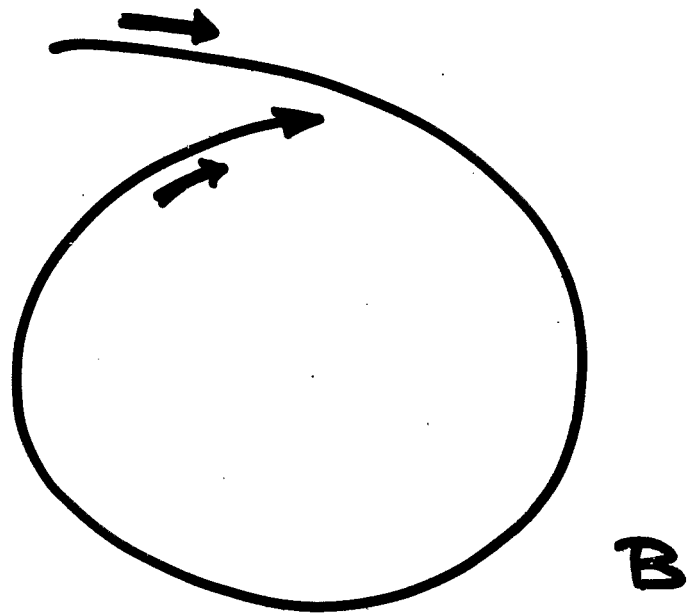
NOW I understand that some theorists feel that maybe even the first order correction might be affected for the μ meson by as much as 10%.

Tamm:

Of course for the μ mesons we can expect quite new effects because of the large mass of the μ meson has to be explained and they may be coming in far earlier than in the case of the electron.

g-2 precession

Non relativistic



$$\omega_c = \frac{eB}{mc}$$

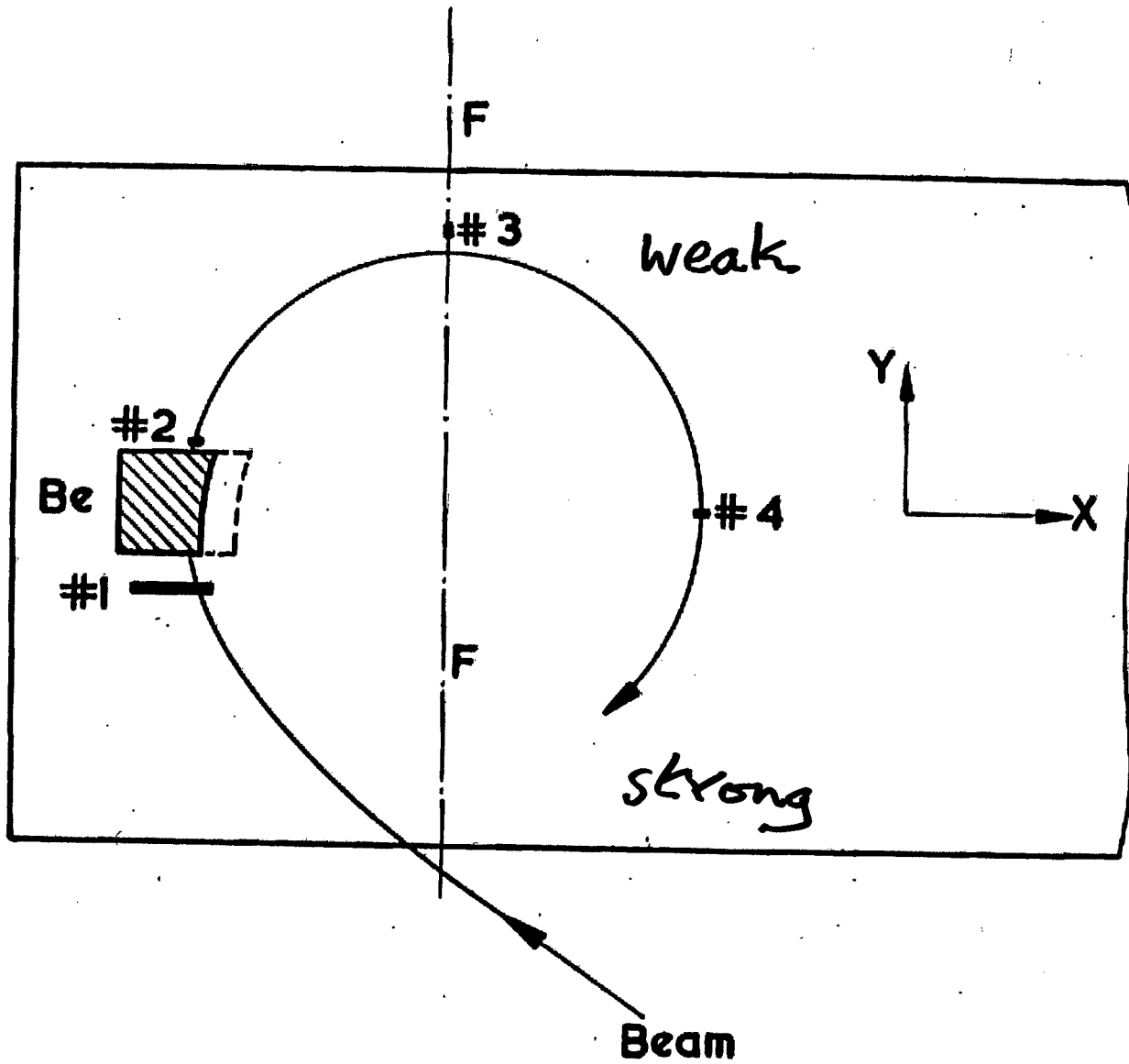
$$\begin{aligned}\omega_s &= g \cdot \frac{e}{2mc} B \\ &= (1+a) \frac{eB}{mc}\end{aligned}$$

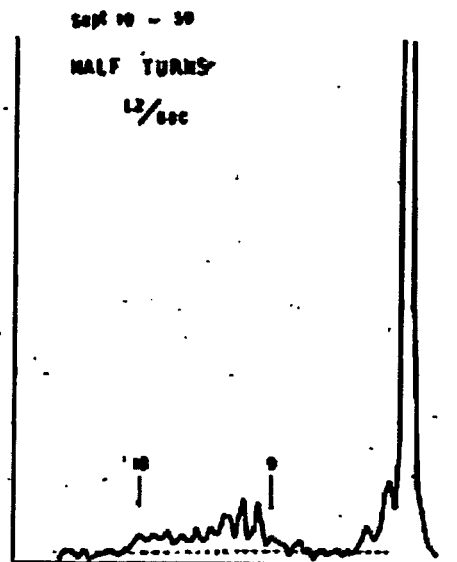
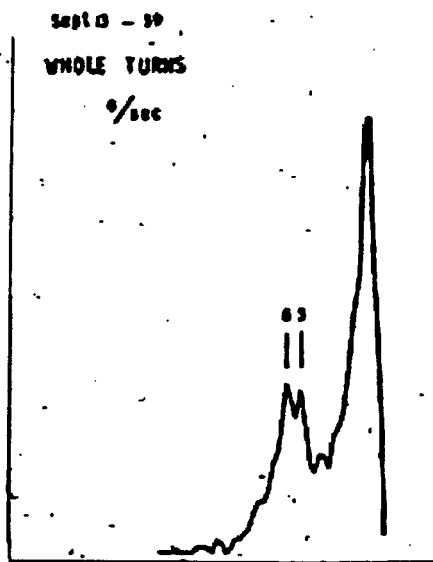
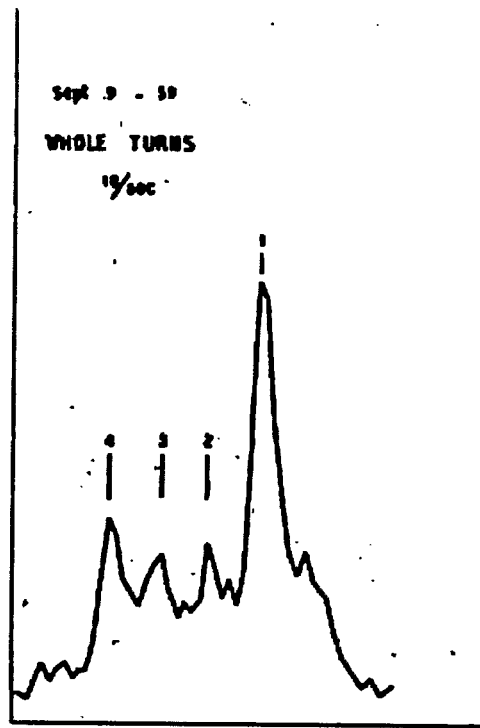
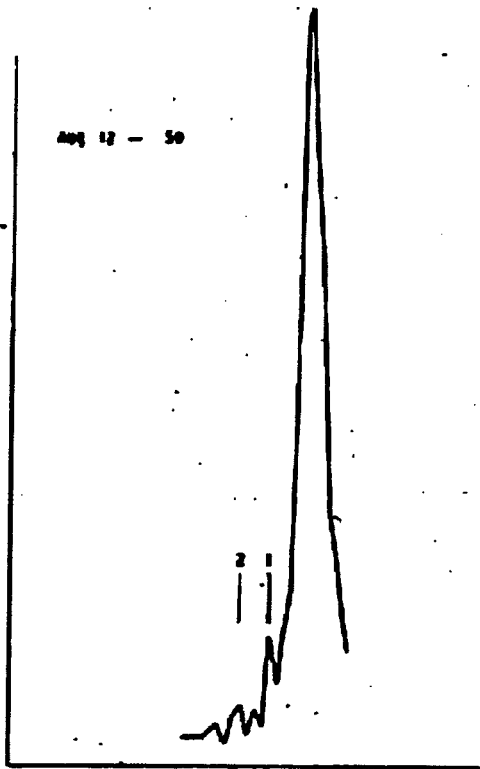
$$\omega_a = \omega_s - \omega_c = a \frac{eB}{mc}$$

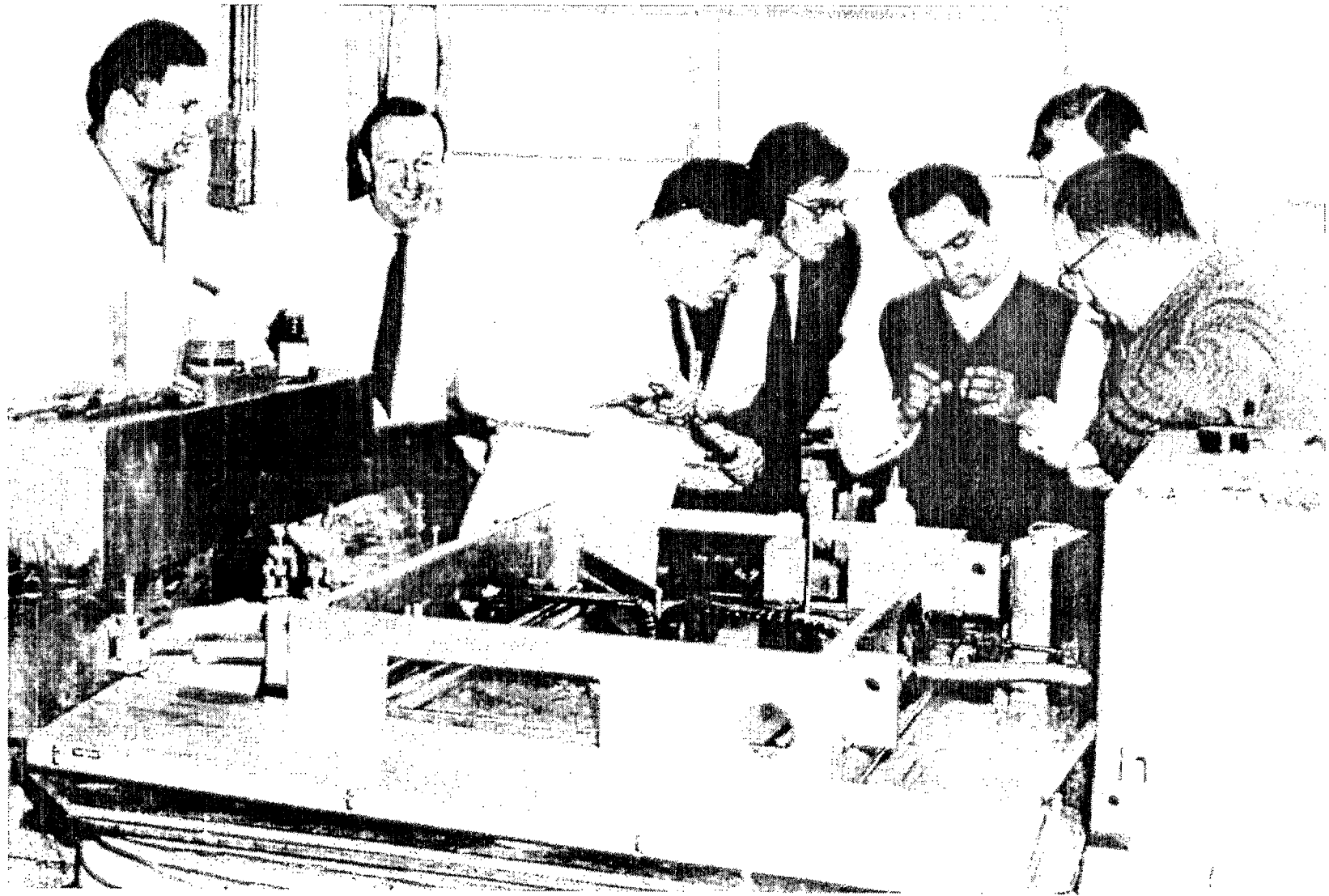
$$\frac{a}{1+a} = \frac{\omega_a}{\omega_s} = \frac{\omega_a / \omega_p}{\omega_s / \omega_p} = \frac{R}{\lambda}$$

$$a = \frac{R}{\lambda - R}$$

16 kG spin turns once in $4 \mu\text{s}$
But $\tau = 2.2 \mu\text{s} \rightarrow 64 \mu\text{s}$

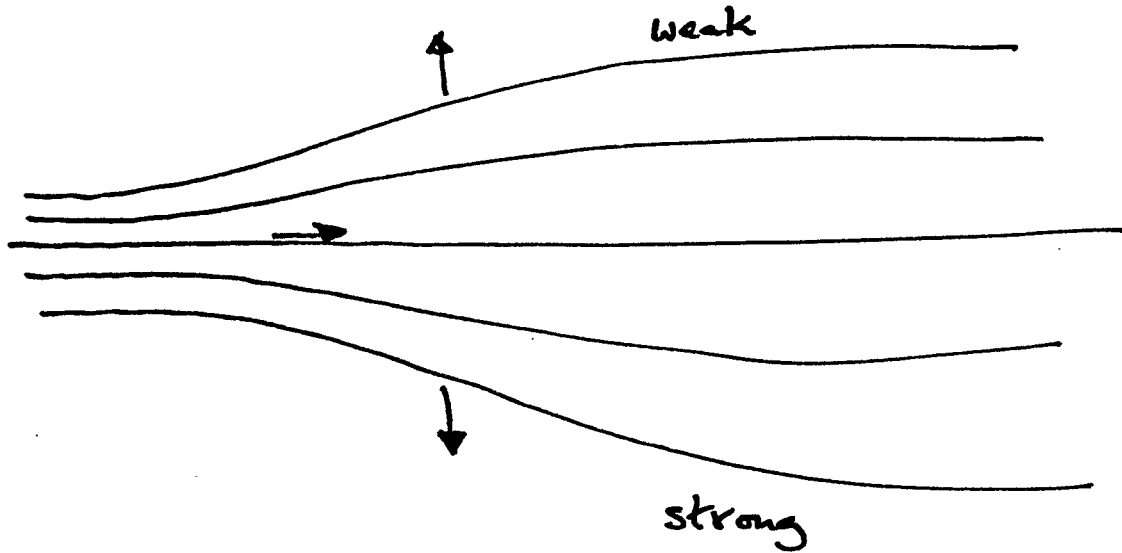




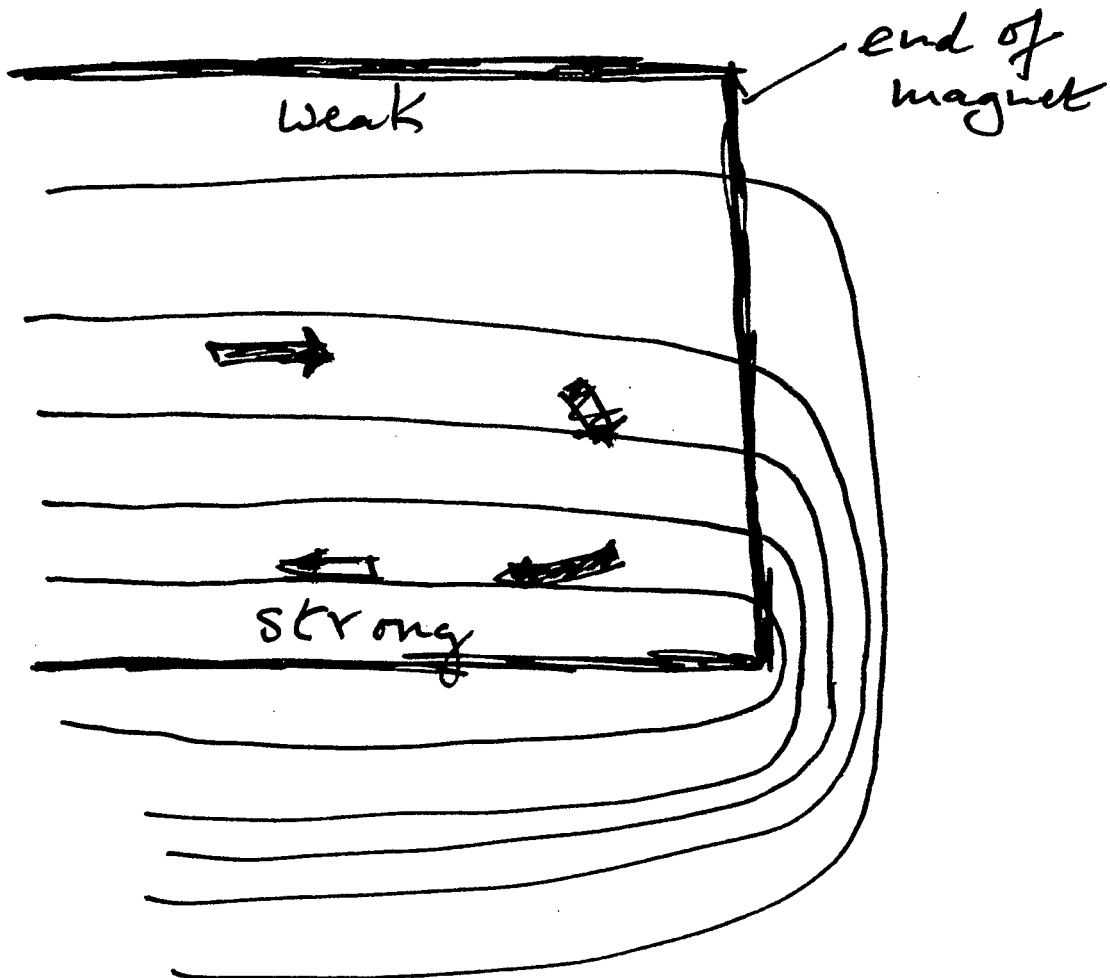


Orbit walks \perp gradient
follows contour lines

Flux through orbit is invariant



orbit smells end of magnet & goes back!



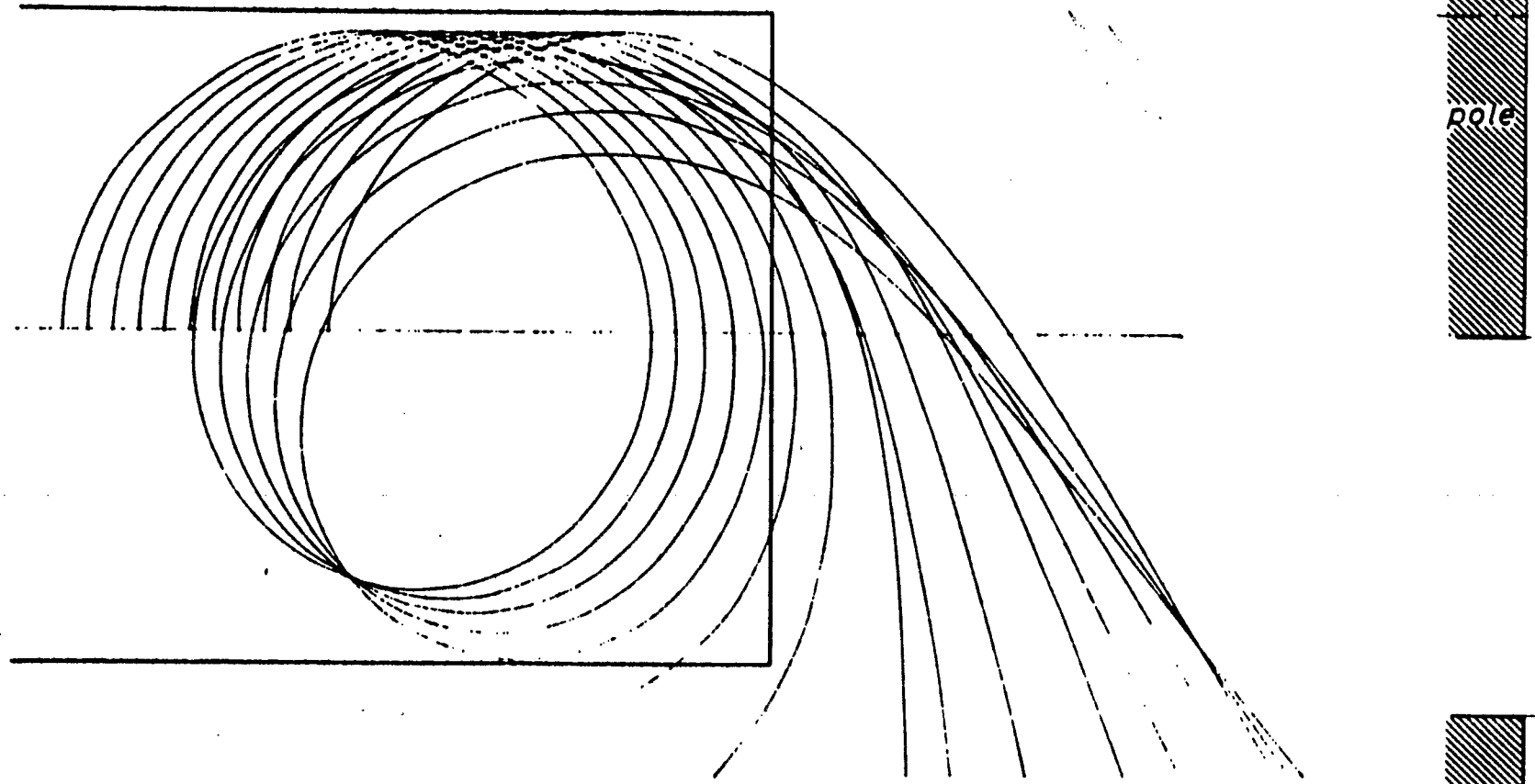


Fig. 6. - Computed orbits at the ejection end of the magnet based on detailed field measurements with the iron block (Fig. 61) in position. The downward movement in the last turn is due to the falling off of the field at the end of the magnet; but this actually improves the ejection by bringing the orbit away from the corner post of the vacuum chamber.

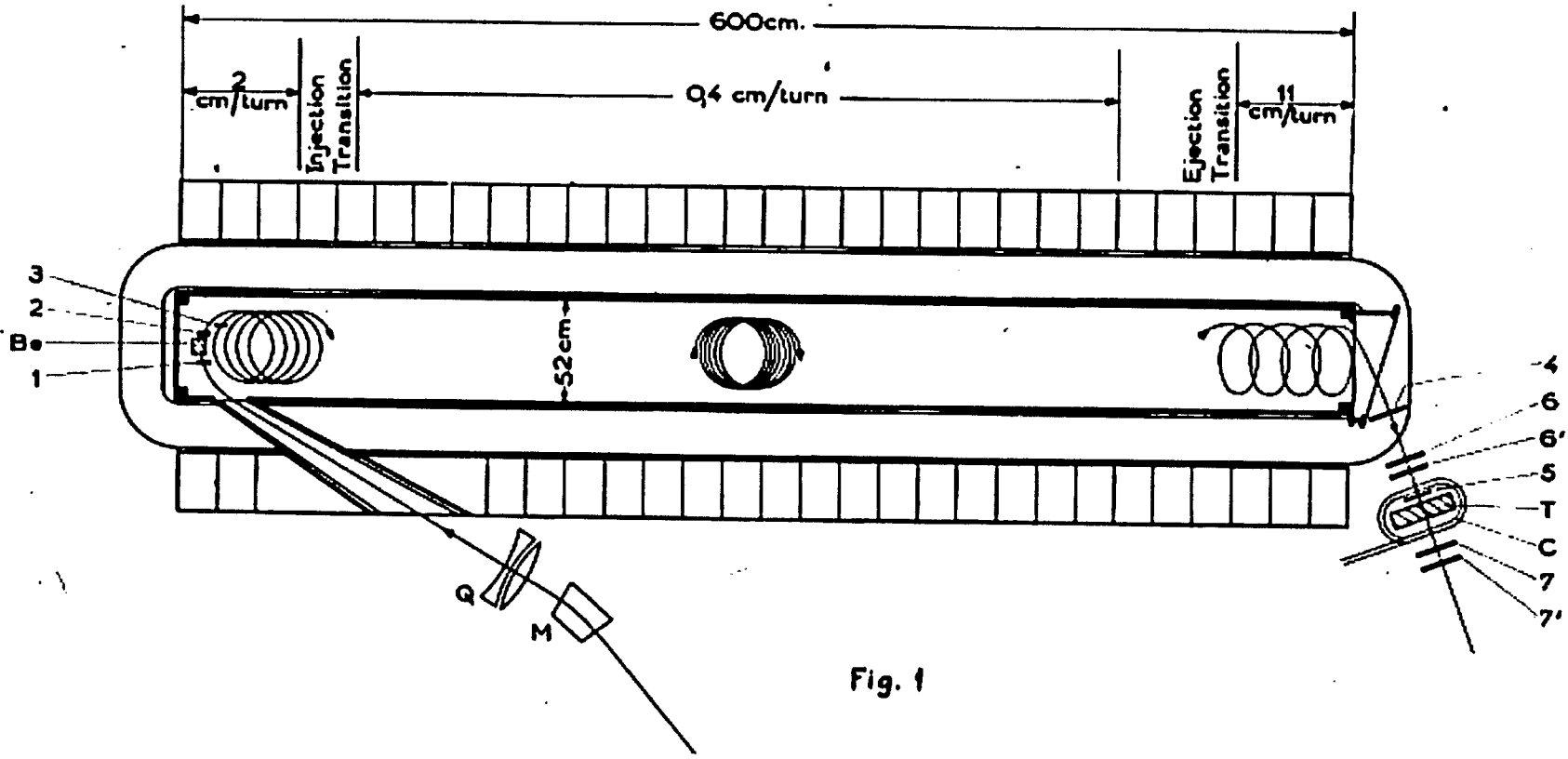
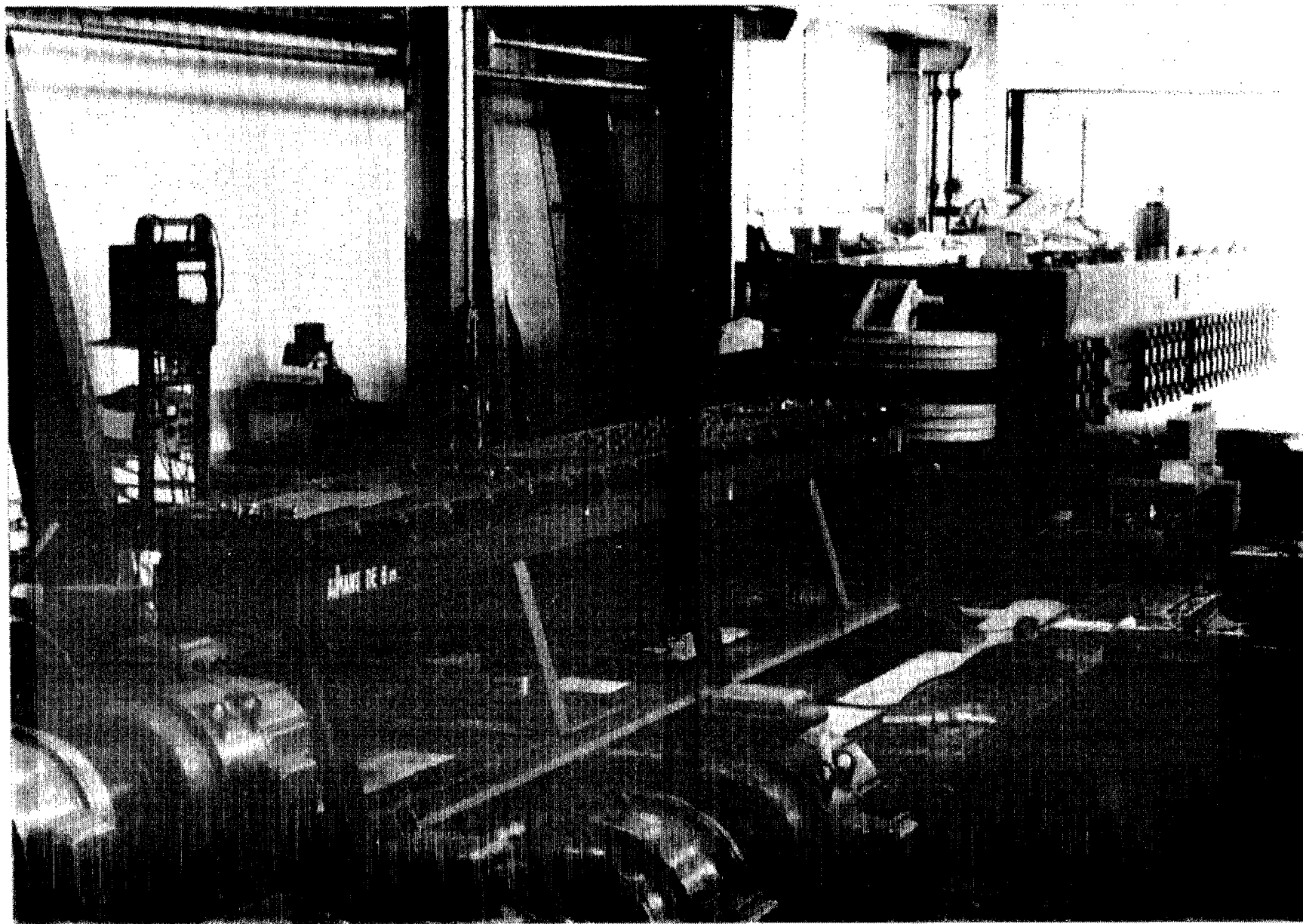
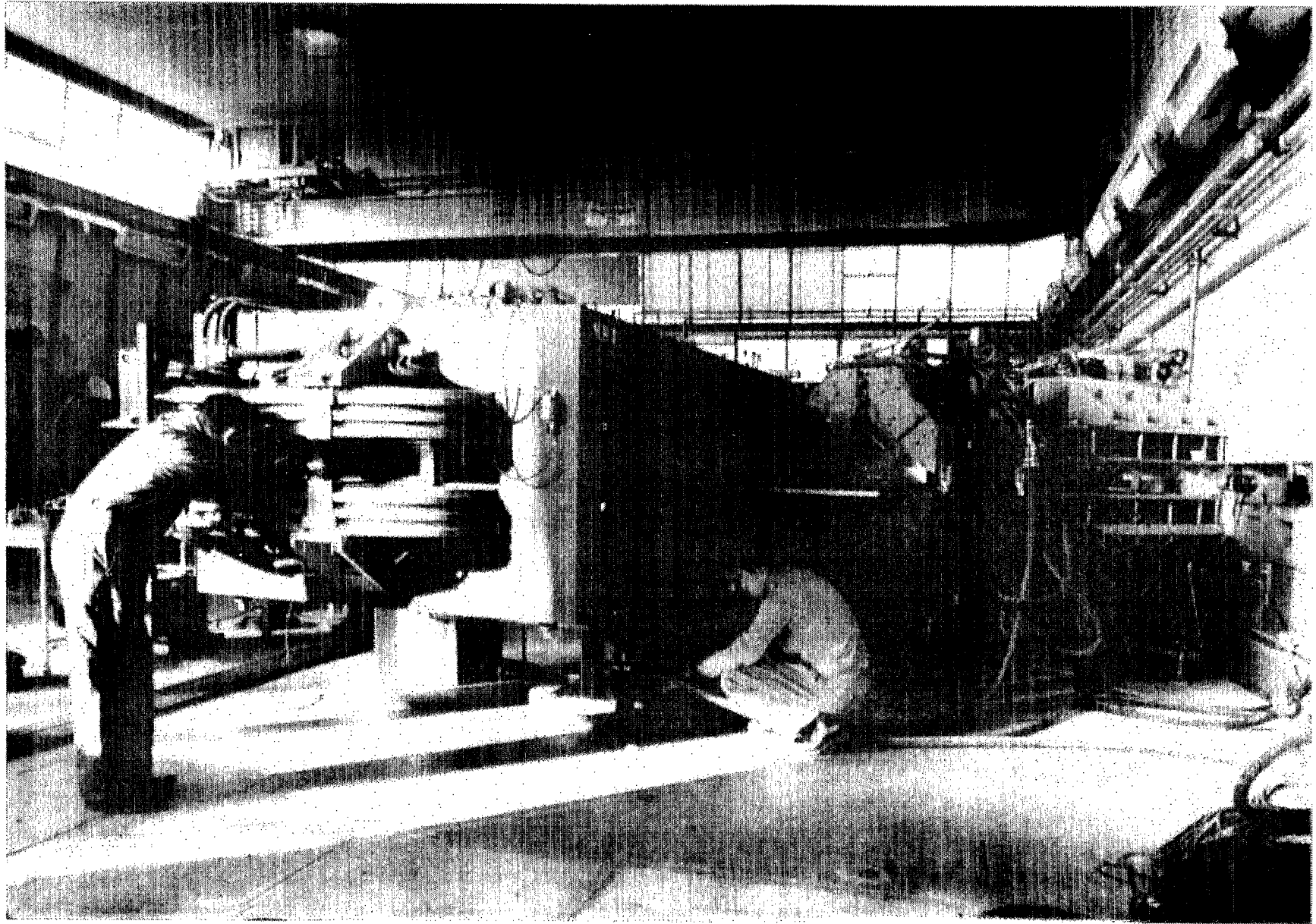


Fig. 1





Shims-profile for injection field

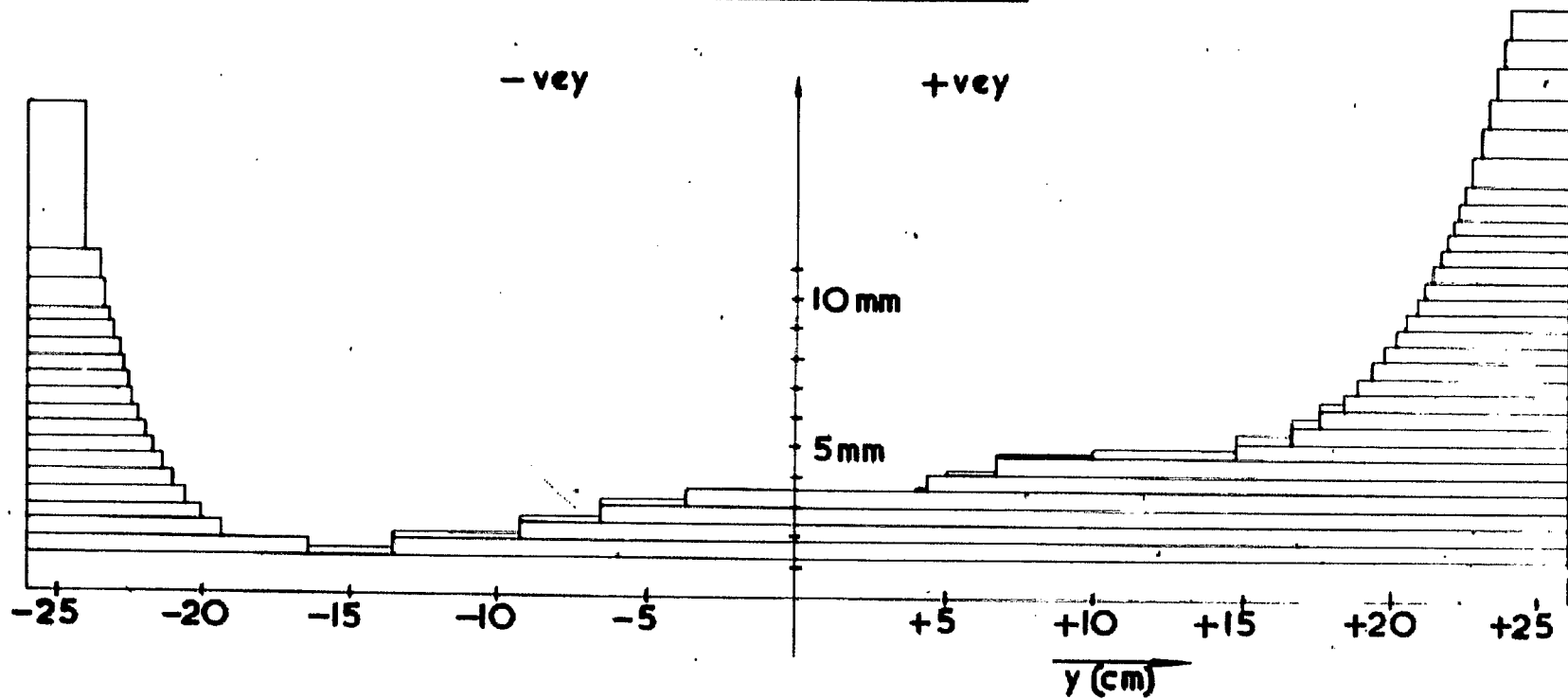


Fig. 13

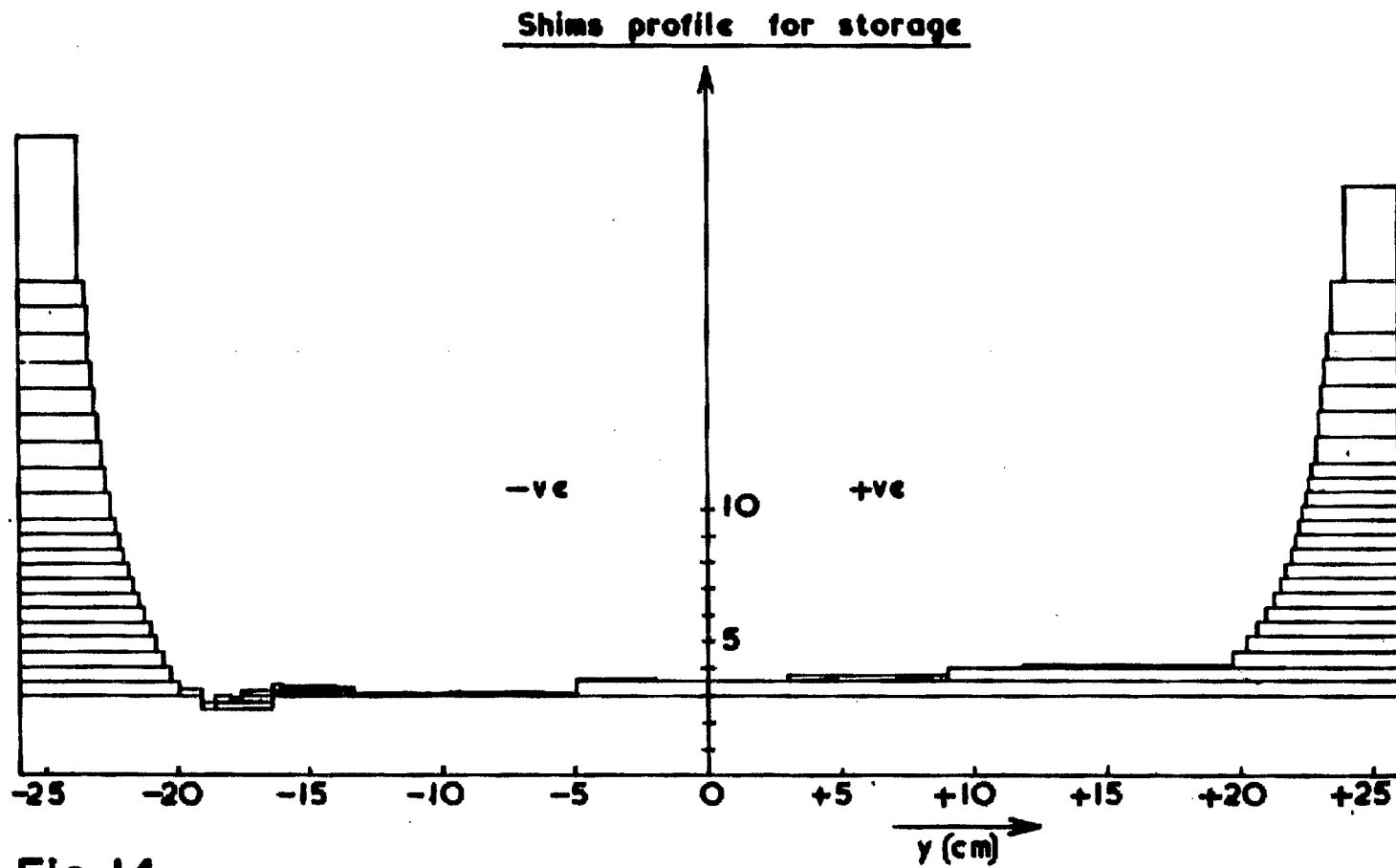


Fig. 14

Shms profile for ejection field ($n=7 \times 10^{13} \text{ cm}^{-3}$)

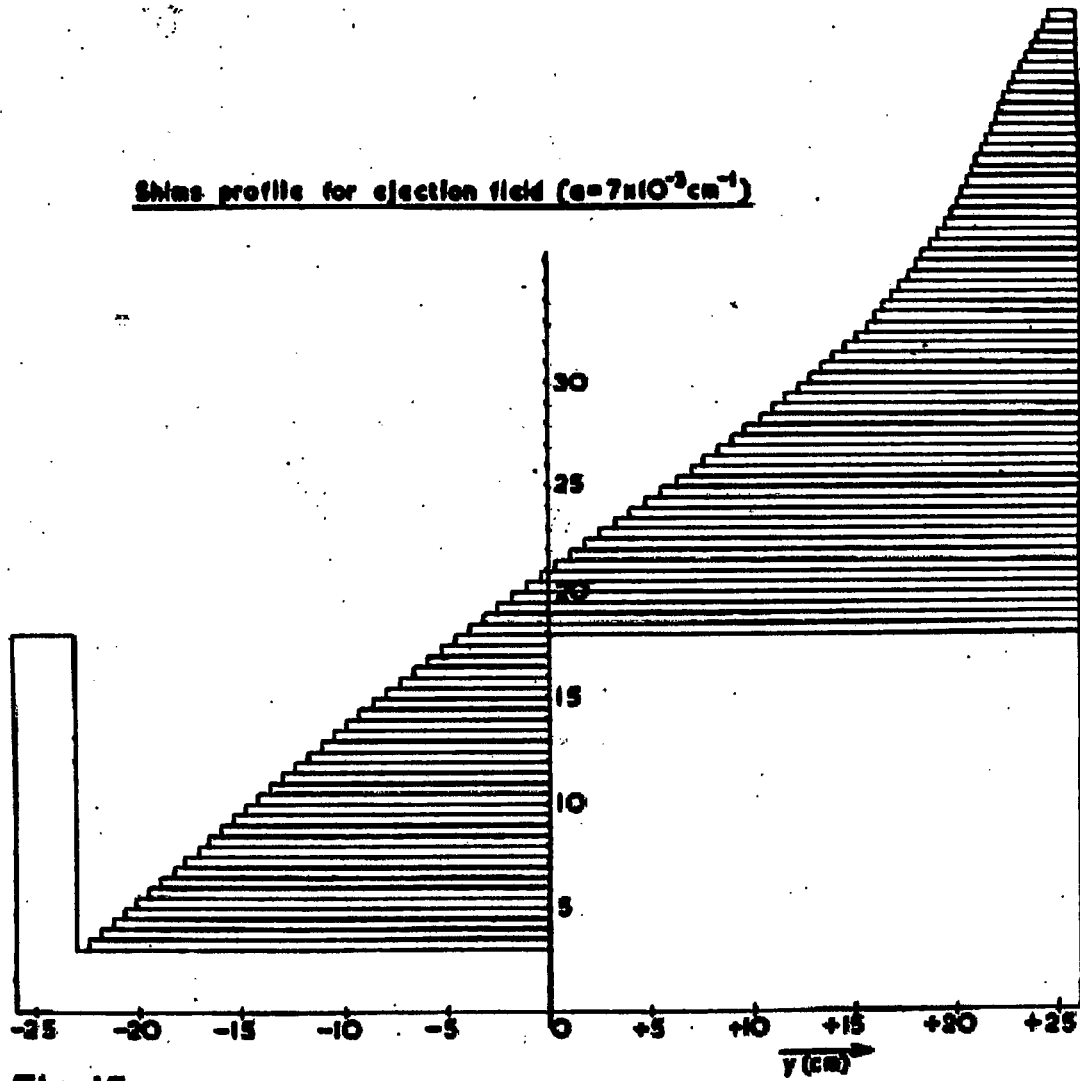
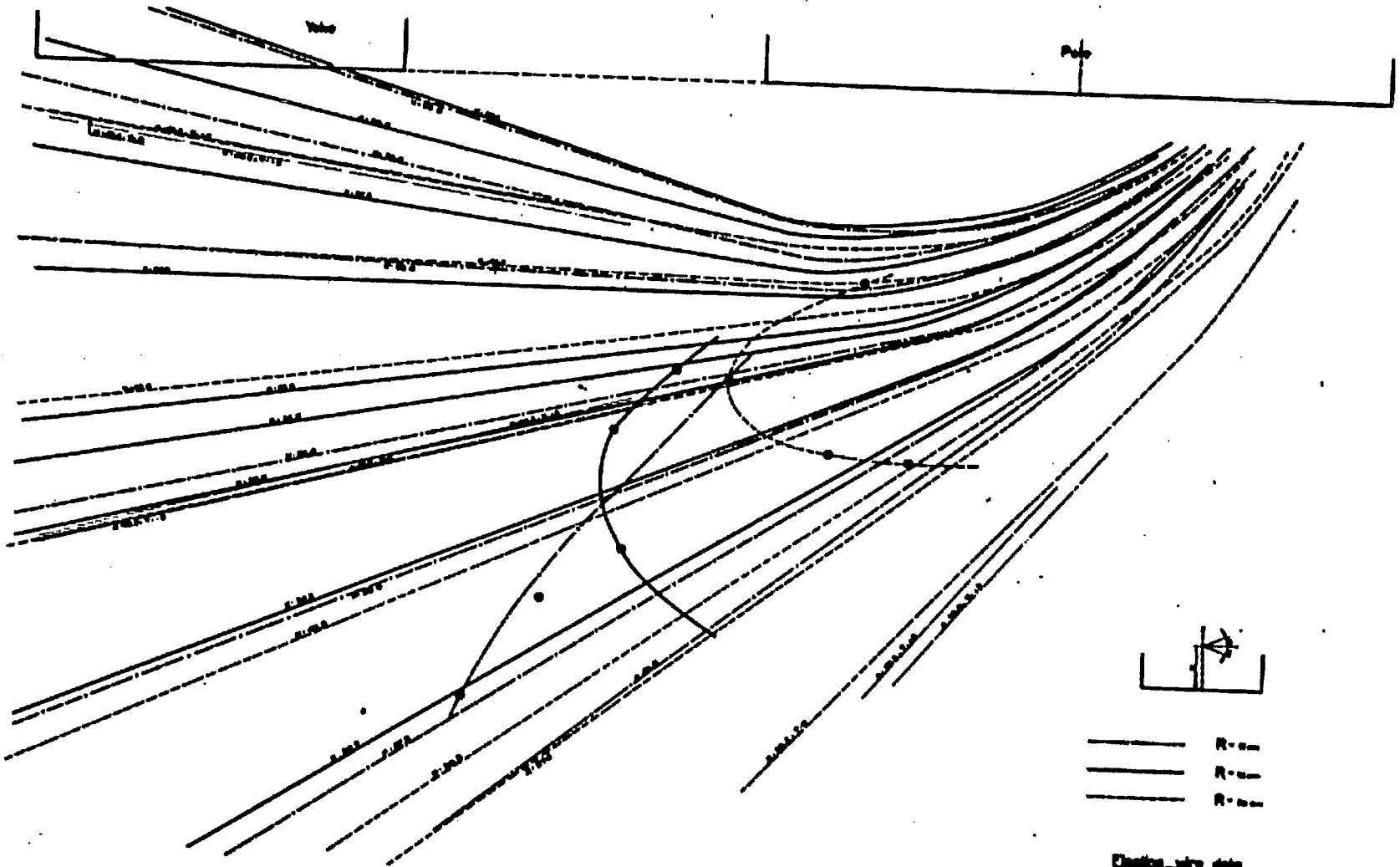
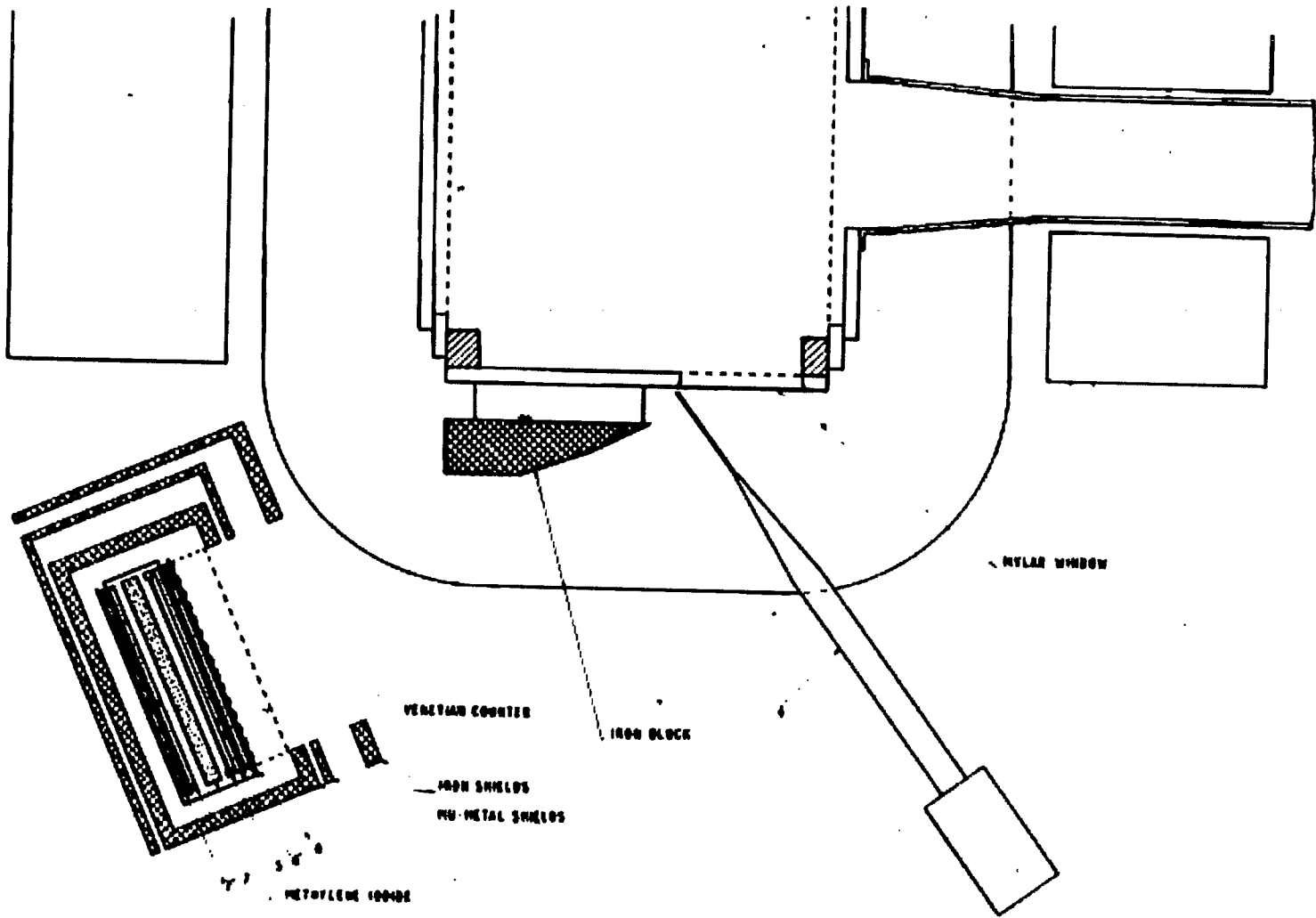


Fig. 15



- R-00
- - - R-00
- · - R-00

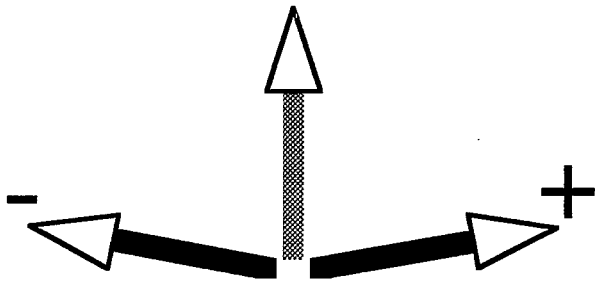
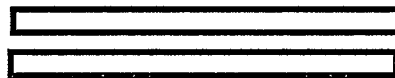
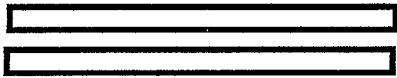
Change wire data
specimen 1111



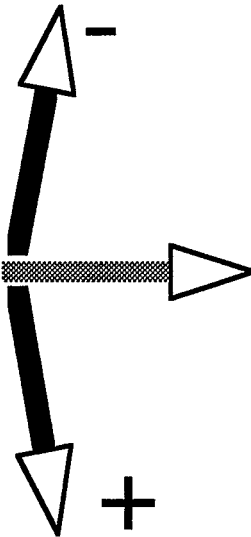
Flip spin $\pm 90^\circ$ after muon stops

$$A = \frac{N_+ - N_-}{N_+ + N_-}$$

Forward telescope



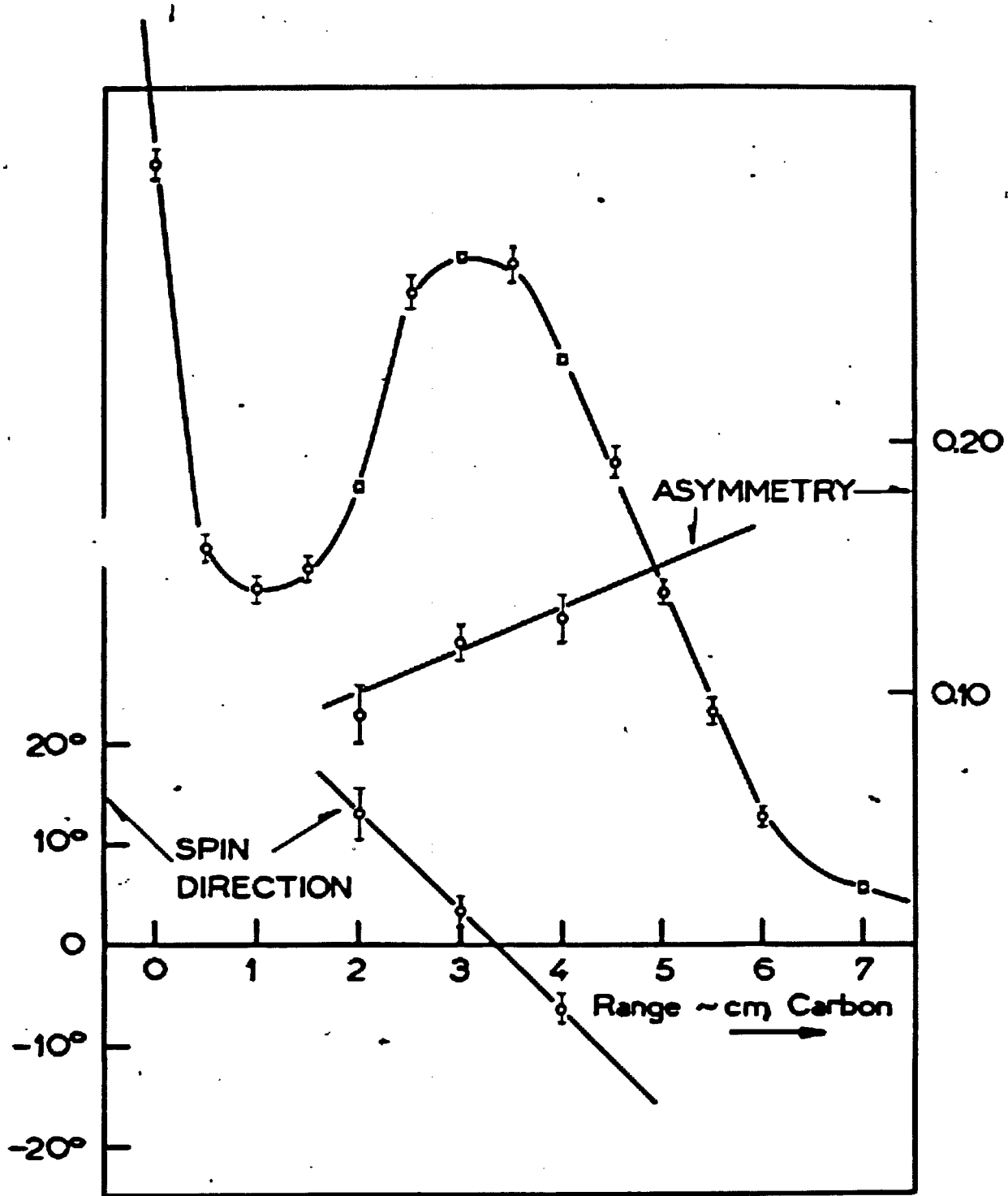
Spin forward - no asymmetry

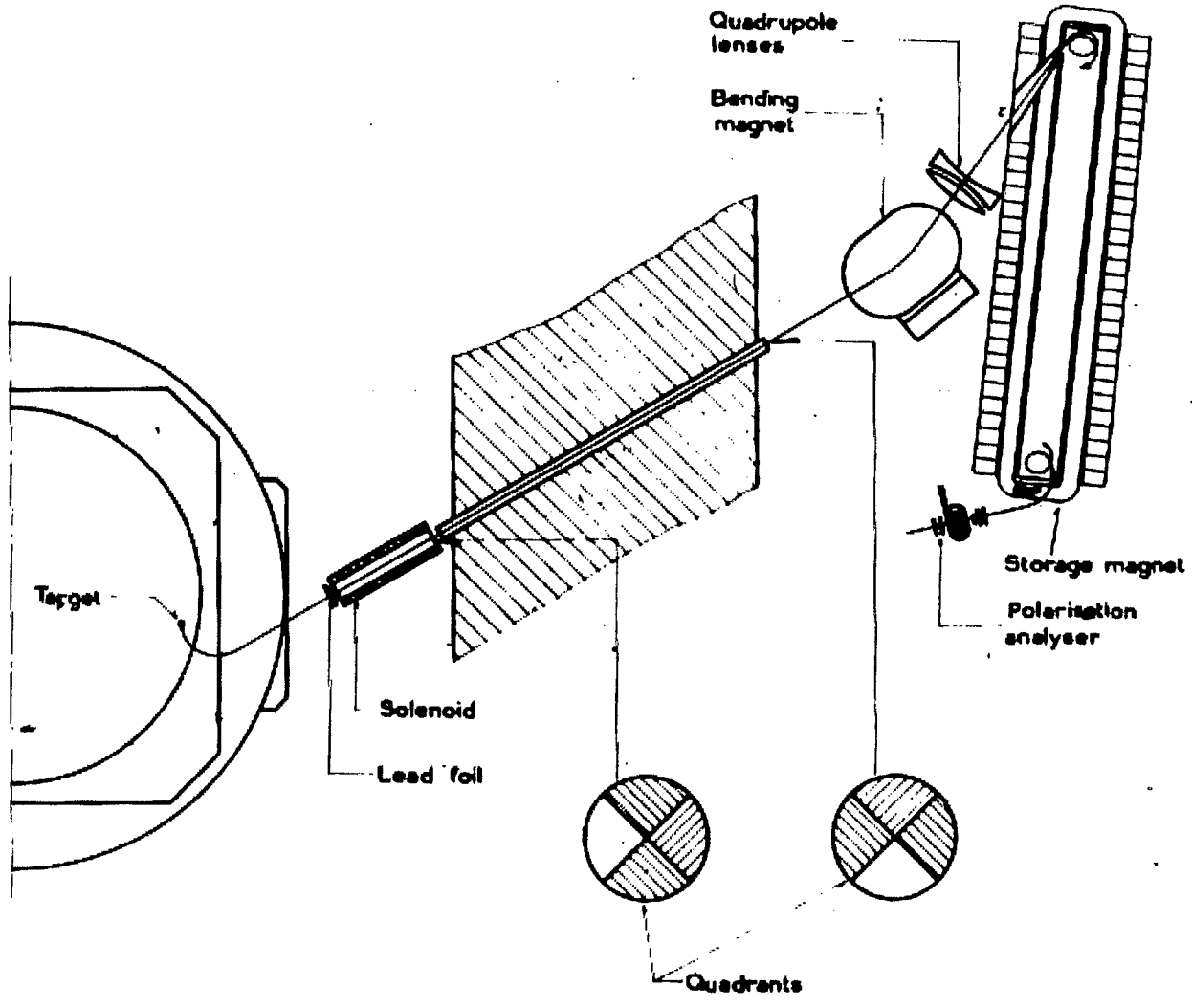


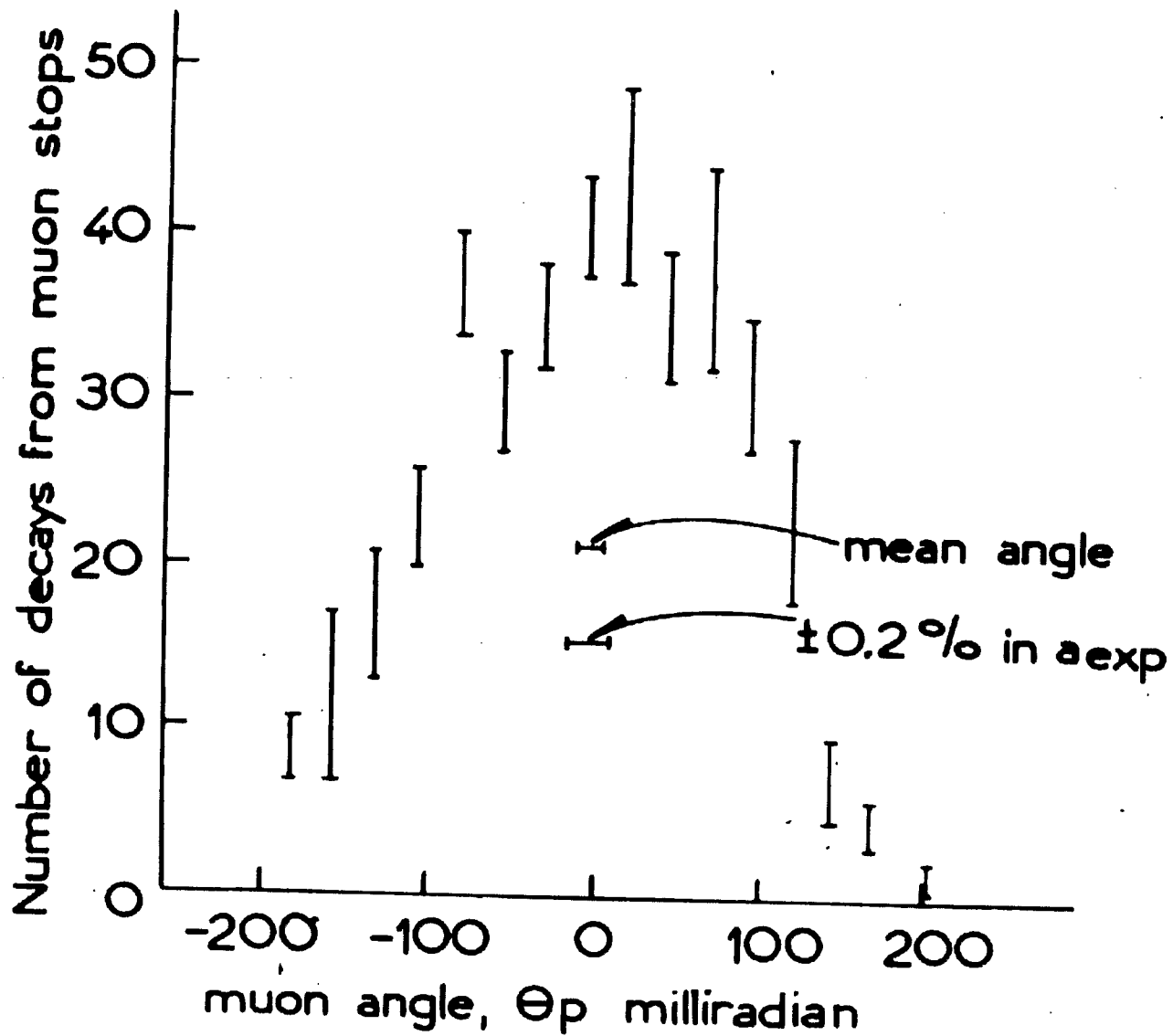
Spin transverse - large asymmetry

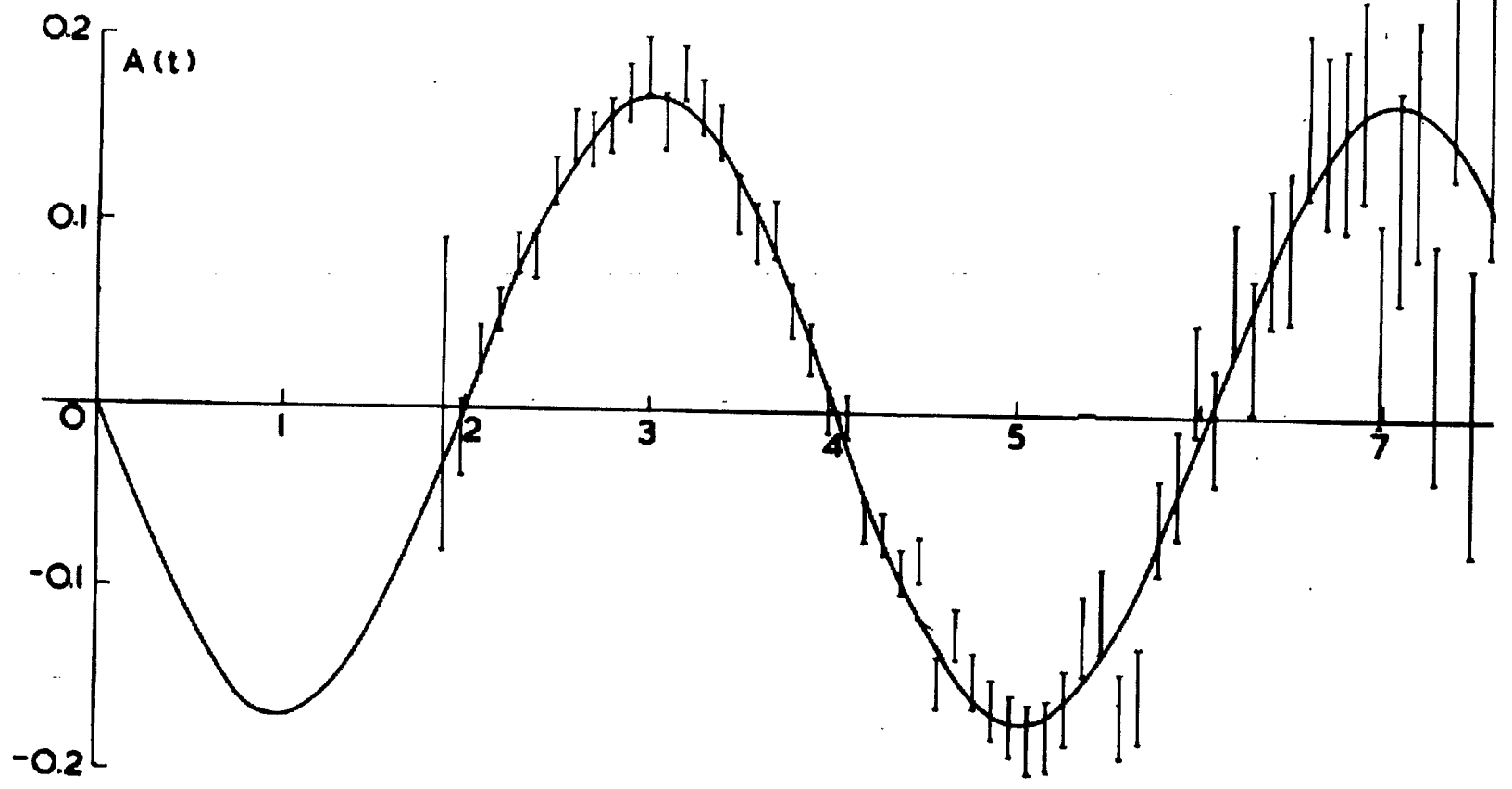
Advantages:

- Flip angle need not be exact
- Telescope efficiency irrelevant









$$A = \frac{n_+ - n_-}{n_+ + n_-}$$

$$= A_0 \sin \theta_s$$

$$= A_0 \sin \left\{ a \frac{e}{mc} \bar{B} t + \theta_P + \theta_s^0 - \theta_P^0 \right\}$$

final direction

initial polarization angle

$\left. \begin{array}{l} |B| \\ \theta_P \\ \theta_s^0 - \theta_P^0 \end{array} \right\}$

are all functions of
orbit radius (range)
& storage time

$$a = 1162 \pm 5 \times 10^{-6}$$

- Charpak, Farley, Garwin, Muller, Sens, Zichichi
Nuovo Cimento 37, 1241 (1965)
- Bailey, Bartl, von Bochmann, Brown, Farley,
Giesch, Jöstlein, van der Meer, Picasso, Williams
Nuovo Cimento 9A, 369 (1972)
- Bailey, Borer, Combley, Drumm, Eck, Farley,
Field, Flegel, Hattersley, Krienen, Lange,
Lebée, McMillan, Petrucci, Picasso,
Runolfsson, von Ruden, Williams, Wójcicki
Nuclear Physics B150, 1 (1979)

Review -

- Farley & Picasso -
in "Quantum Electrodynamics"
T. Kinoshita (ed)
World Scientific, 1990
- g-2 note #96

Muon now seen as
heavy electron

QED works
good to any order.

Can we do better?

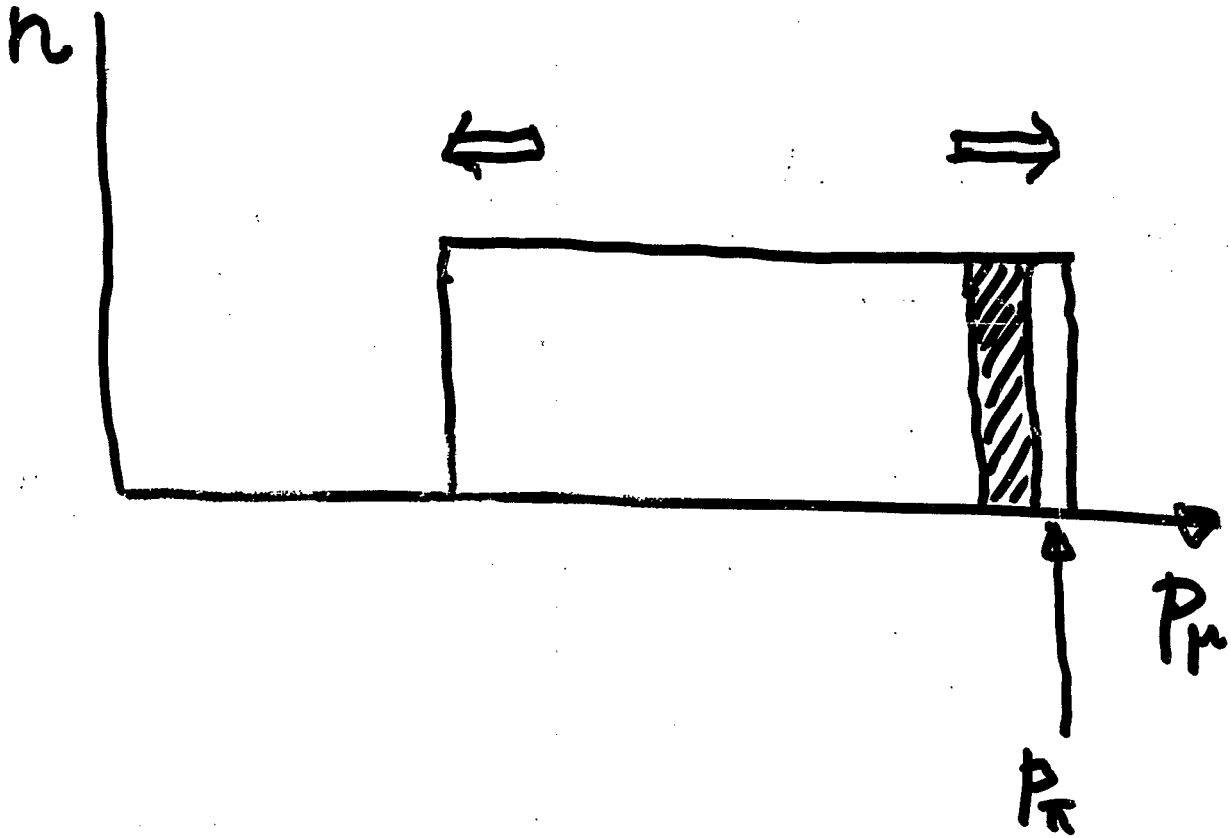
PS is now running

Dilate lifetime to get
more $g-2$ cycles

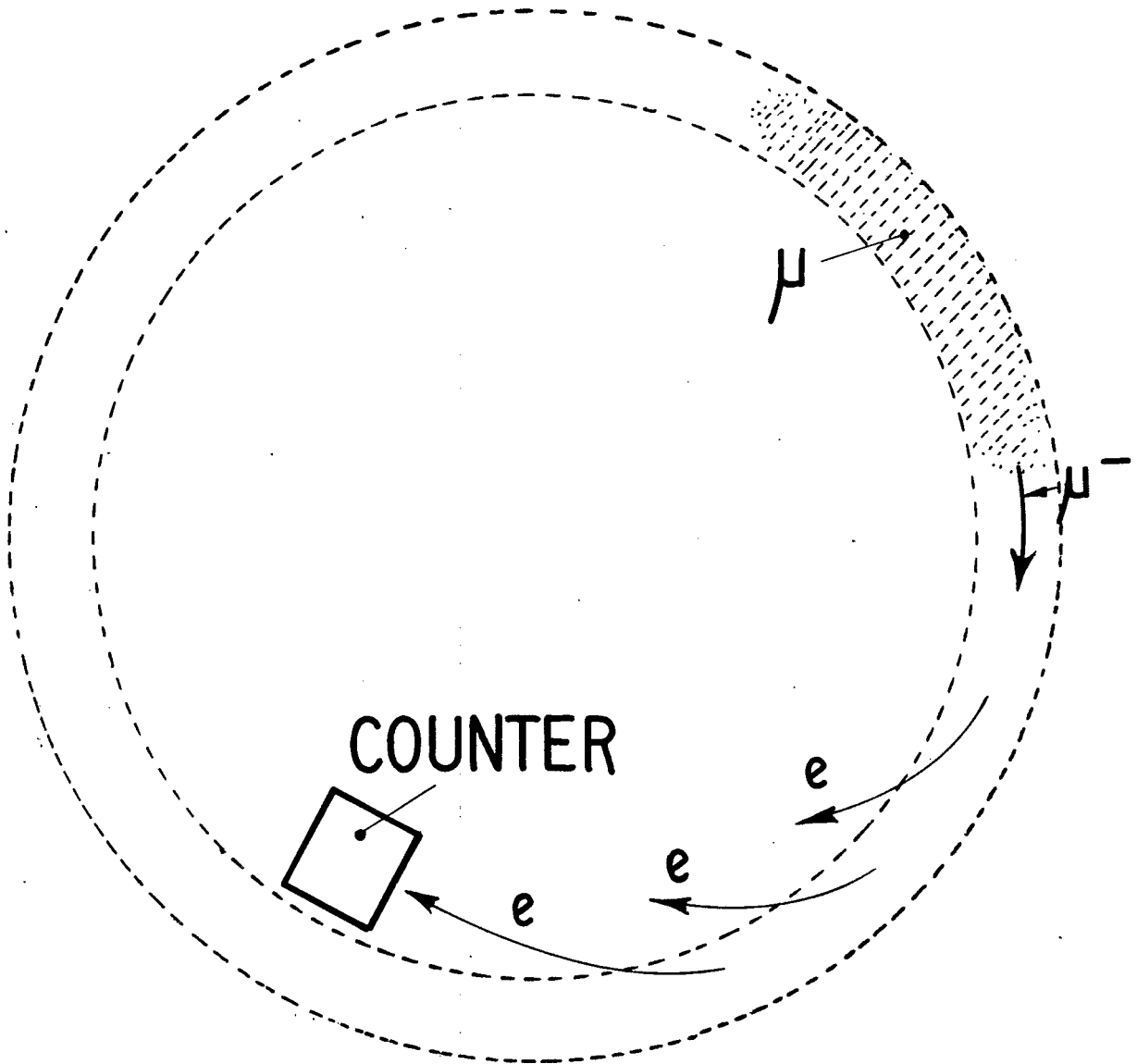
Four miracles of Nature

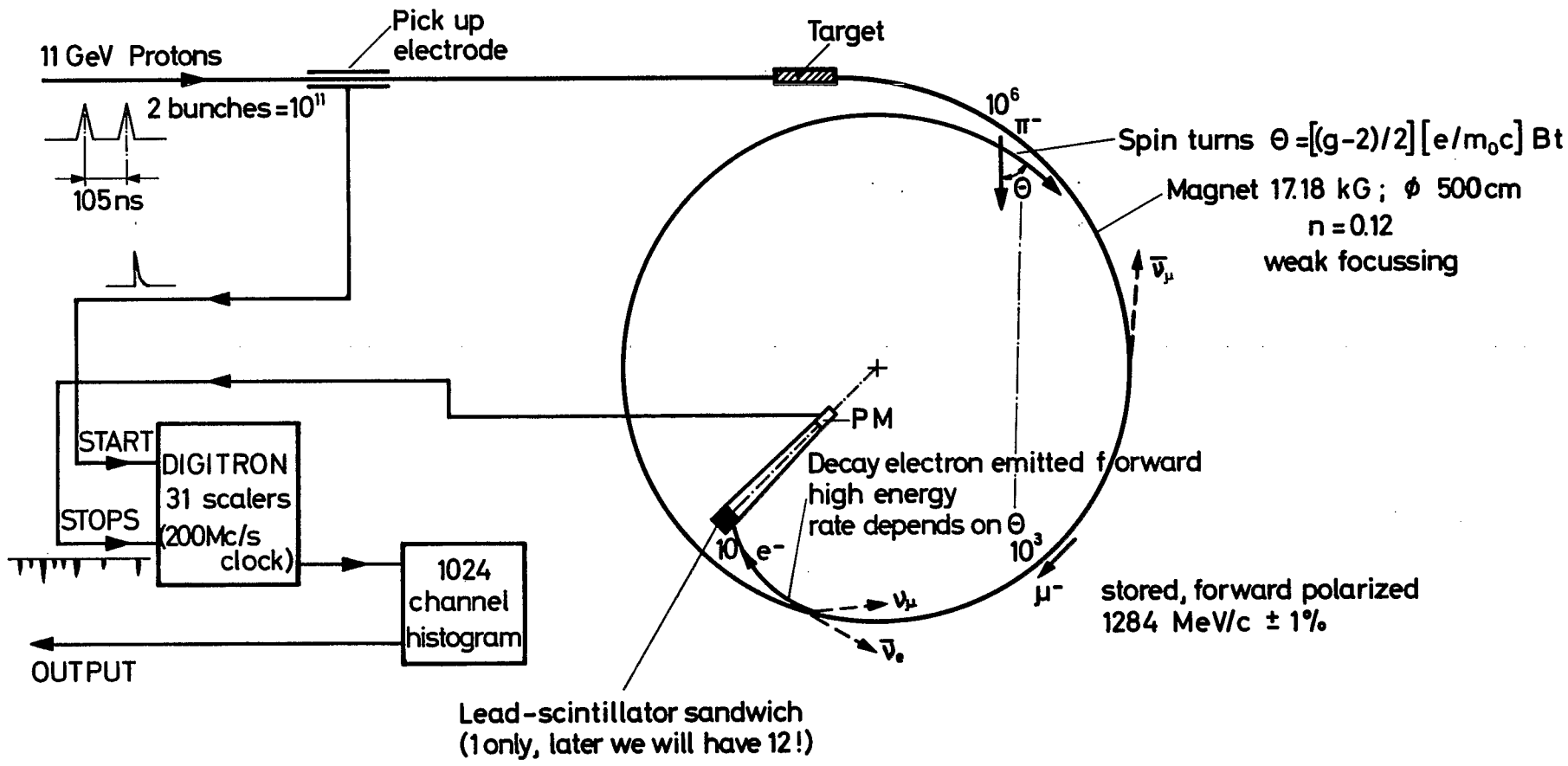
- 1) Easy to store muons in a magnetic ring. Just inject pions and wait.
- 2) The stored muons are longitudinally polarised
- 3) Decay electrons come out on the inside of the ring and tell us which way the spin is pointing.

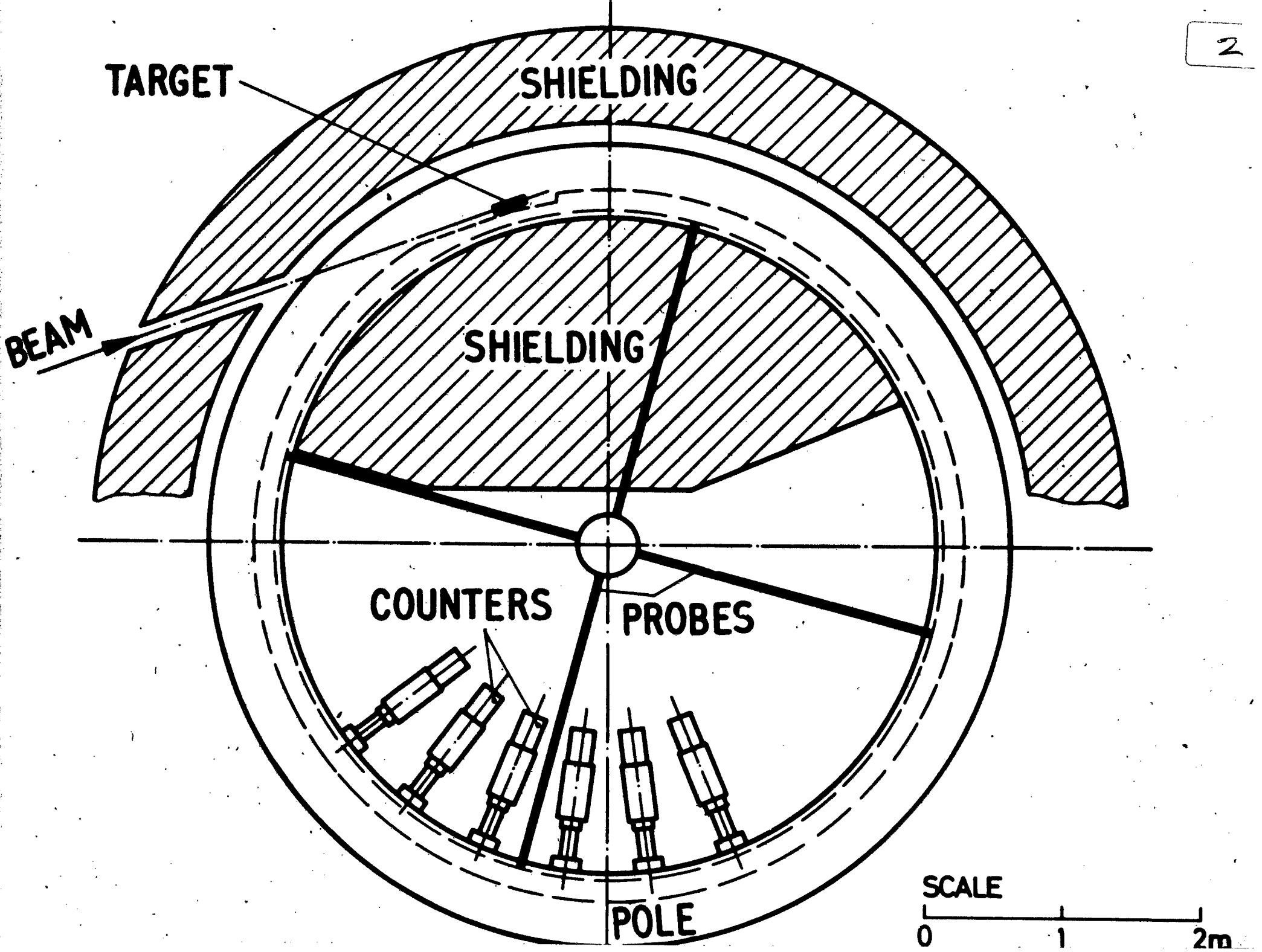
Inject by $\pi - \mu$ decay

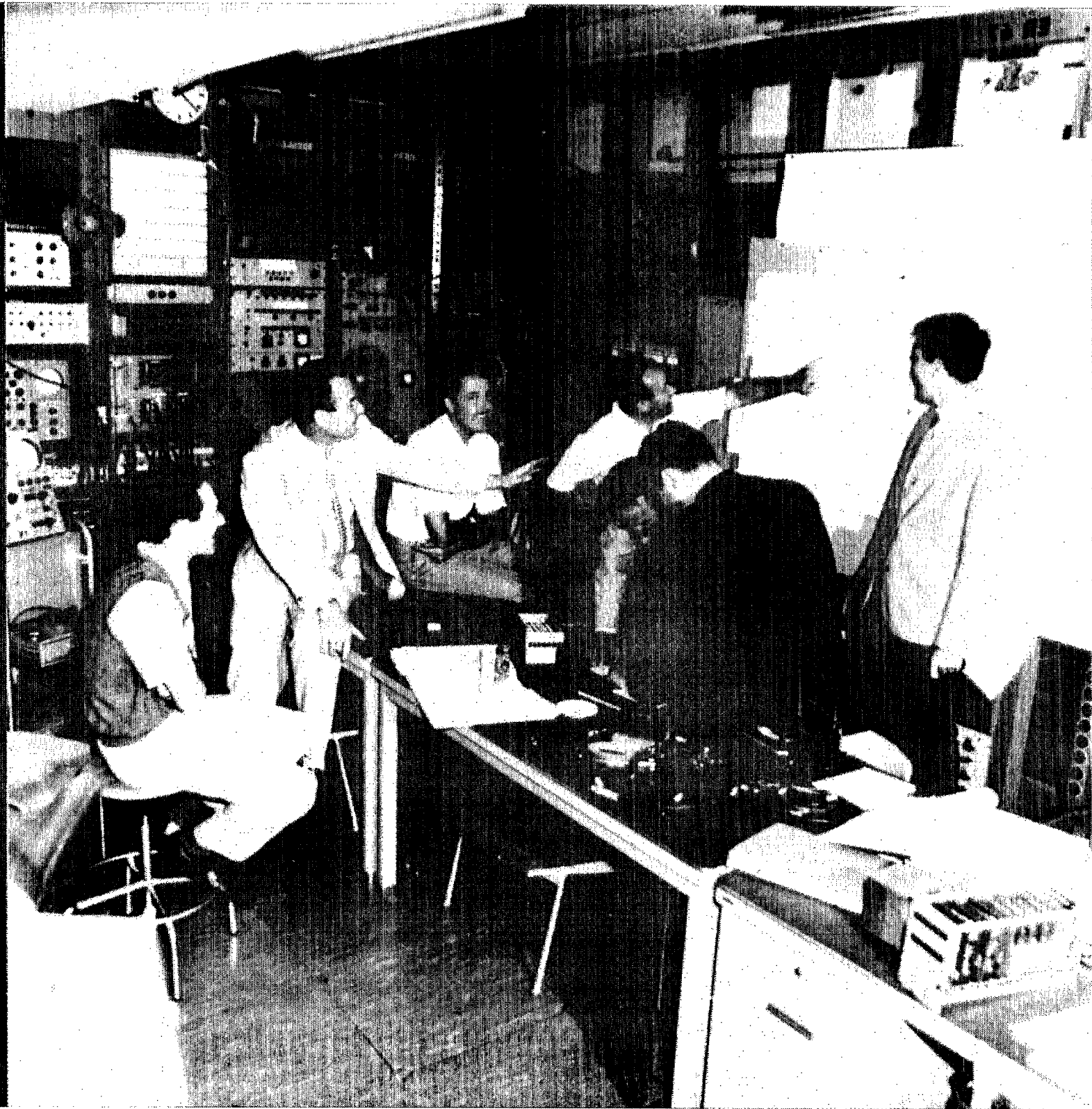


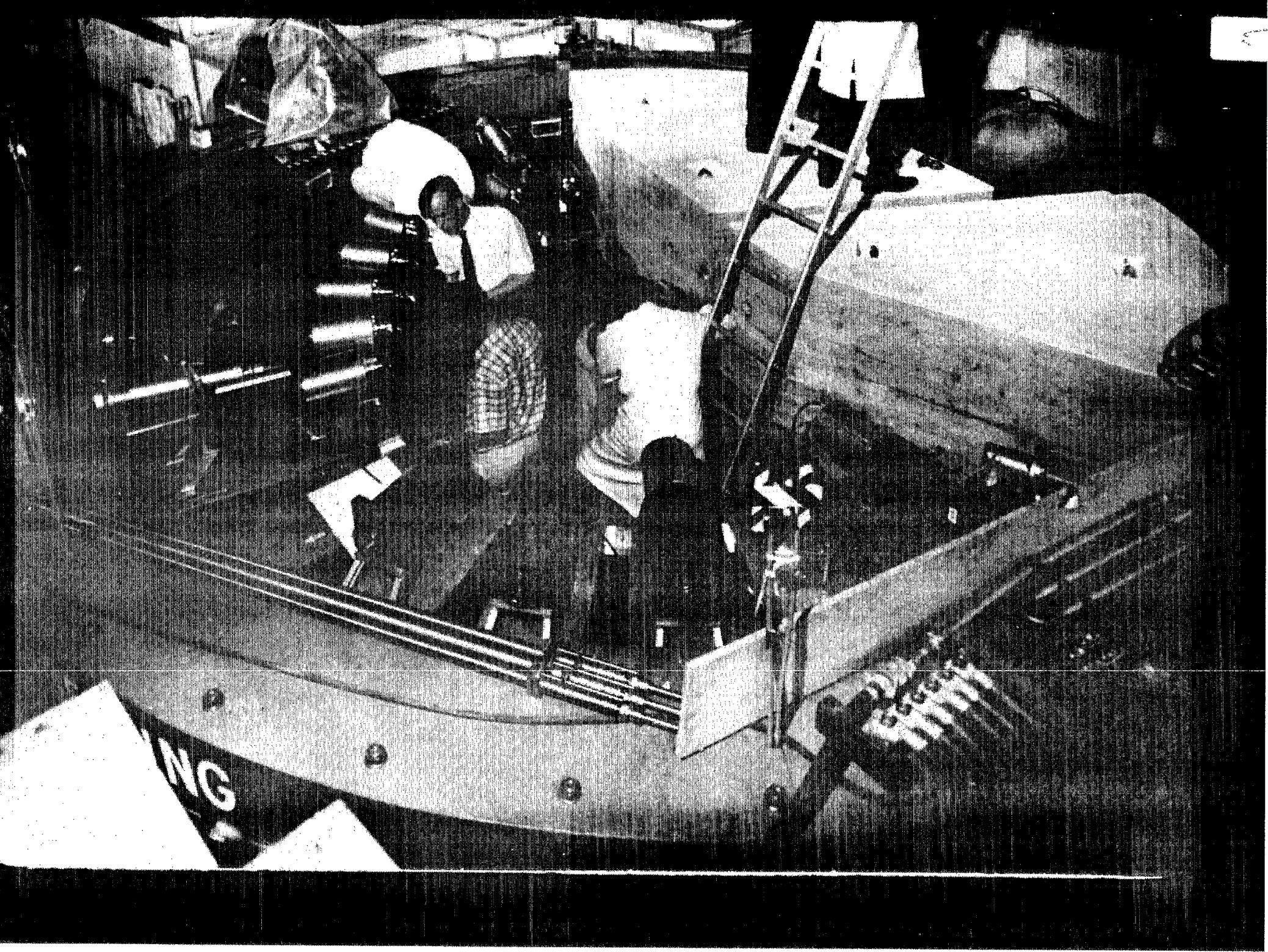
- Inject π for $\frac{1}{2}$ turn
- Forward polarized μ are automatically stored

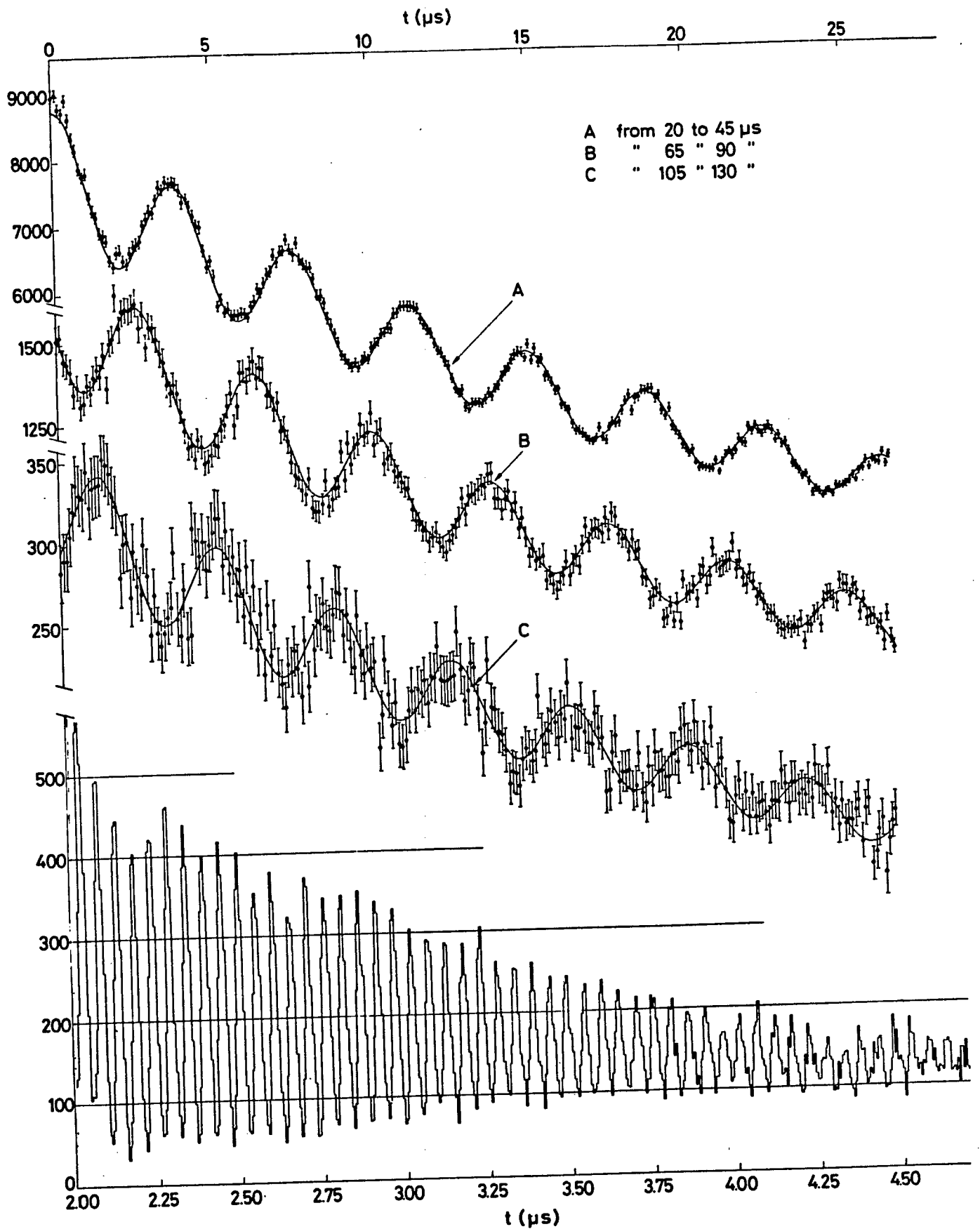


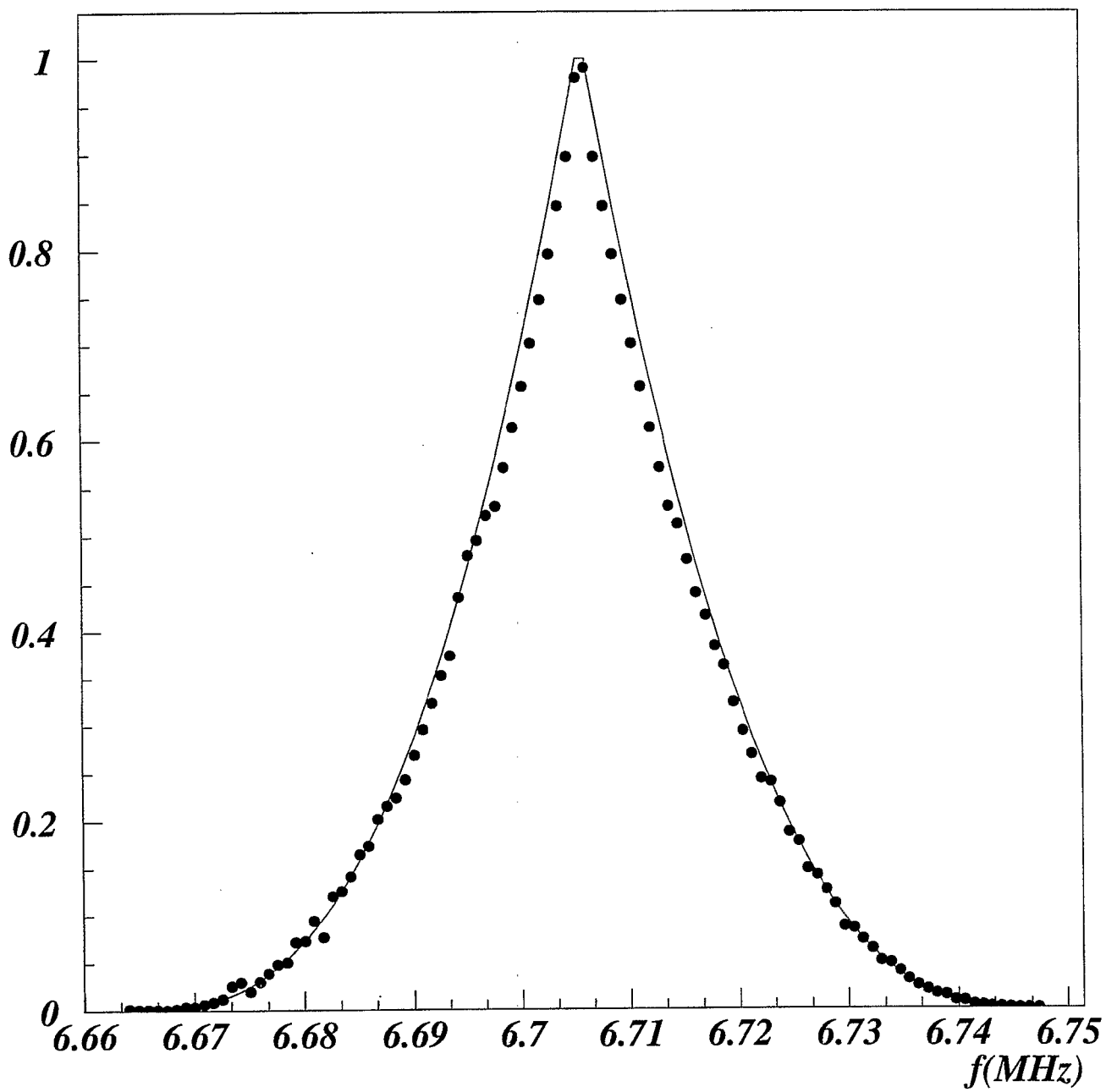












$$a = 1162 \pm 5 \times 10^{-6}$$

$$a = 116\,616 \text{ (31)} \times 10^{-8}$$

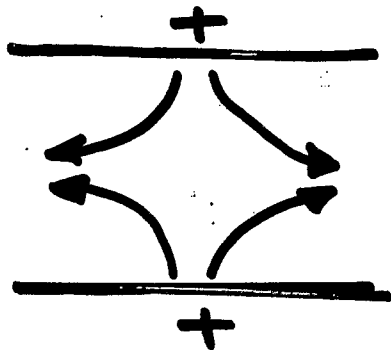
$$450 \pm 270 \quad \text{ppm}$$

$$\longrightarrow 240 \pm 270 \quad \text{ppm}$$

14

Vertical focusing

- Magnetic gradient would create error in B
- Use electric quadrupole



- Radial electric field does not affect spin motion if

$$\gamma = \sqrt{1 + \frac{1}{\alpha}} = 29.3$$

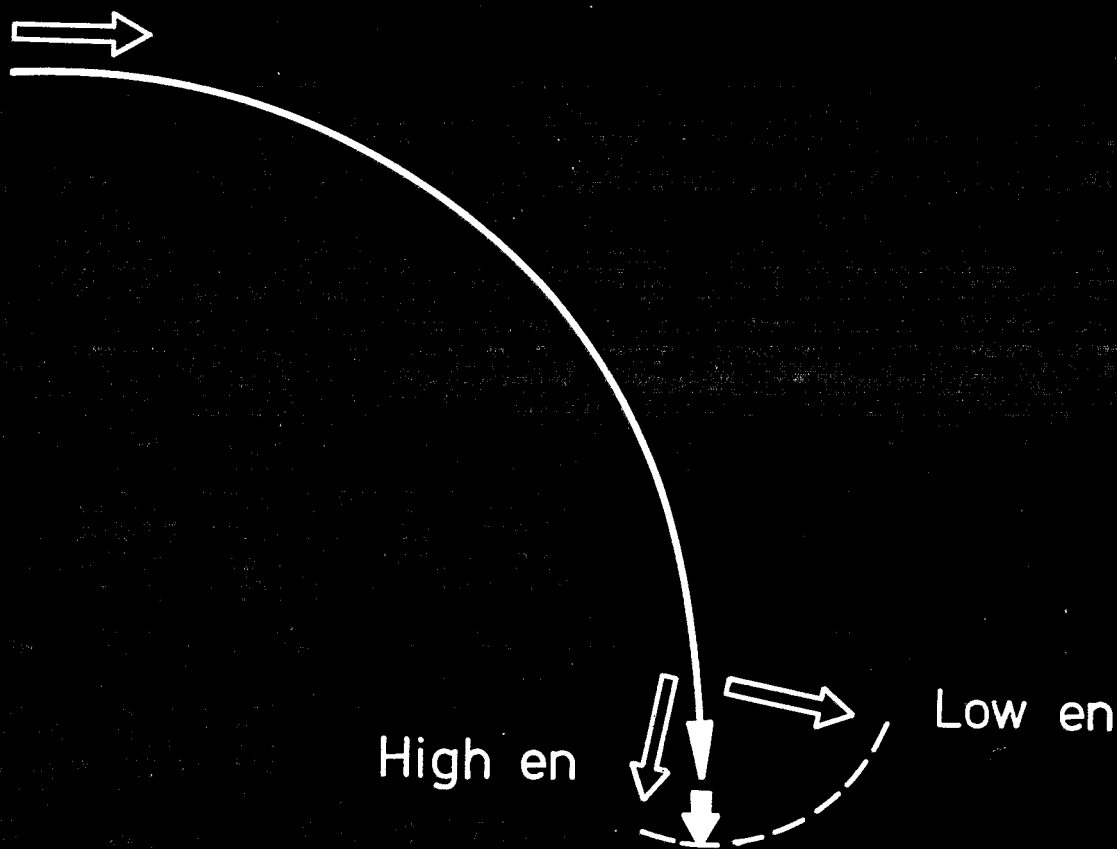
$$p_m = 3.094 \text{ GeV}/c$$

MAGIC GAMMA

- UNIFORM B
ELECTRIC FOCUSING

$$\theta = \theta_s - \theta_i - \theta_p = a_p \left(\frac{c}{mc} \right) Bt [1 + (\beta - 1/\alpha) \beta \gamma^2 (E_r/B)]$$

RADIAL ELECTRIC FIELD



$$\gamma = \sqrt{\frac{1}{\alpha} + 1} = 29.304$$

$$E_{\mu} = 3.096 \text{ GeV}$$

$$f'_a / f_a = 1 + \left(\beta - \frac{1}{\alpha \beta \gamma^2} \right) E/B$$

IMPROVEMENTS

Uniform B, electric focusing

Pion injection

40 point NMR stabilizer

Grinding magnet pole

Electric scraping

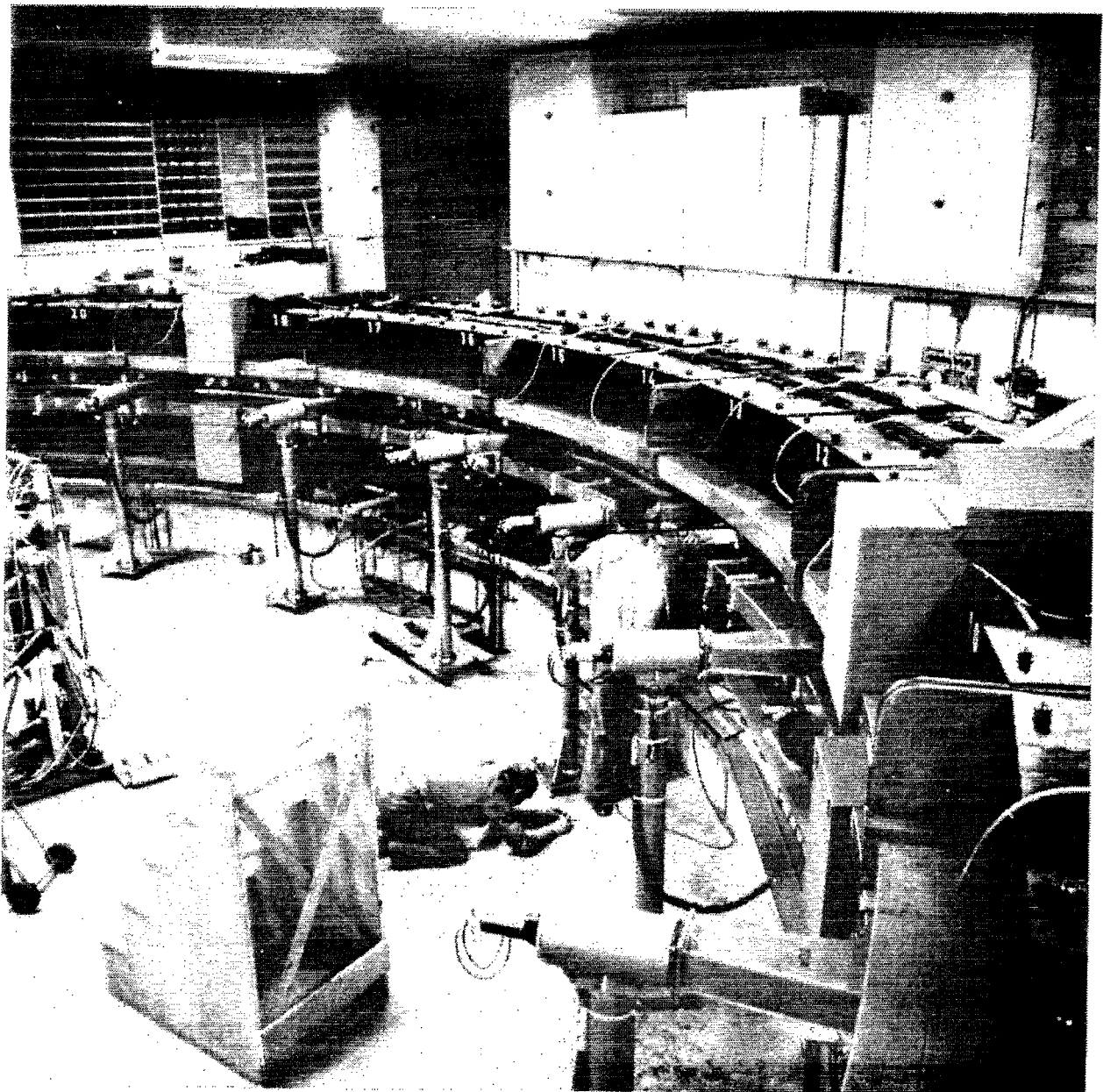
3 GeV $\tau \rightarrow 64 \mu\text{s}$

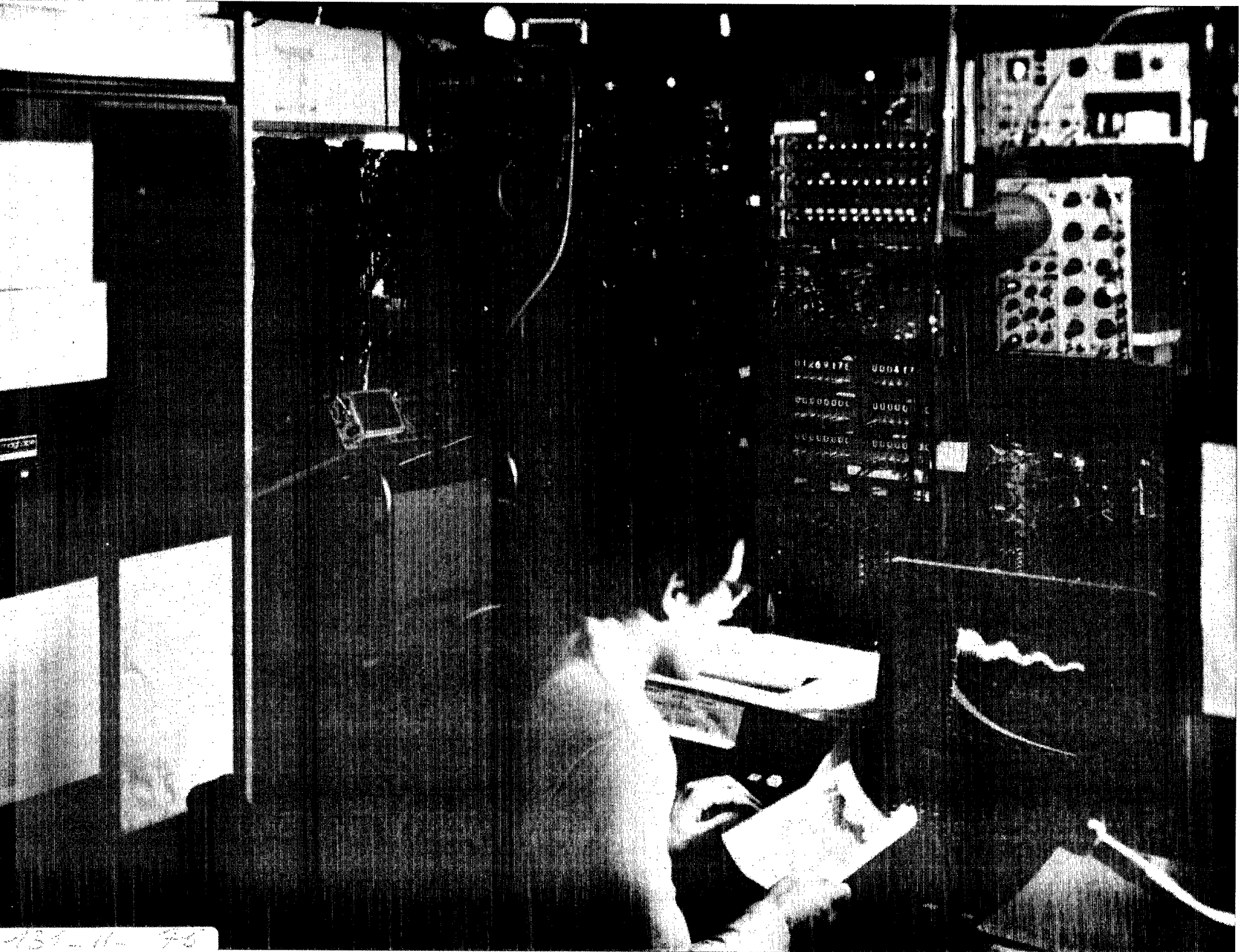
Nuclear Physics B150 (1979) 1-75
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CERN
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SCIENTIFIQUE

**FINAL REPORT ON THE CERN MUON STORAGE RING
INCLUDING THE ANOMALOUS MAGNETIC MOMENT AND
THE ELECTRIC DIPOLE MOMENT OF THE MUON,
AND A DIRECT TEST OF RELATIVISTIC TIME DILATION**

J. BAILEY ¹⁾, K. BORER ²⁾, F. COMBLEY ³⁾, H. DRUMM ⁴⁾,
C. ECK ⁴⁾, F.J.M. FARLEY ⁵⁾, J.H. FIELD ⁶⁾, W. FLEGEL ⁶⁾, P.M.
HATTERSLEY ⁷⁾, F. KRIENEN ⁶⁾, F. LANGE ⁴⁾, G. LEBÉE ⁶⁾,
E. McMILLAN ⁸⁾, G. PETRUCCI ⁶⁾, E. PICASSO ⁶⁾, O. RÚNOLFSSON ⁶⁾,
W. von RÜDEN ⁴⁾, R.W. WILLIAMS ⁹⁾ and S. WOJCICKI ¹⁰⁾
CERN-Mainz-Daresbury Collaboration

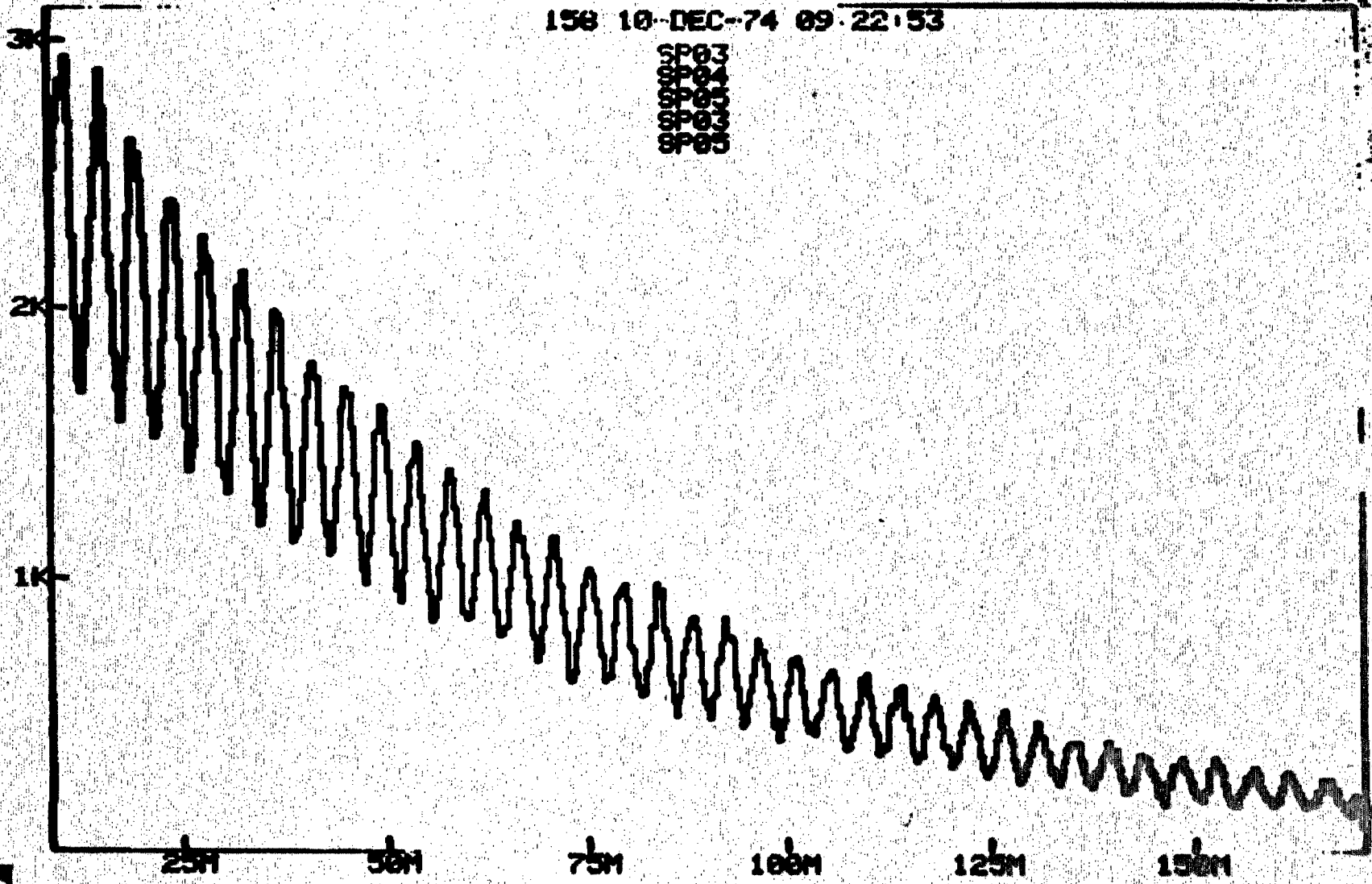




434-11-78

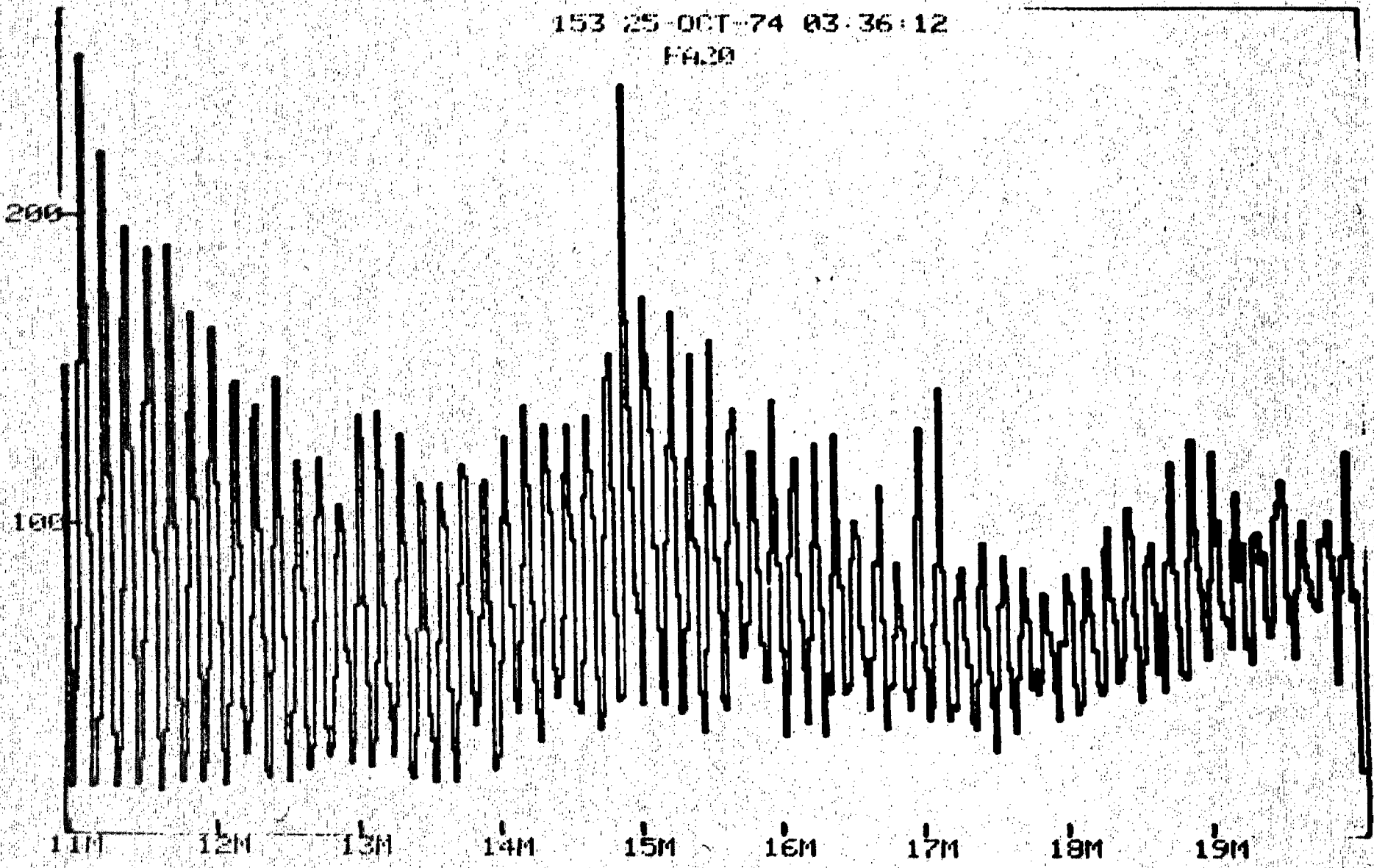
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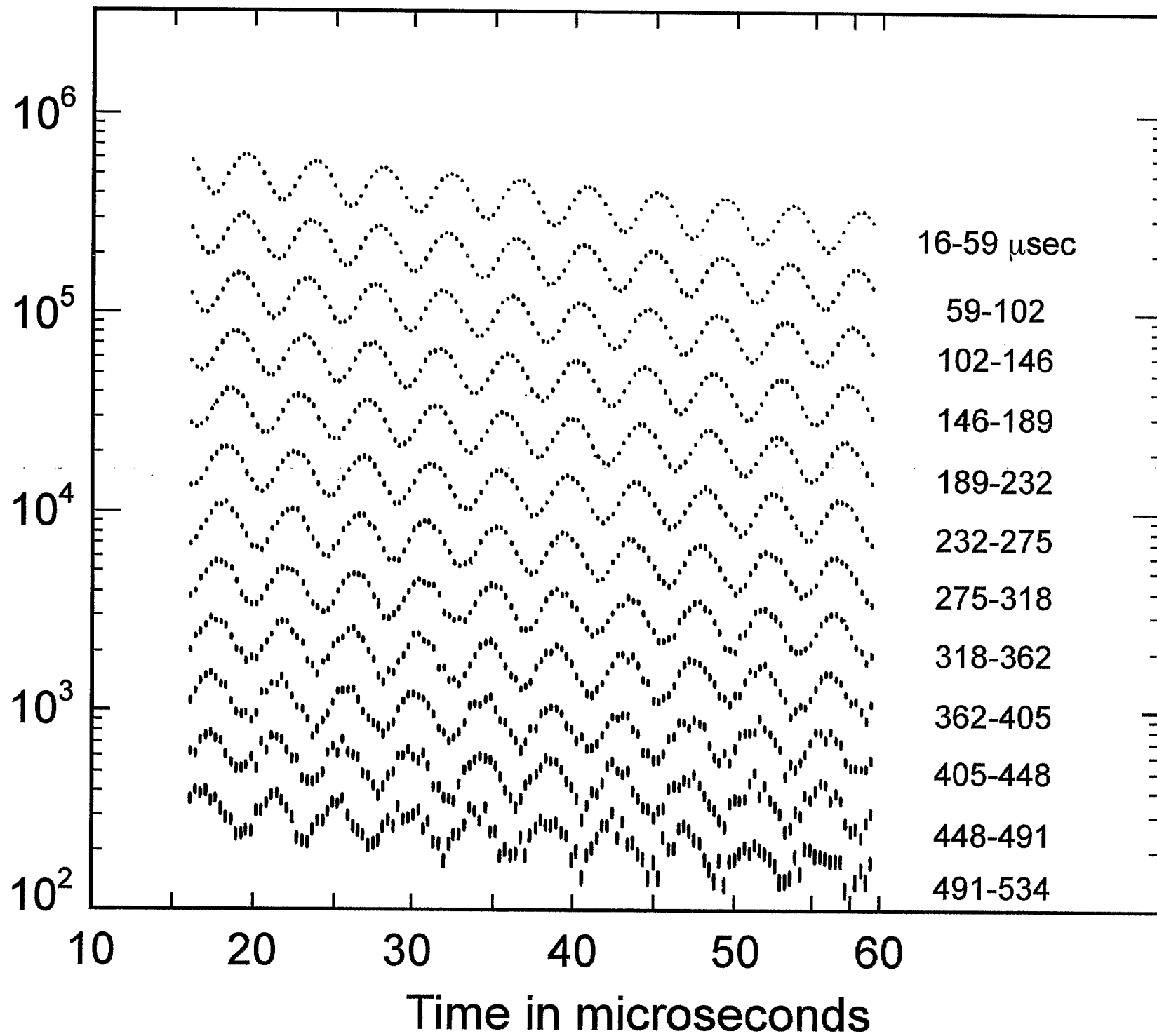
SP03
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SP09



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FA20





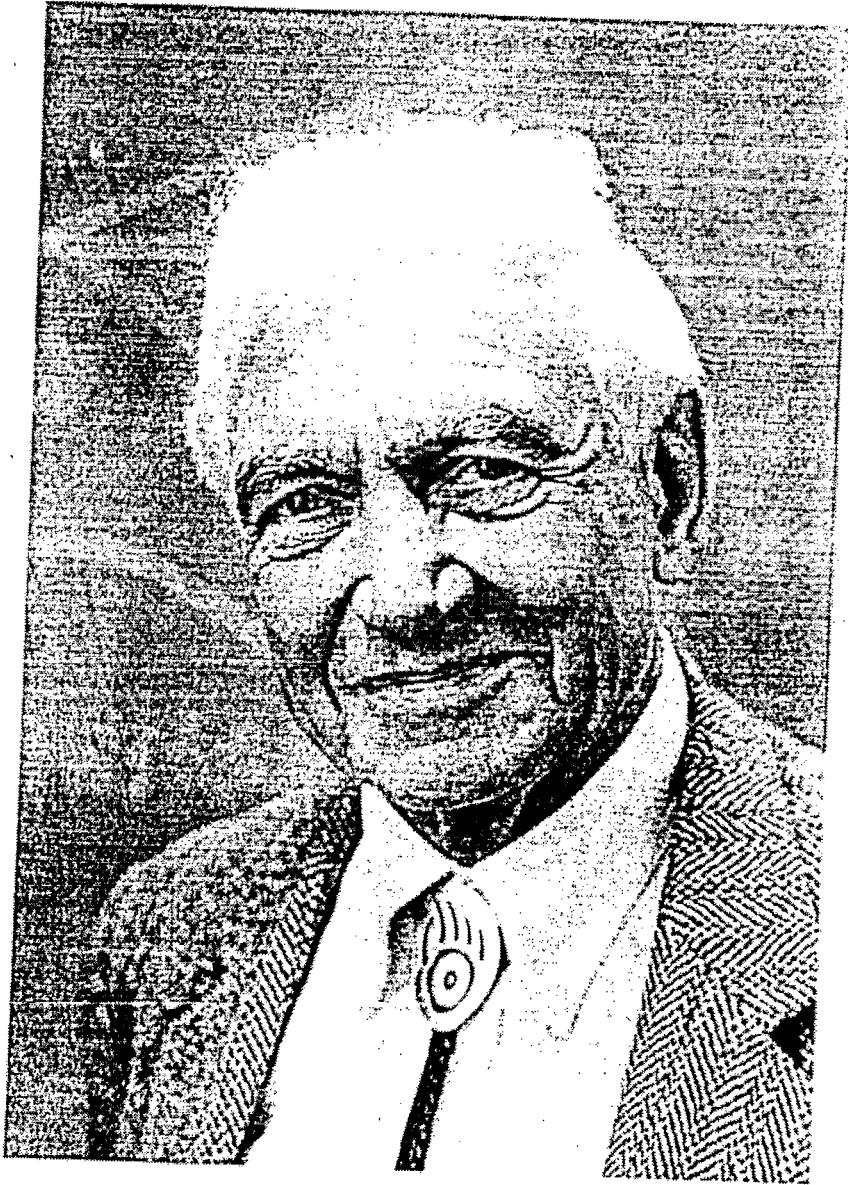




Figure 24 At Brookhaven National Laboratory, summer 1984. *Standing, from left:* Gordon Danby, John Field, Francis Farley, Emilio Picasso, and Frank Krienen; *kneeling, from left:* John Bailey, Vernon Hughes, and Fred Combley.

Precise Measurement of the Positive Muon Anomalous Magnetic Moment

H. N. Brown,² G. Bunce,² R. M. Carey,¹ P. Cushman,⁹ G. T. Danby,² P. T. Debevec,⁷ M. Deile,¹¹ H. Deng,¹¹ W. Deninger,⁷ S. K. Dhawan,¹¹ V. P. Druzhinin,³ L. Duong,⁹ E. Efstathiadis,¹ F. J. M. Farley,¹¹ G. V. Fedotovitch,³ S. Giron,⁹ F. Gray,⁷ D. Grigoriev,³ M. Grosse-Perdekamp,¹¹ A. Grossmann,⁶ M. F. Hare,¹ D. W. Hertzog,⁷ V. W. Hughes,¹¹ M. Iwasaki,¹⁰ K. Jungmann,⁶ D. Kawall,¹¹ M. Kawamura,¹⁰ B. I. Khazin,³ J. Kindem,⁹ F. Krienen,¹ I. Kronkvist,⁹ R. Larsen,² Y. Y. Lee,² I. Logashenko,^{1,3} R. McNabb,⁹ W. Meng,² J. Mi,² J. P. Miller,¹ W. M. Morse,² D. Nikas,² C. J. G. Onderwater,⁷ Y. Orlov,⁴ C. S. Özben,² J. M. Paley,¹ C. Polly,⁷ J. Pretz,¹¹ R. Prigl,² G. zu Putnitz,⁶ S. I. Redin,¹¹ O. Rind,¹ B. L. Roberts,¹ N. Ryskulov,³ S. Sedykh,⁷ Y. K. Semertzidis,² Yu. M. Shatunov,³ E. P. Sichtermann,¹¹ E. Solodov,³ M. Sossong,⁷ A. Steinmetz,¹¹ L. R. Sulak,¹ C. Timmermans,⁹ A. Trofimov,¹ D. Urner,⁷ P. von Walter,⁶ D. Warburton,² D. Winn,⁵ A. Yamamoto,⁸ and D. Zimmerman⁹

Muon ($g - 2$) Collaboration

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⁴Newman Laboratory, Cornell University, Ithaca, New York 14853

⁵Fairfield University, Fairfield, Connecticut 06430

⁶Physikalisches Institut der Universität Heidelberg, 69120 Heidelberg, Germany

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⁸KEK, High Energy Accelerator Research Organization, Tsukuba, Ibaraki 305-0801, Japan

⁹Department of Physics, University of Minnesota, Minneapolis, Minnesota 55455

¹⁰Tokyo Institute of Technology, Tokyo, Japan

¹¹Department of Physics, Yale University, New Haven, Connecticut 06520

(Received 8 February 2001)

See a more complete

review knowledge of magnetic field

Review article:

Farley & Picasso in QED (1990)

Improvements introduced in BNL 821

Larger magnet aperture
Superconducting coils
Pole face windings

Field more uniform
Field more stable
To trim field

Trolley with NMR probes in vacuum
Pulsed NMR good to 0.1 ppm

Circular magnet aperture for muons
Trace back to find mu position

 less sensitive to multipoles

Superconducting inflector

Muon injection
Muon kicker

More stored mu, less flash

Improved electric quadrupoles

Avoid resonances, no losses

Scalloped vacuum chamber

Higher asymmetry

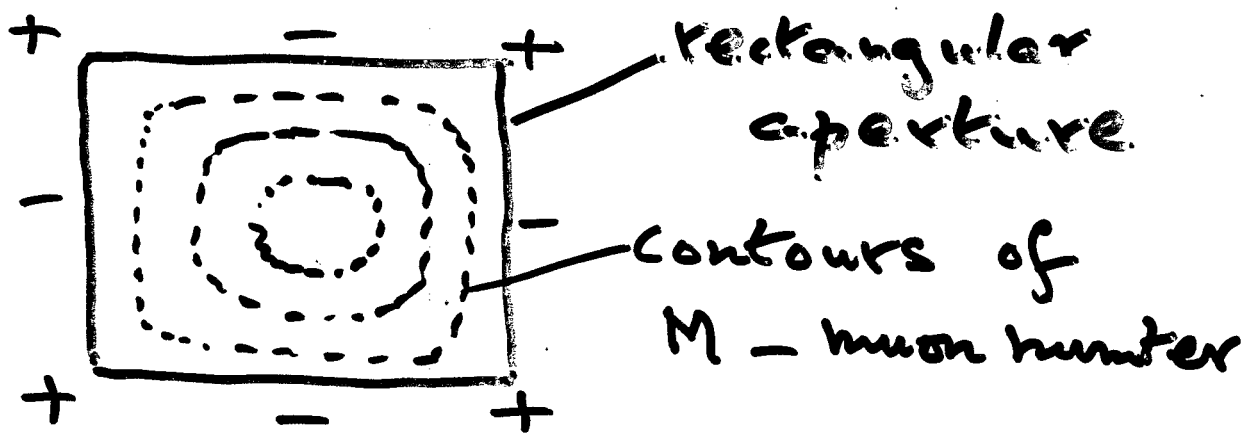
Increase in primary proton intensity

Many counter channels

Reduce overlap at high rates

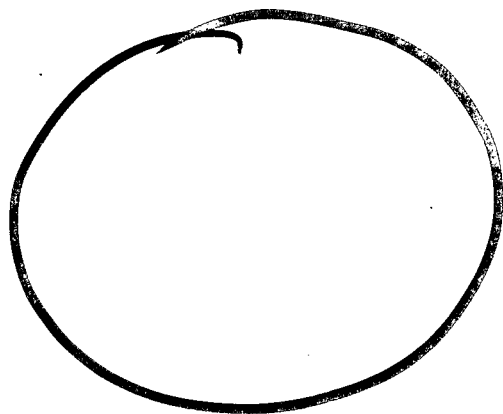
12 separate fills per AGS cycle

" " " "



I_4 is large - $r^4 \cos 4\theta$
 \therefore 4th multipole of field must be small

I_8, I_{12} also large !!



Circular aperture

Higher moments all small
 Multipoles of field do not matter - \underline{B} stays same

**Magnetic field averaged in azimuth,
expanded as multipoles**

$$B = \sum c_n r^n \cos(n\theta)$$

Muon distribution $M(r, \theta)$

$$\bar{B} = \sum_n \int M(r, \theta) c_n r^n \cos(n\theta) r dr d\theta$$

$$= \sum_n c_n \int M(r, \theta) r^n \cos(n\theta) r dr d\theta$$

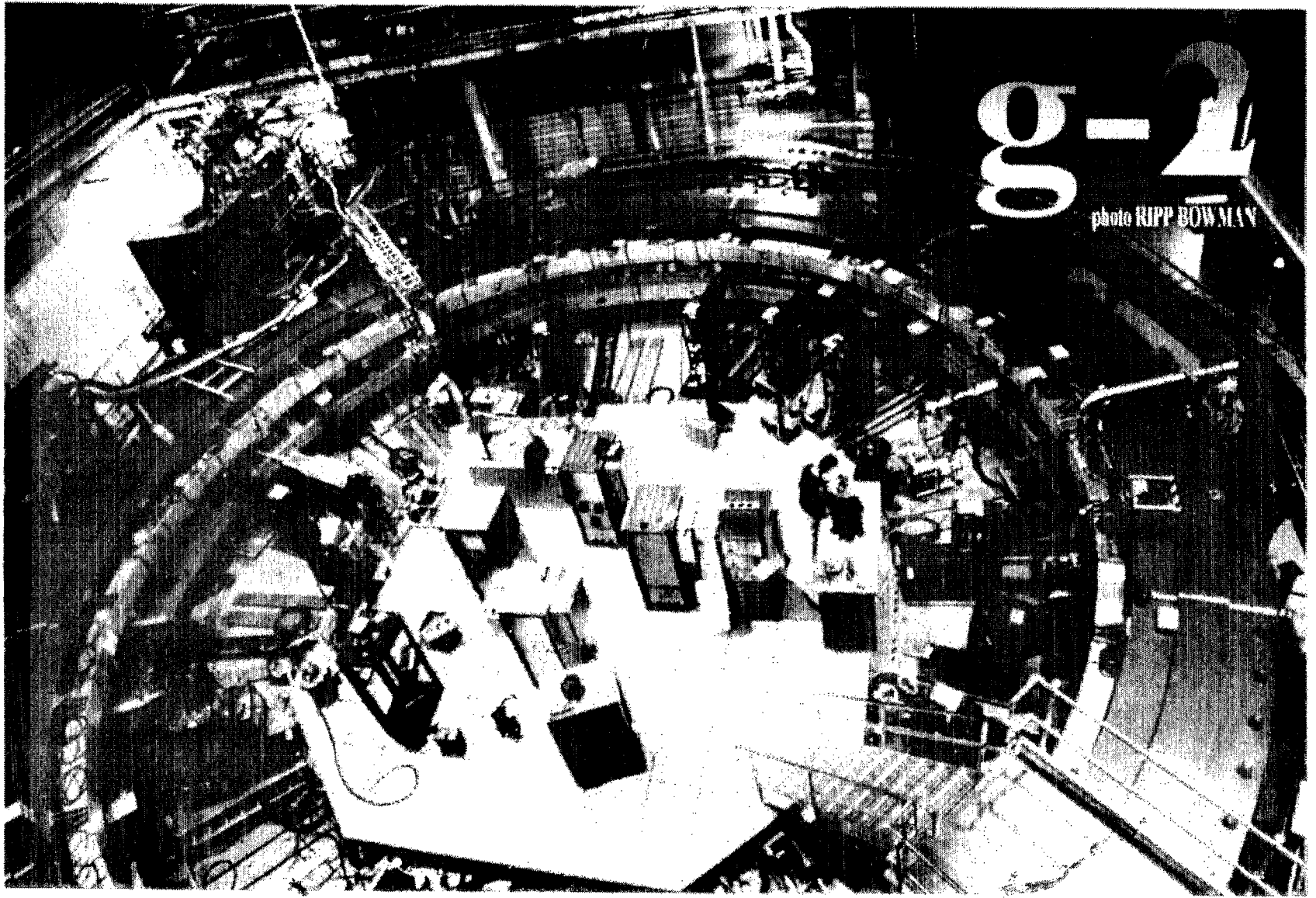
$$= \sum_n c_n I_n$$

$$I_n = \int M(r, \theta) r^n \cos(n\theta) r dr d\theta$$

I_n is n th moment of the muon distribution

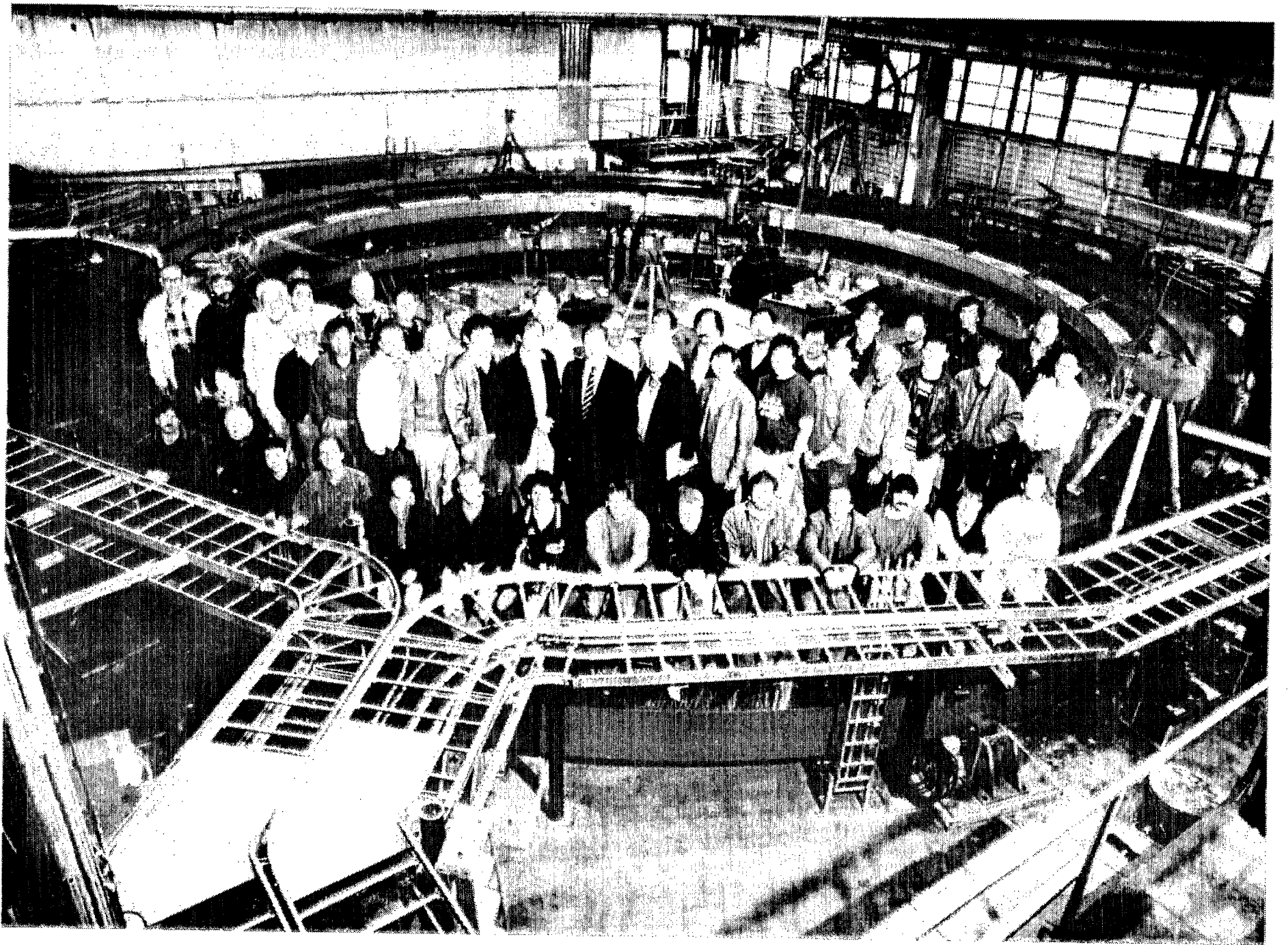
n th moment only couples to multipole of same n

If I_n is small, higher multipoles do not change \bar{B}

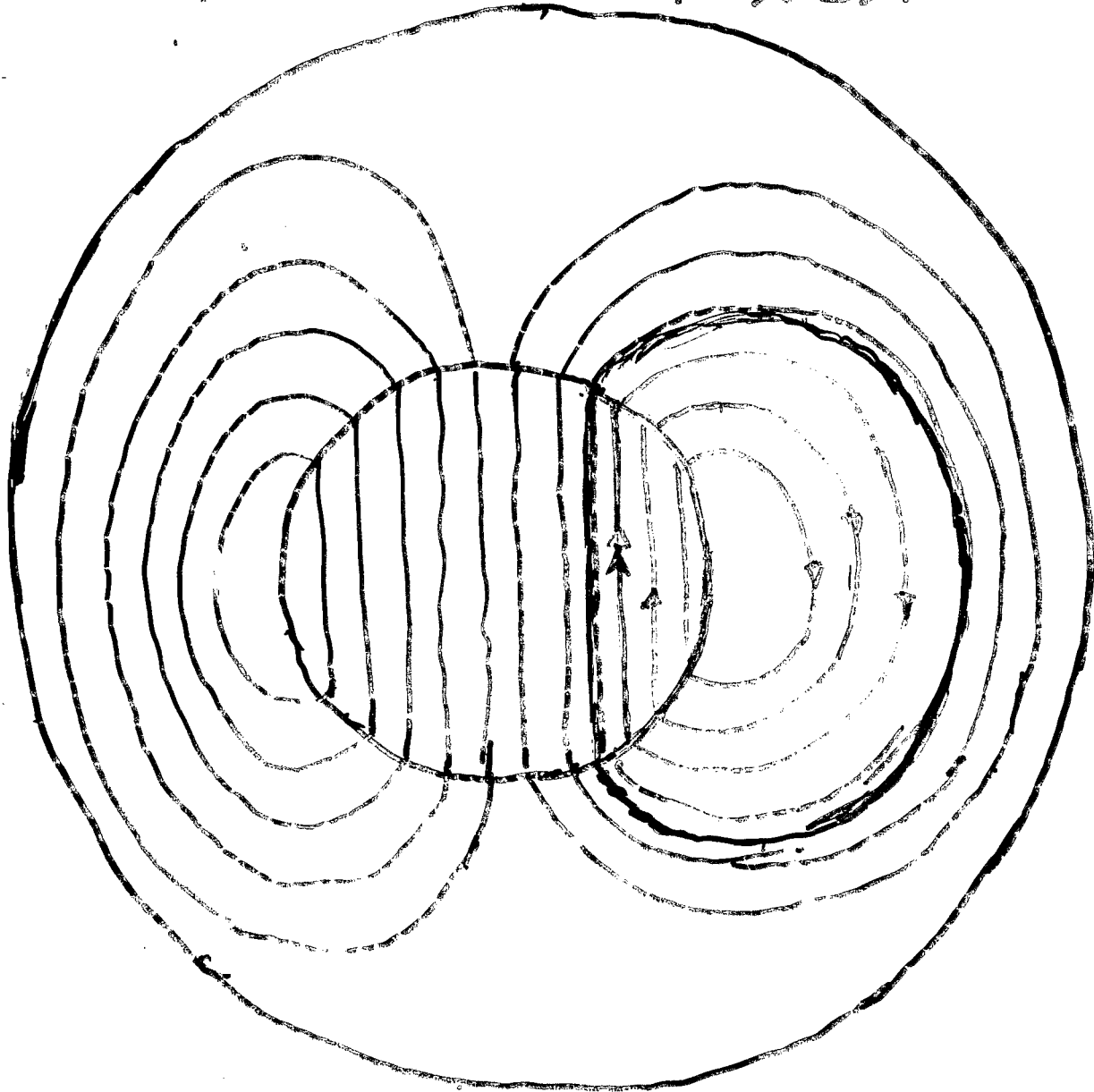


g-2

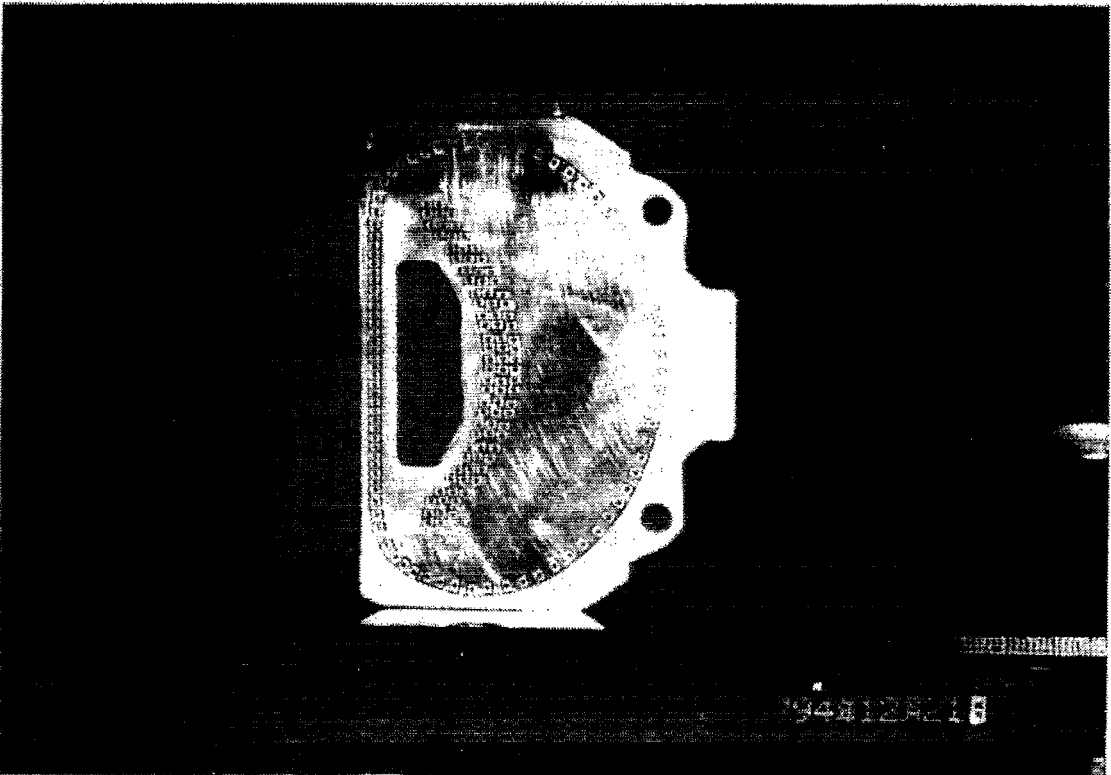
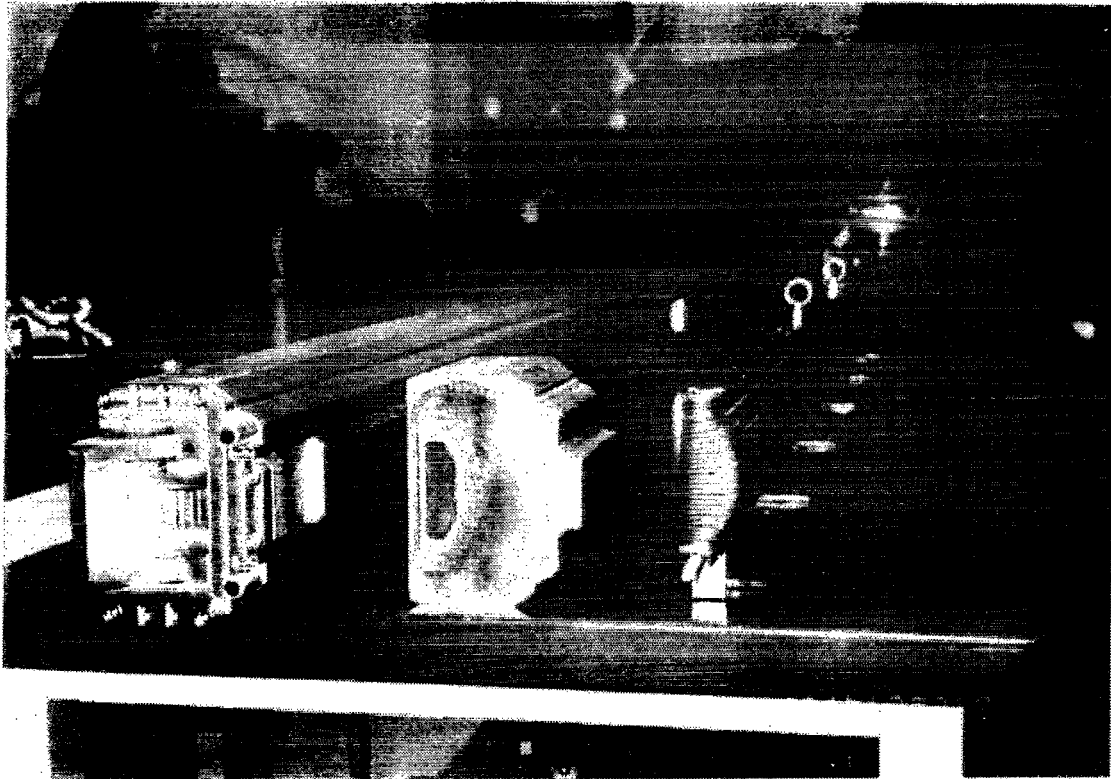
photo RIPP BOWMAN



FRANK KRIEKEN

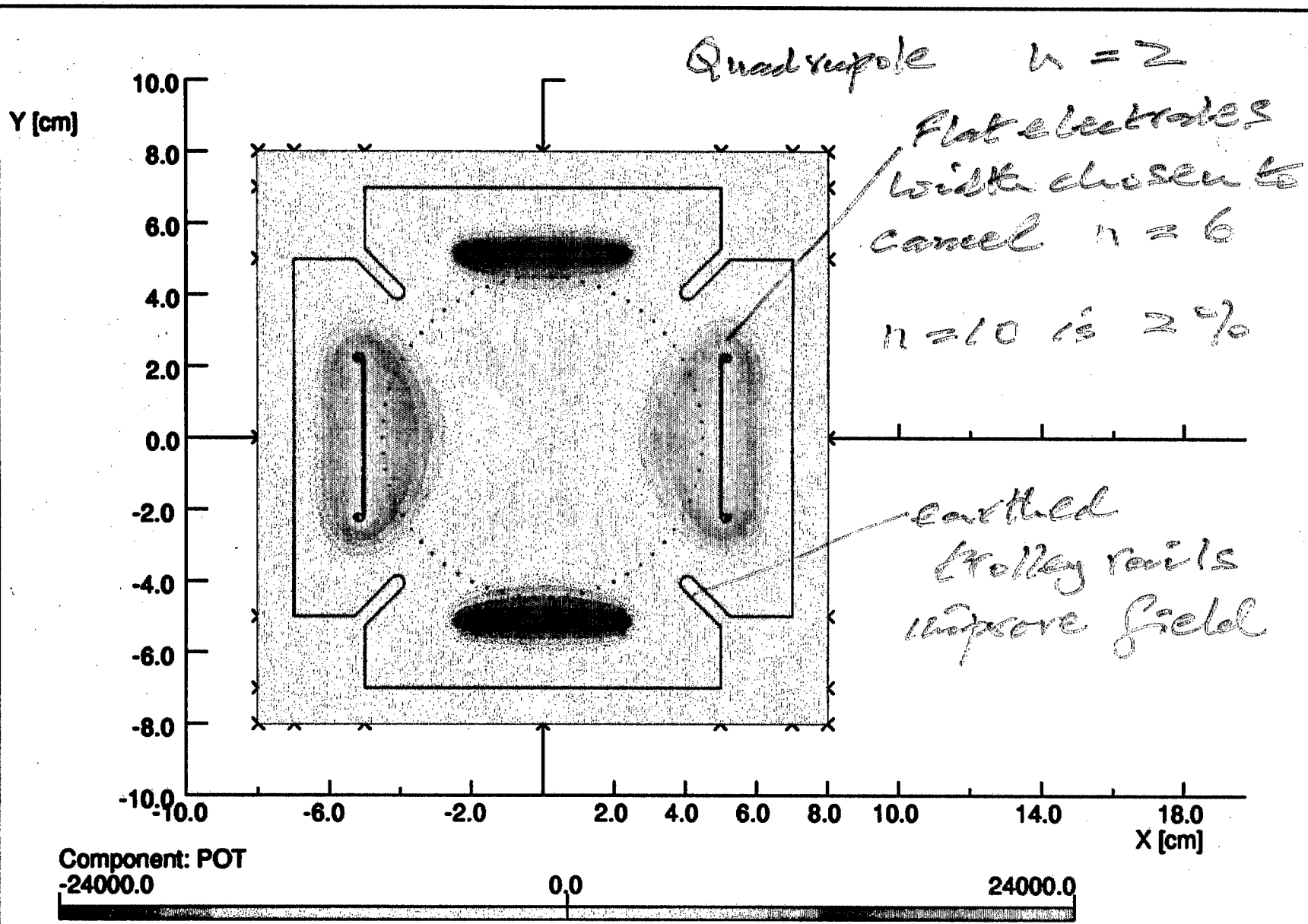


SUPERCONDUCTING INFLECTOR



SEMERTZIDIS

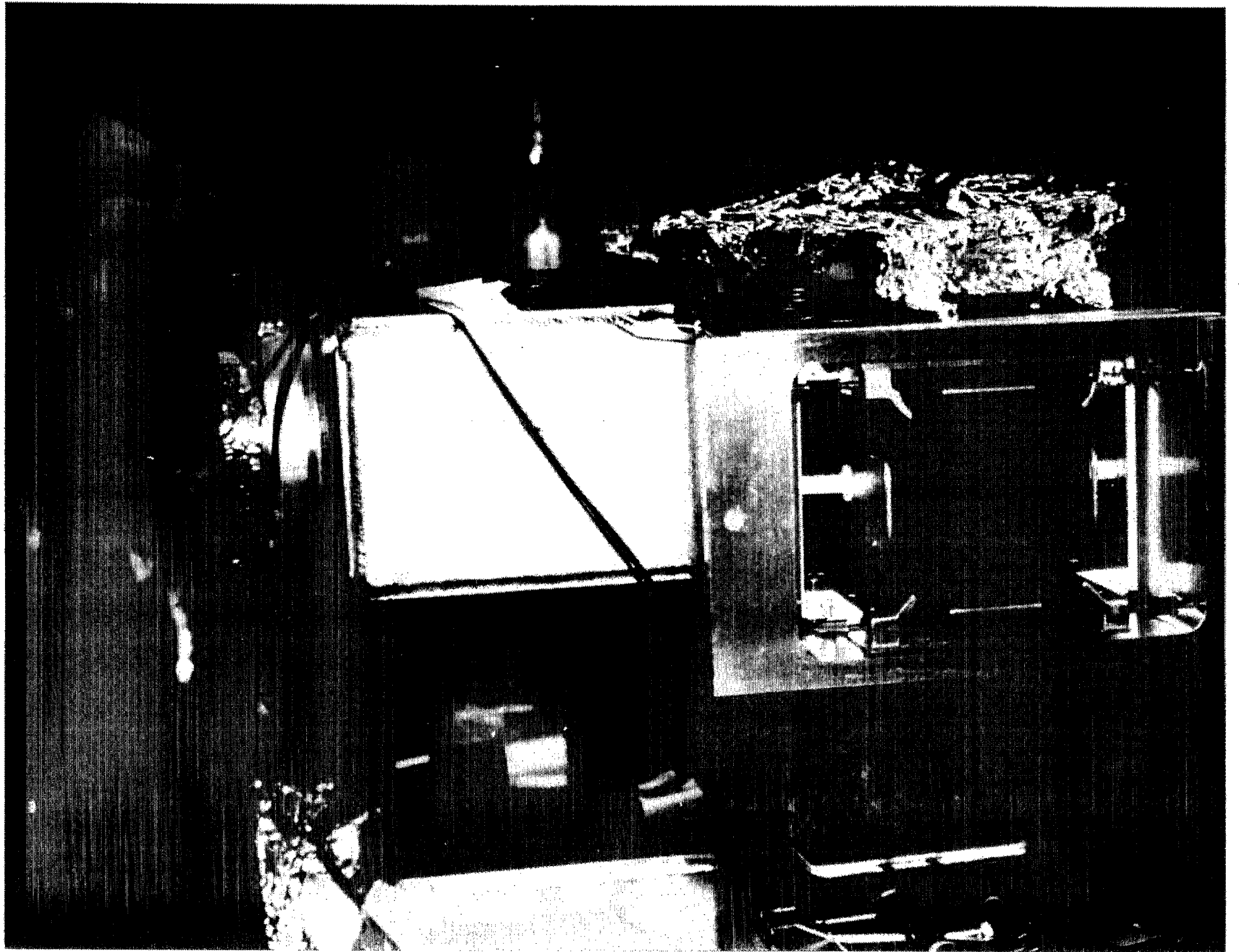
$$V = r^n \cos n\theta$$



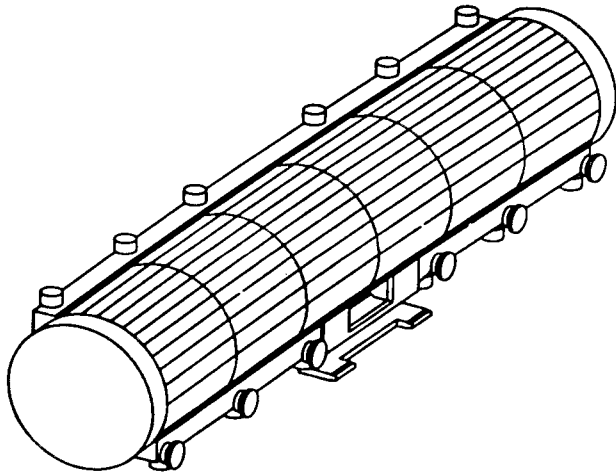
UNITS	
Length	: c
Flux density	: C
Field strength	: V
Potential	: V
Conductivity	: S
Source density	: C
Power	: W
Force	: N
Energy	: J
Mass	: kg

PROBLEM D	
quad_no	offsets
Linear elements	
XY symmetry	
Scalar potential	
Electric fields	
Static solution	
Scale factor	= 1.0
6084 elements	
3155 nodes	
52 regions	

25/Feb/1999 18:20:

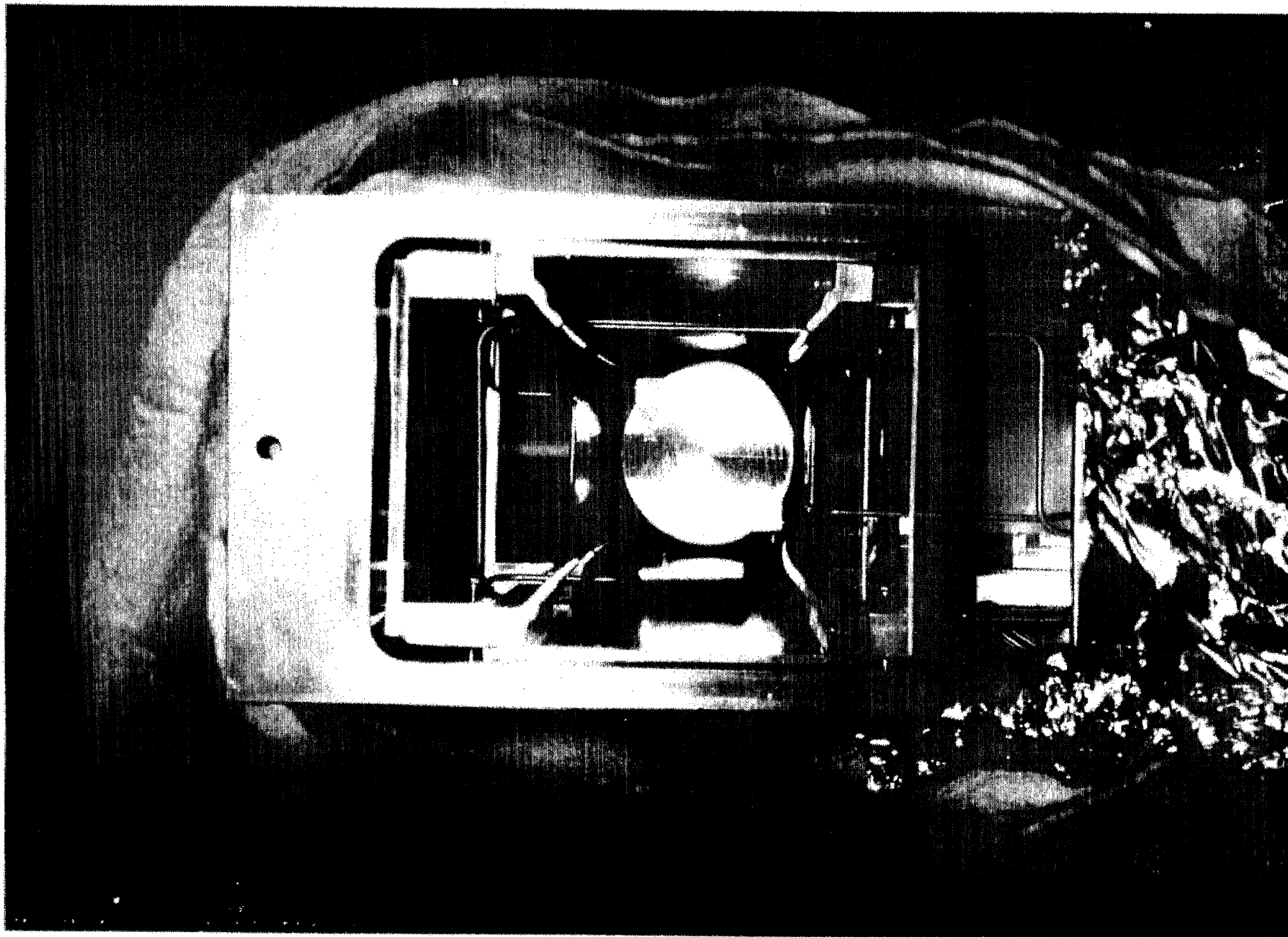


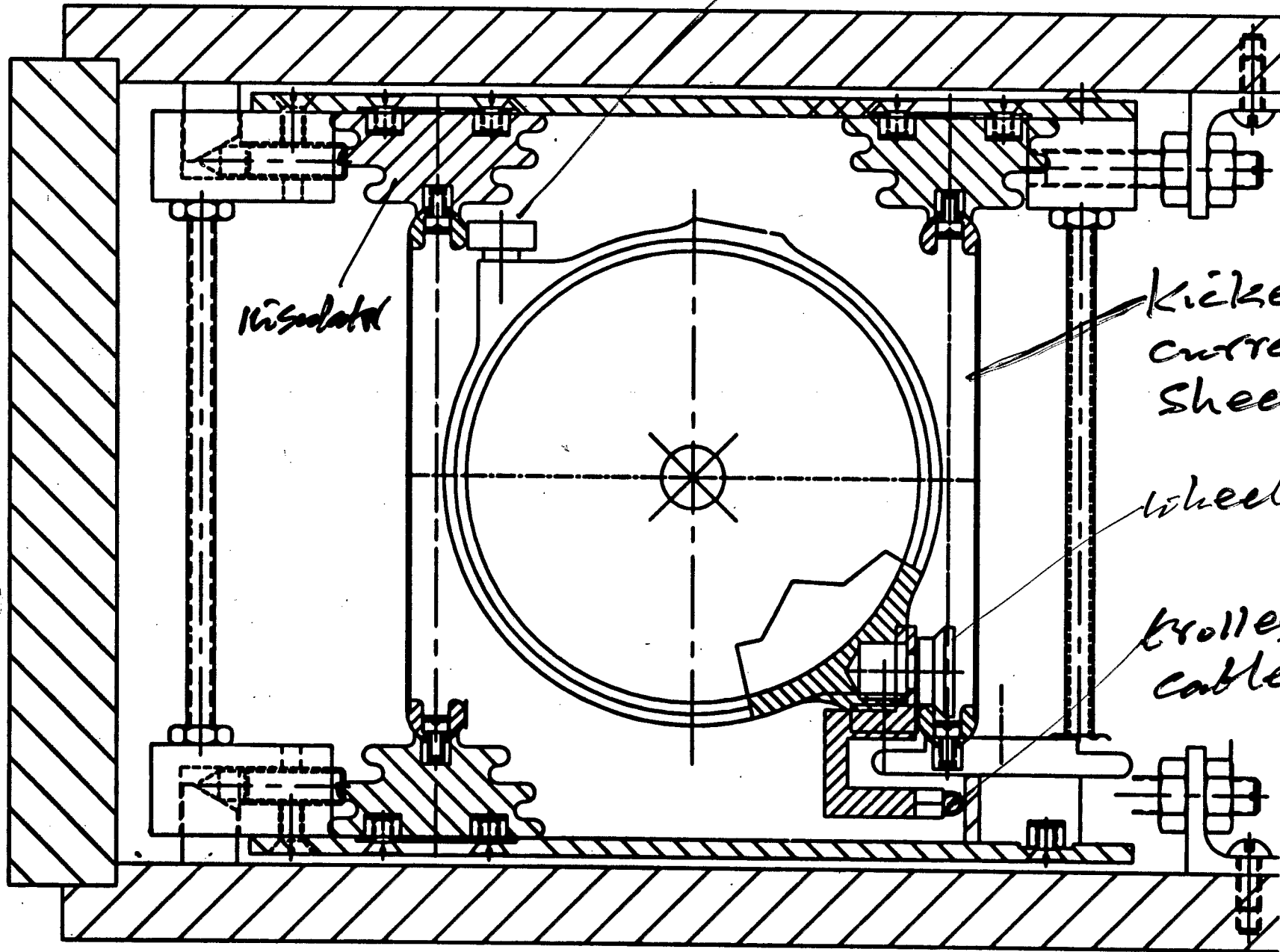
Alex Grossmann } Heidelberg
Klaus Jungmann }
Ralf Prigl



Runs on rails
Pulled by cable
Goes through
Quadrupoles
Kicker
Pick up electrodes
A tour de force!
17 NMR probes
Computer
reference f via cable







wheel



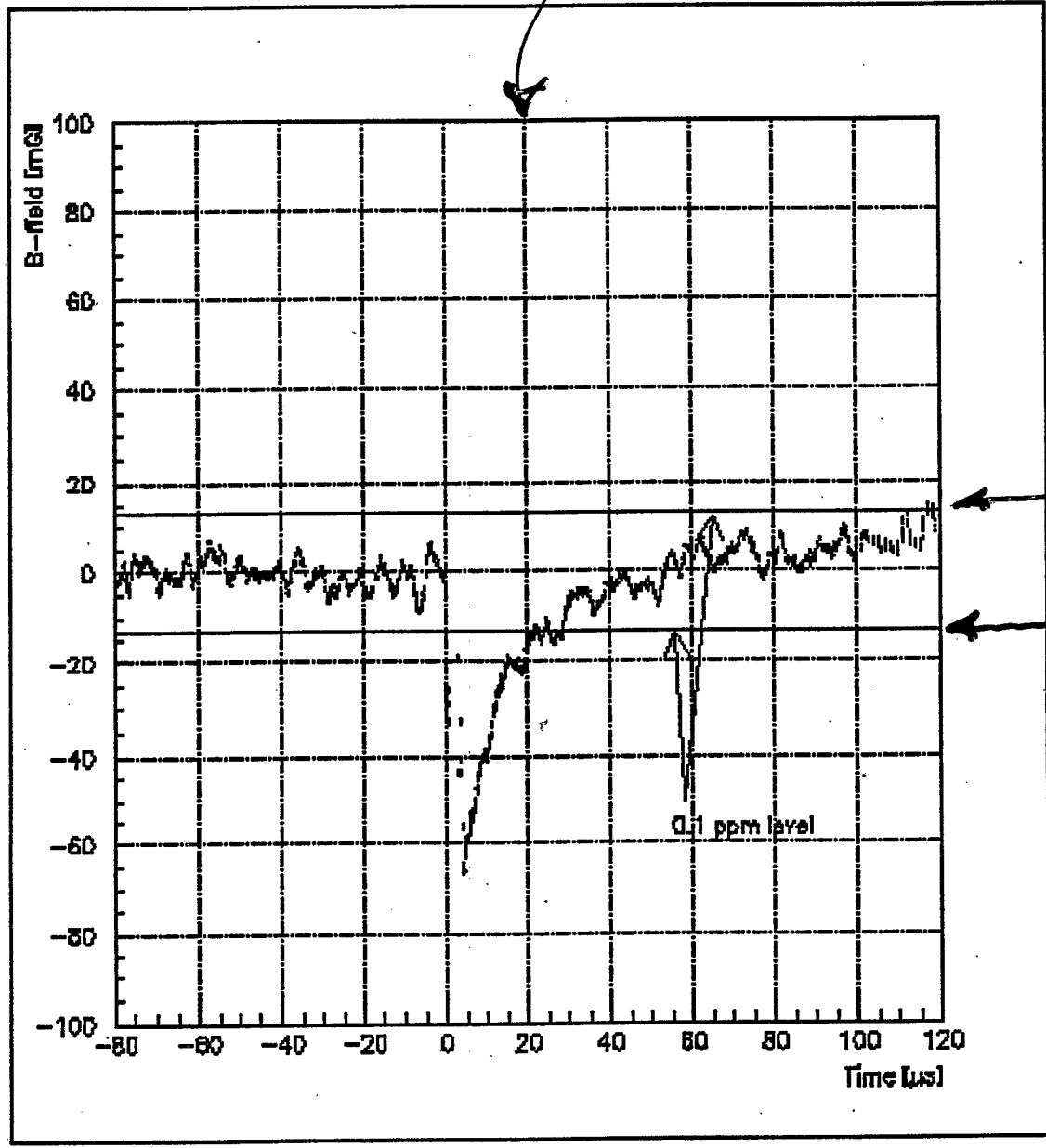
insulator

kicker
current
sheet

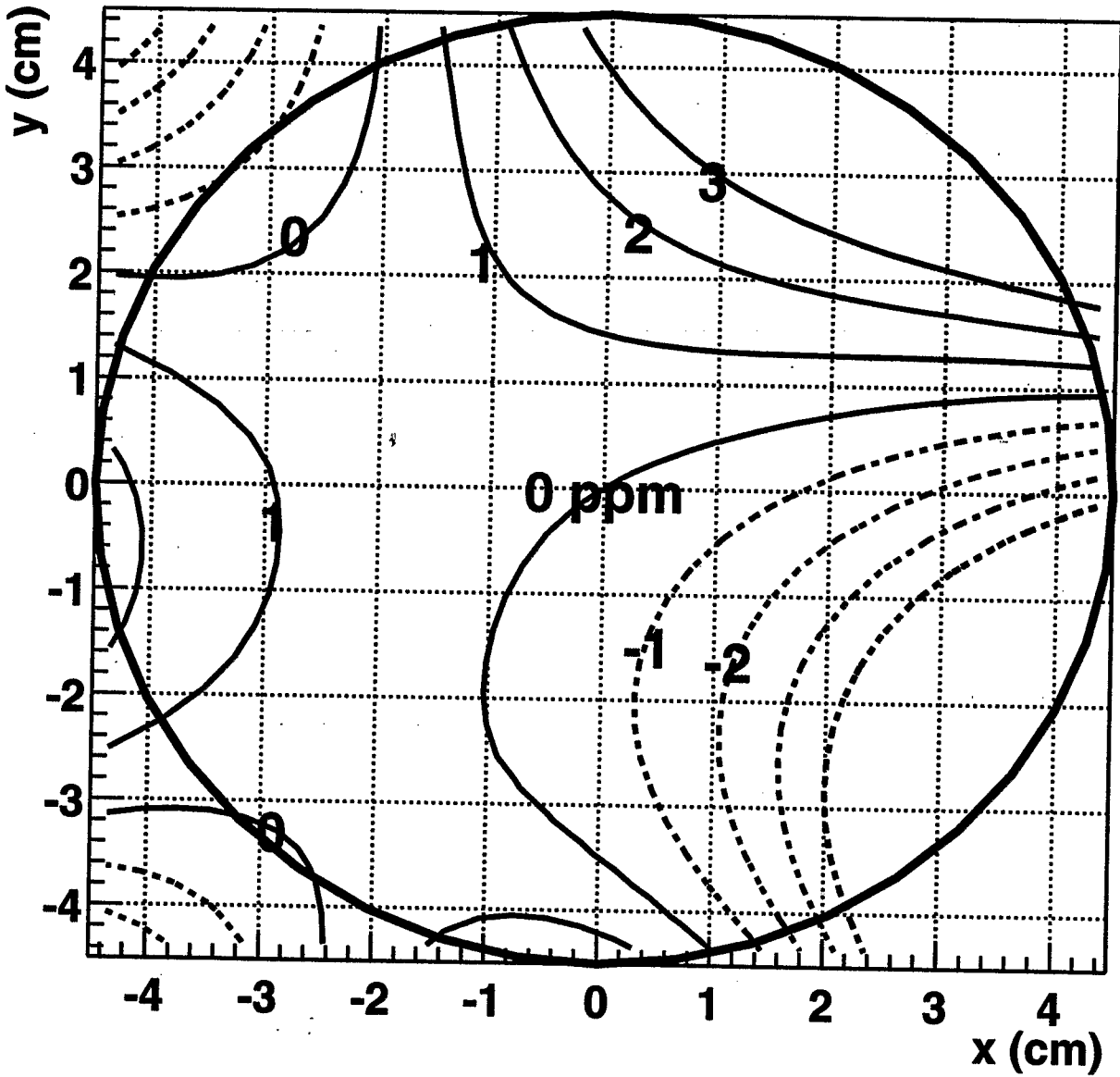
wheel

trolley
cable

↑
Residual field negligible
after 20 μ s



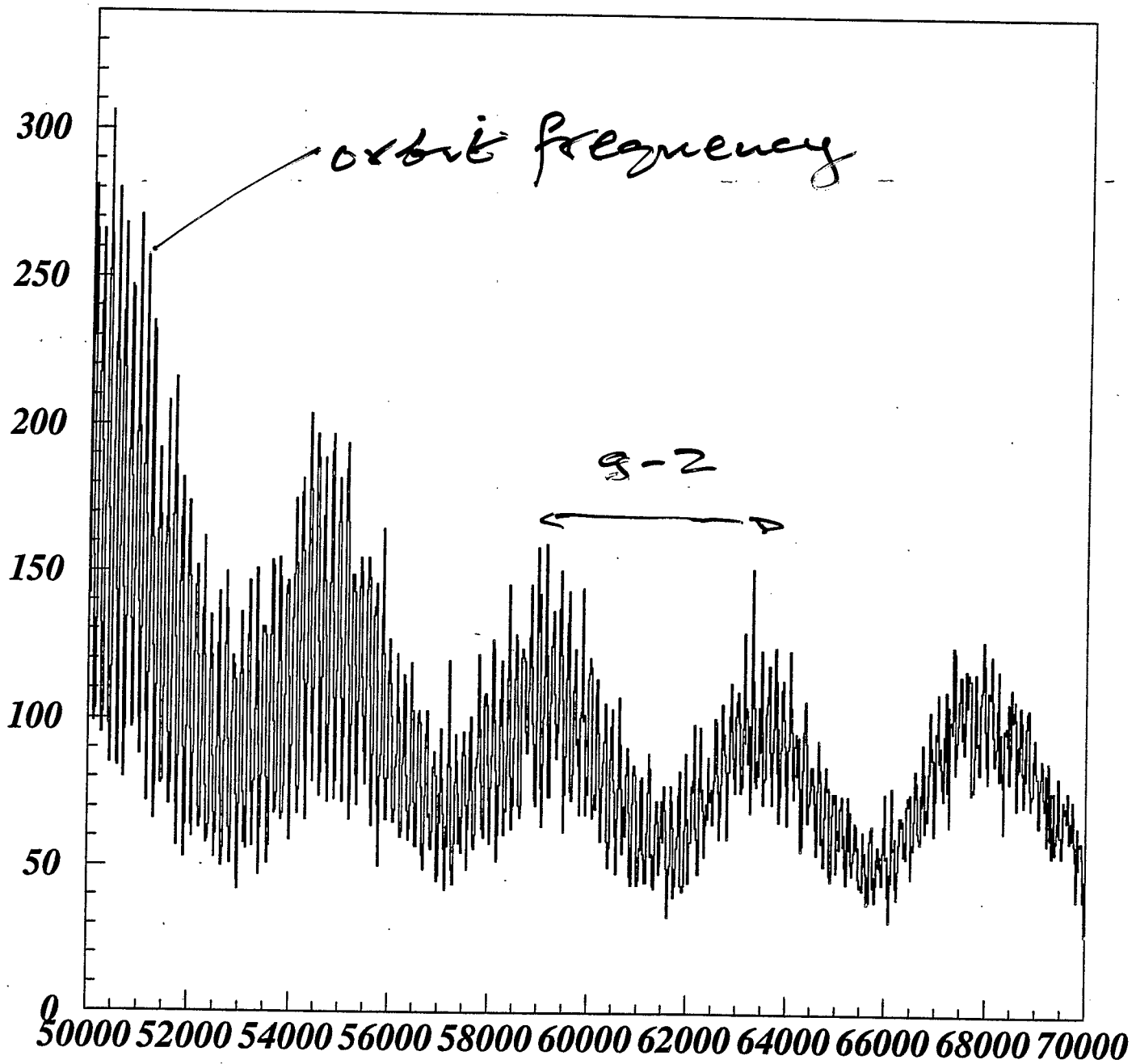
Multipole expansion of B field



B=1.451229 T (804.6 ppm)

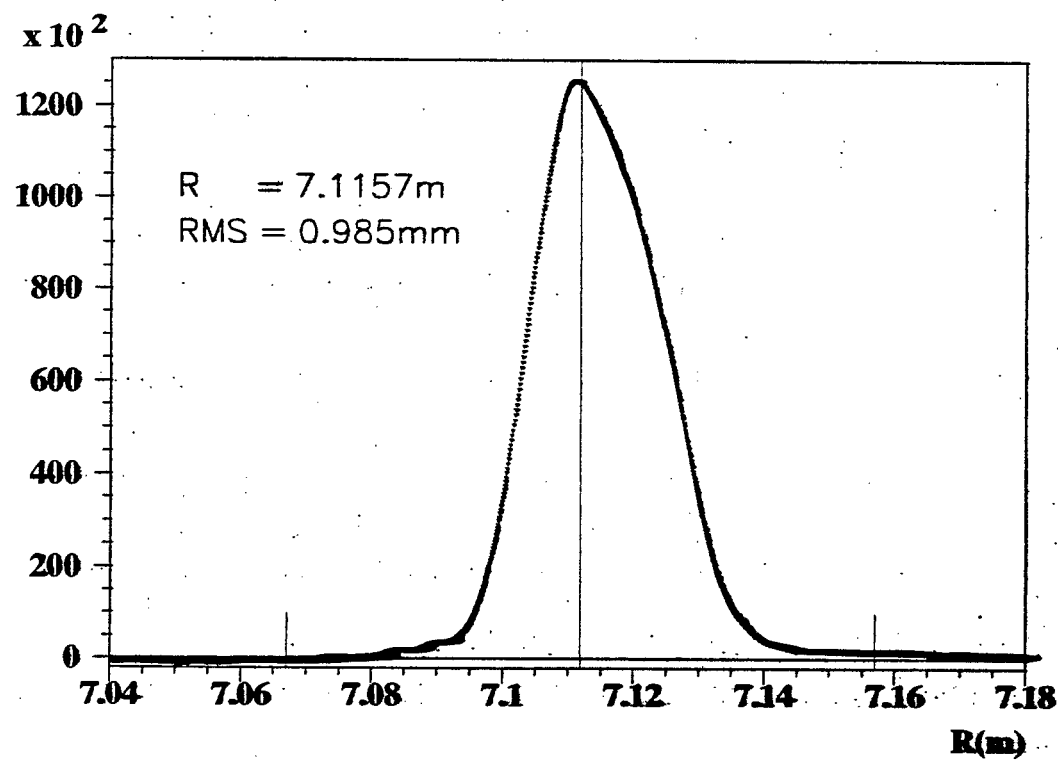
Multipoles (ppm)

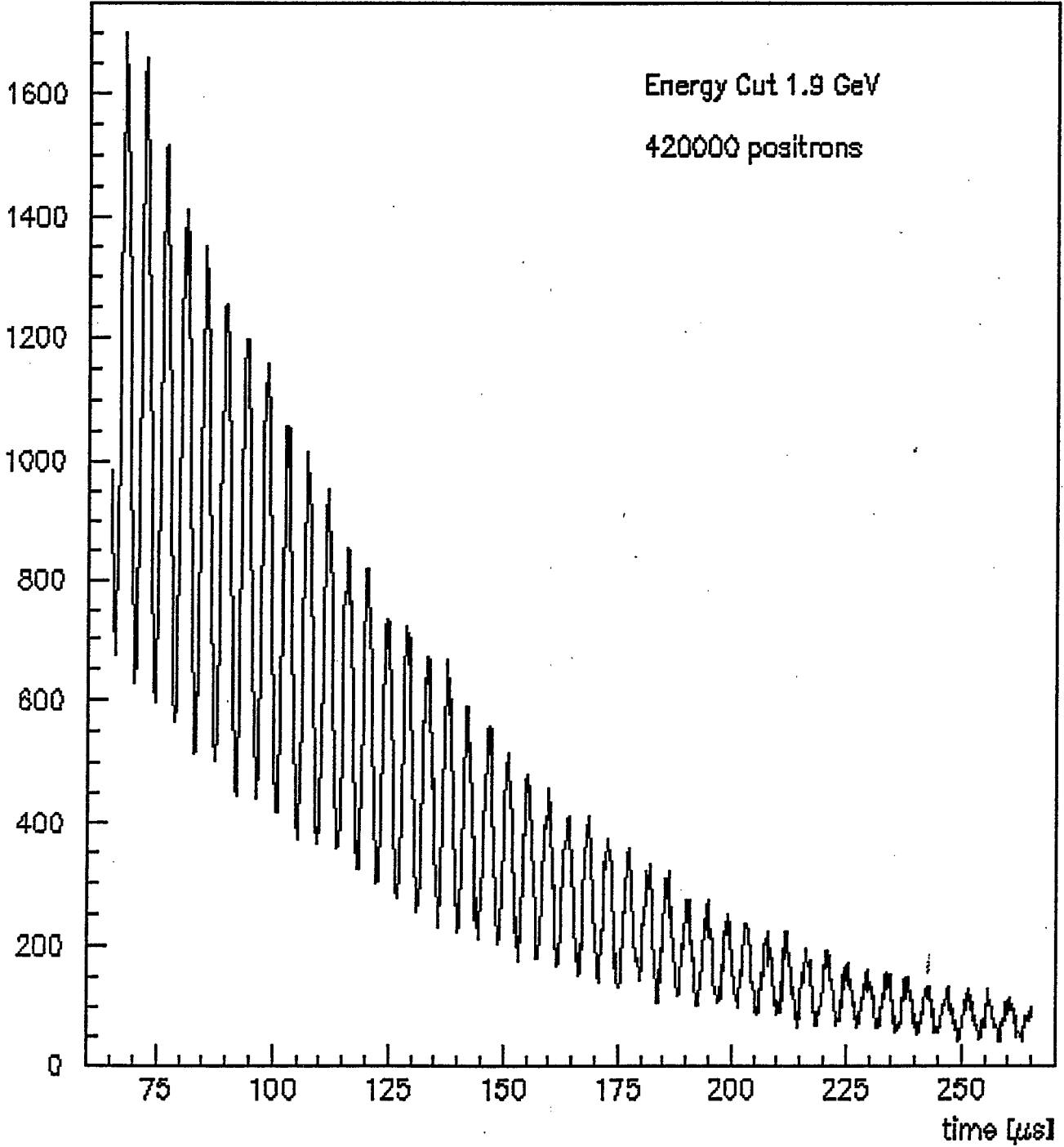
	normal	skew
Quad	-1.80	2.73
Sext	-1.61	3.70
Octu	-1.29	1.76
Decu	0.91	0.72



FOURIER TRANSFORM OF ORBIT FREQUENCY

N_{μ} vs Radius





Small effects in the counting data

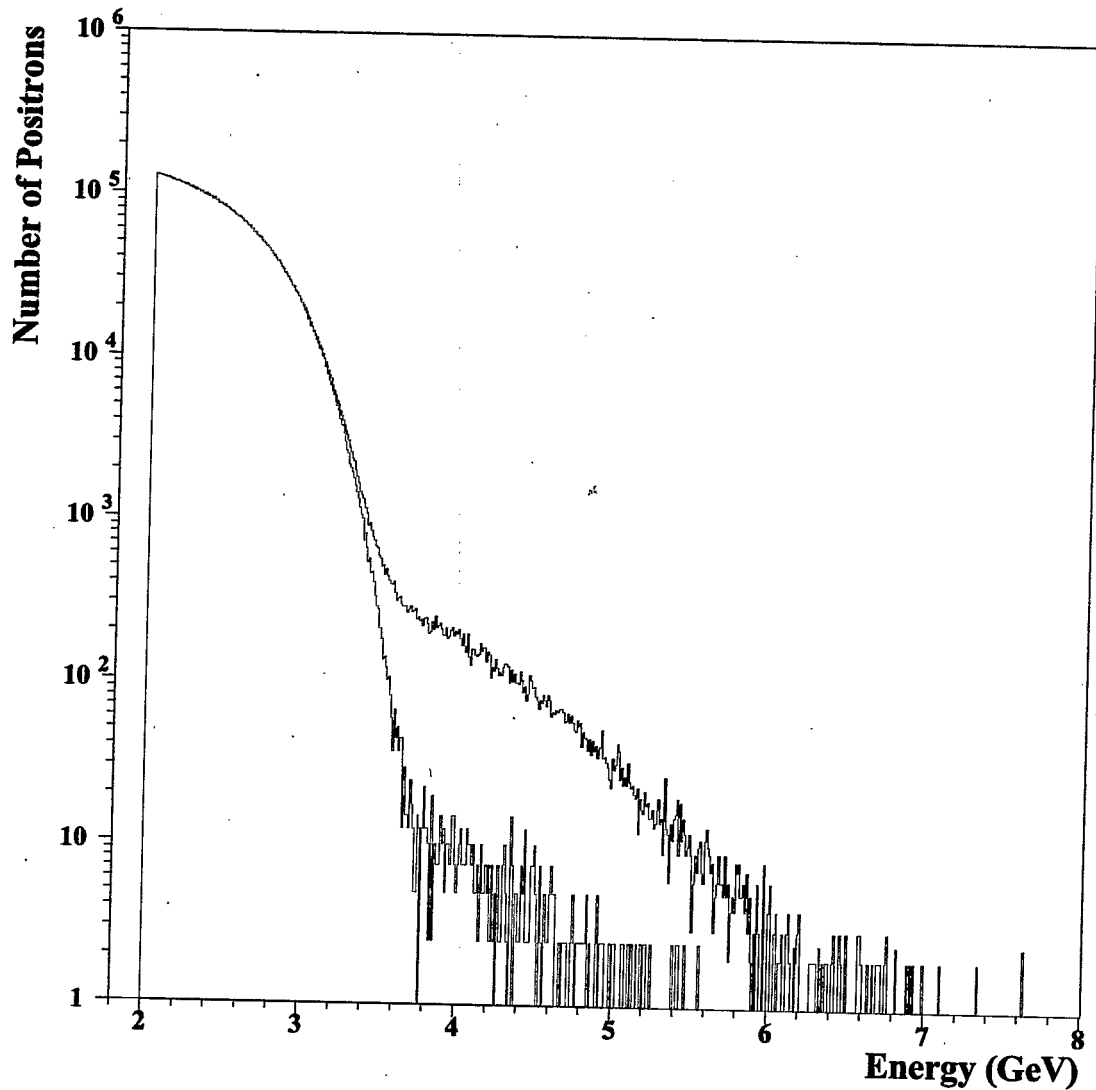
Signal overlap

Extra counts at early times

Gain change with time

Horizontal oscillations

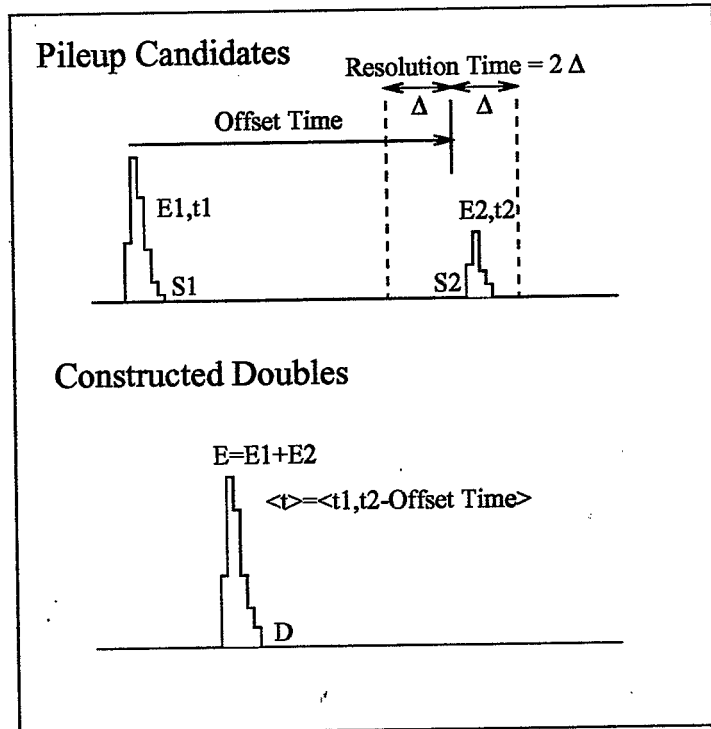
SIGNAL OVERLAP



— early times

— late times

overlap increased by g-2 modulation
& bunch structure

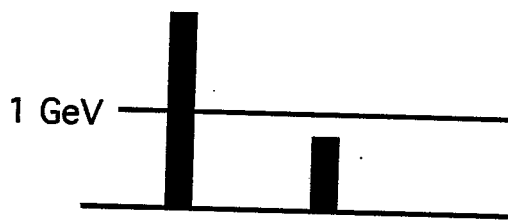
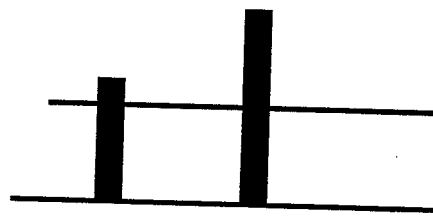
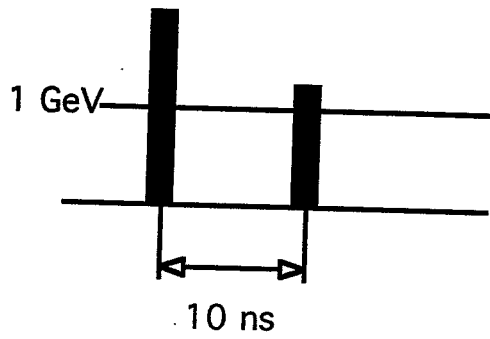


overlap adds

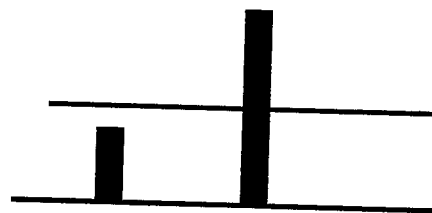
D

kills

S_1 & S_2



count this twice



not seen

Pseudo overlaps - with 1 GeV threshold

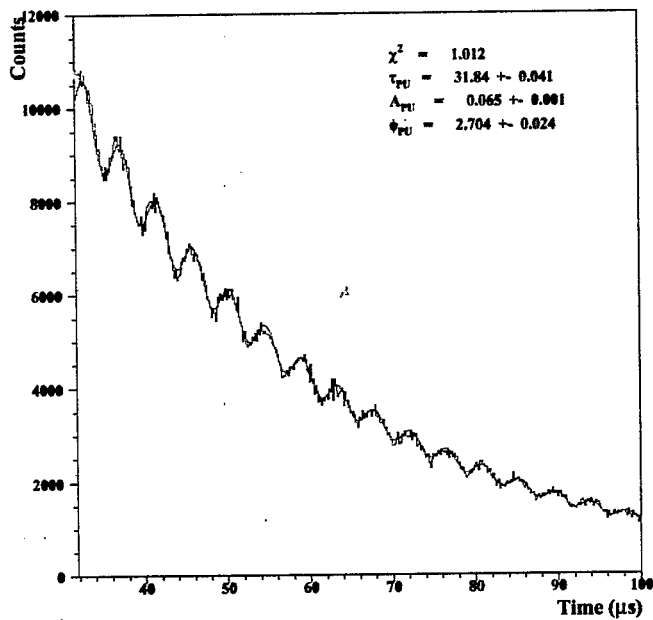
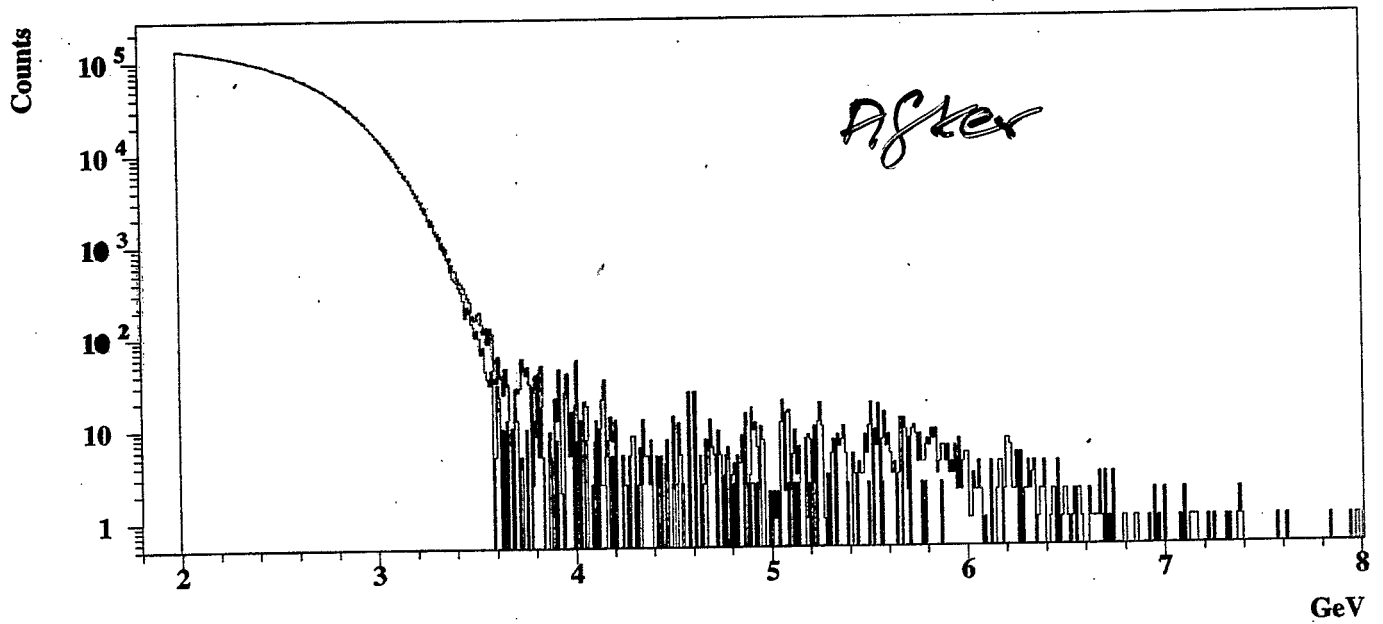
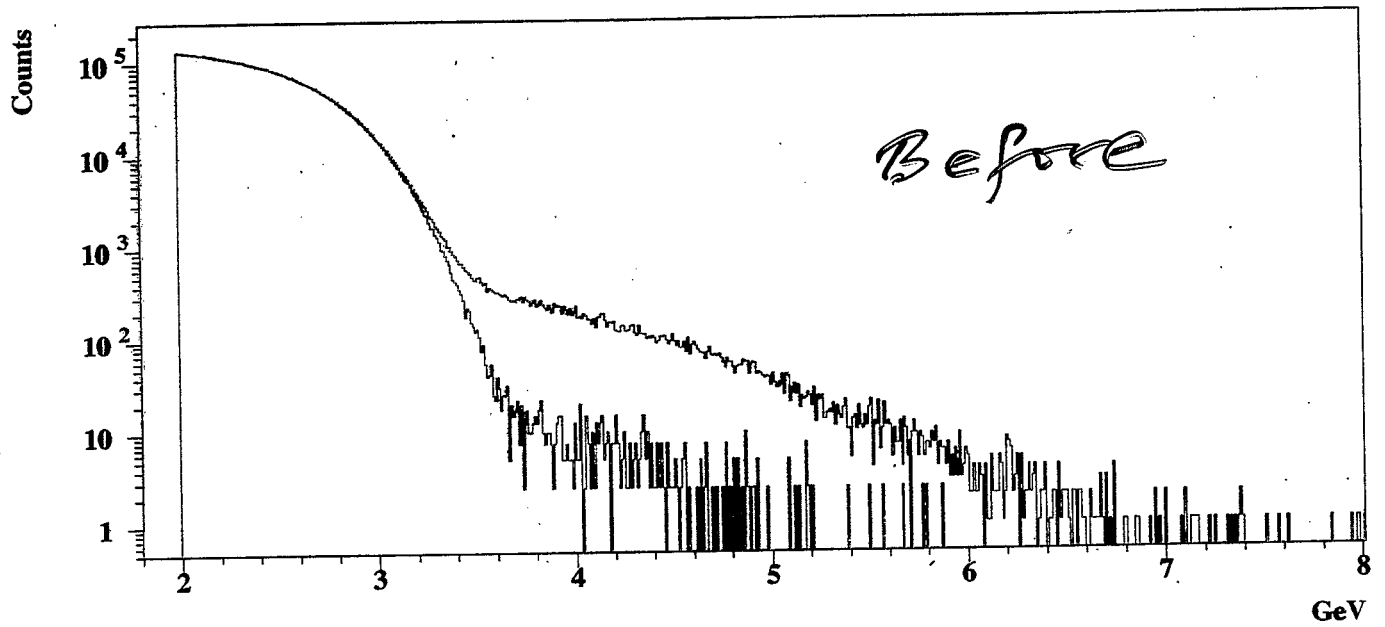


Figure 20 : Fit to pileup ($D - S_1 - S_2$).

$$\tau = 32 \mu s$$

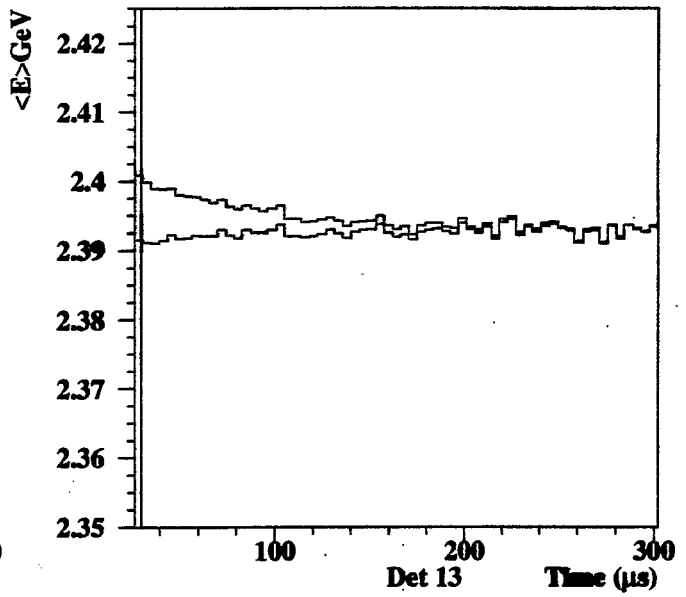
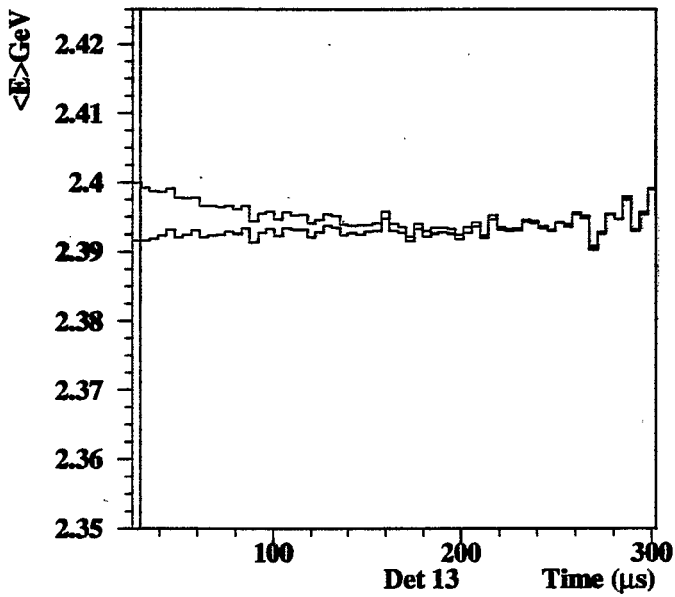
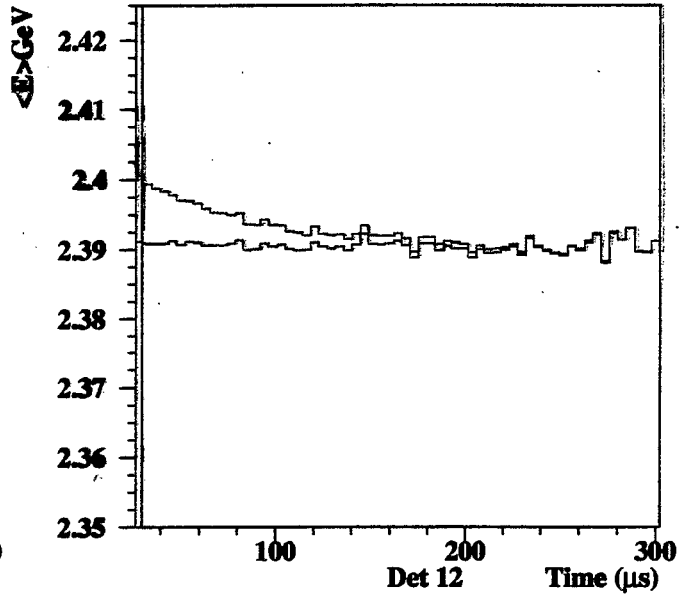
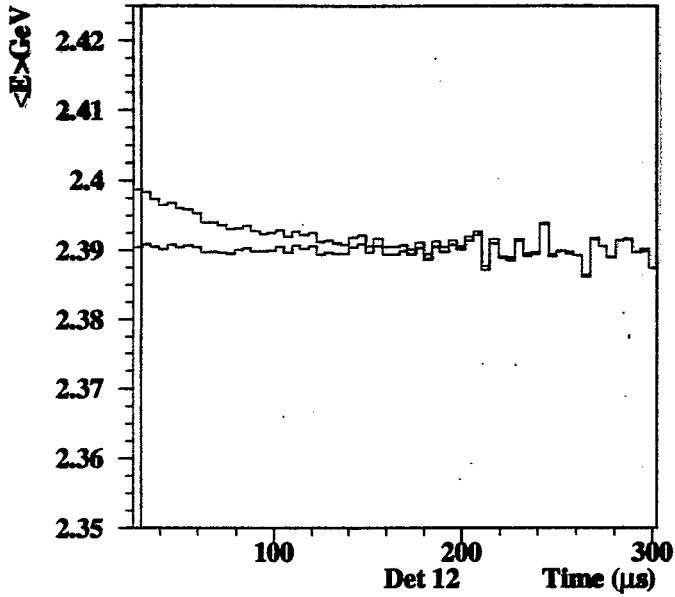
$$A = 0.06$$

wrong phase !!



— early
— late

Average Energies ; before and after removing the pileup



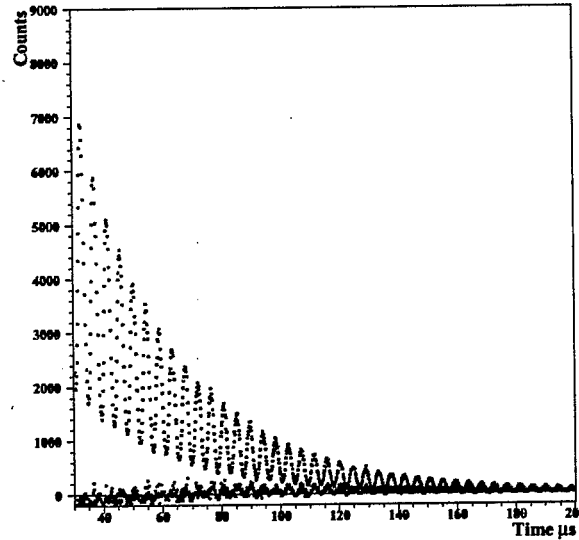
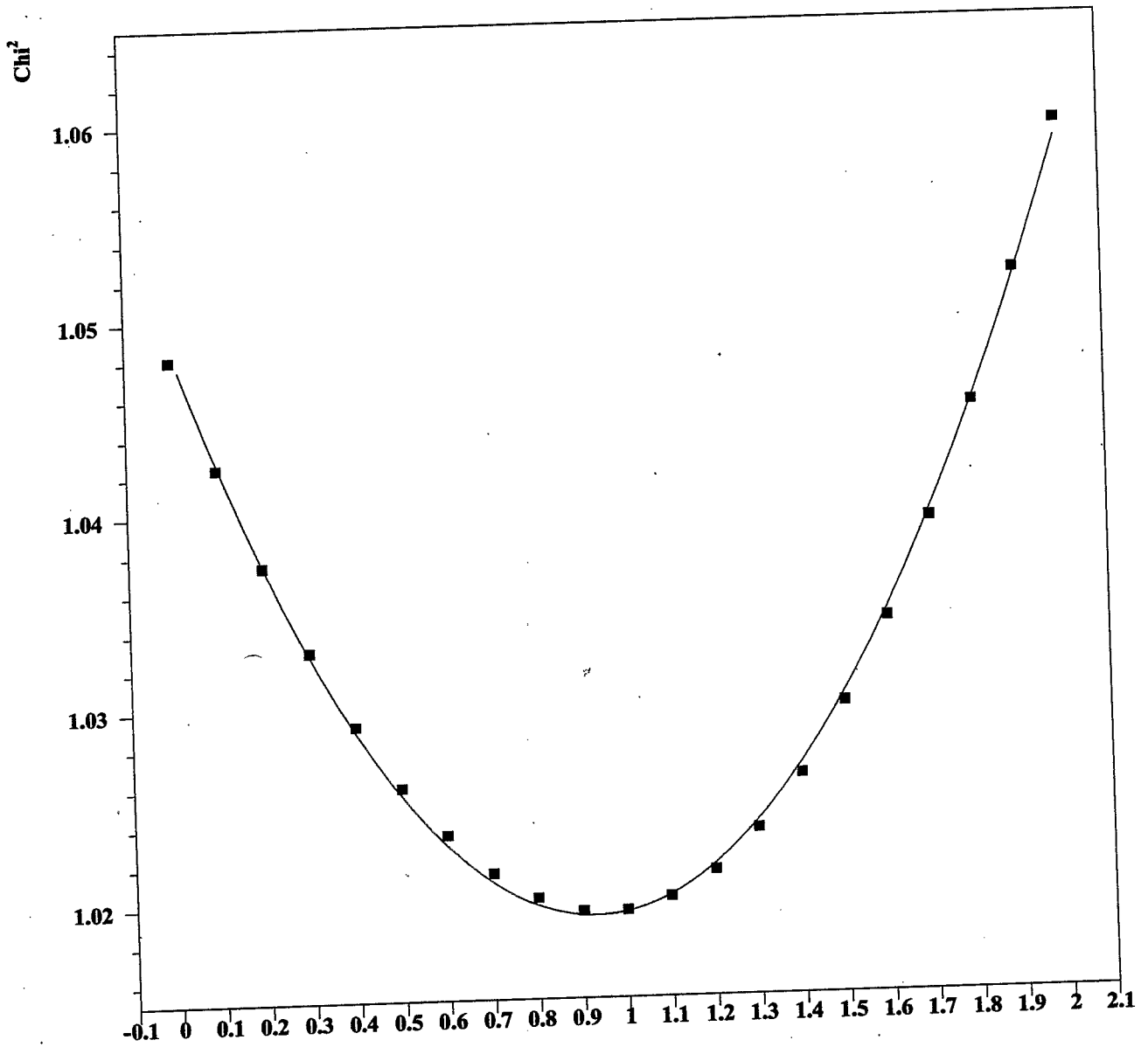


Figure 23 : The comparison of time spectra ($E \geq 3.5\text{GeV}$) with and without the pileup subtraction. No energy scale correction was done.

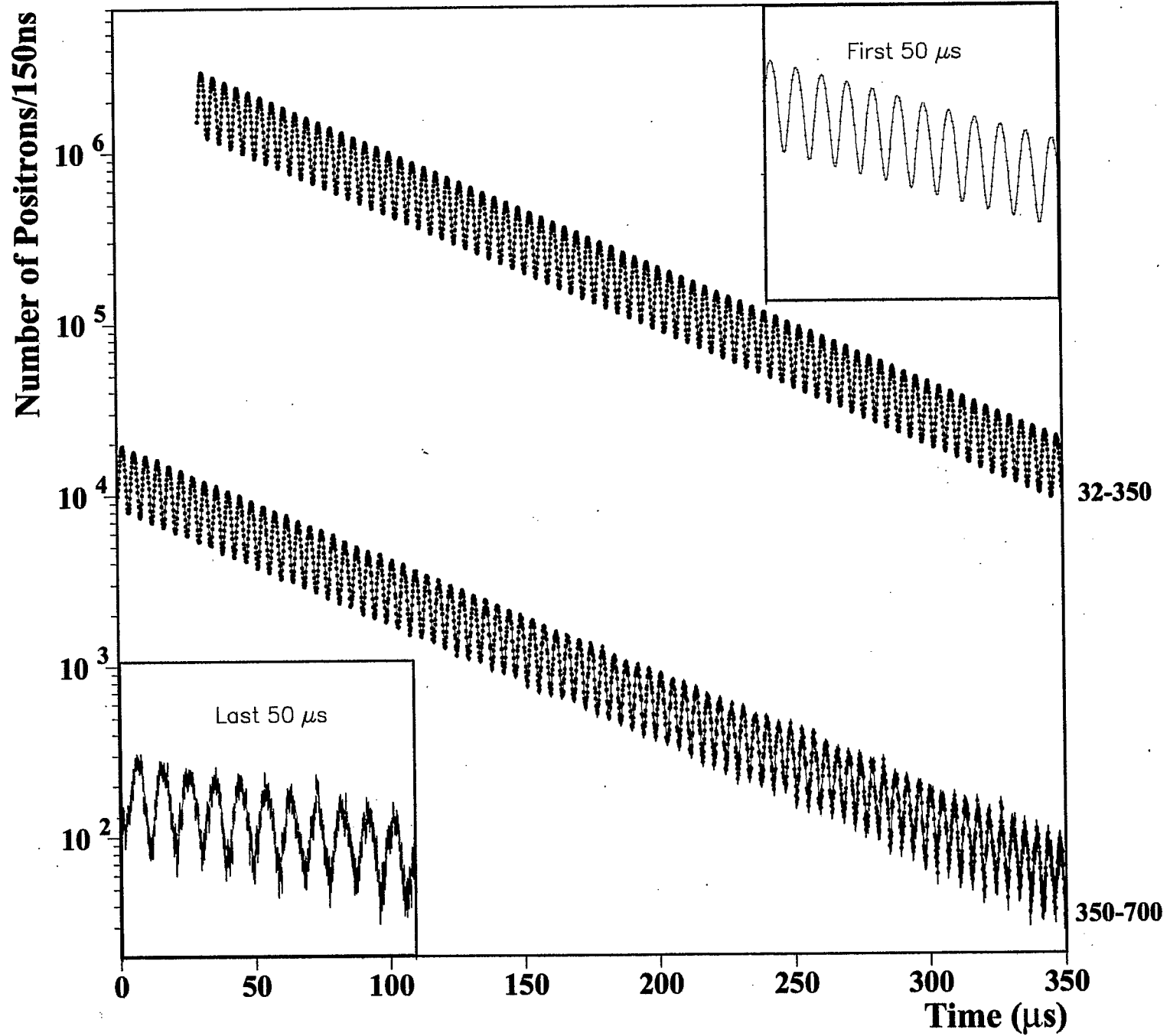
$E > 3.5 \text{ GeV}$ raw data
— after subtracting pileup

Pileup Systematic with smaller steps

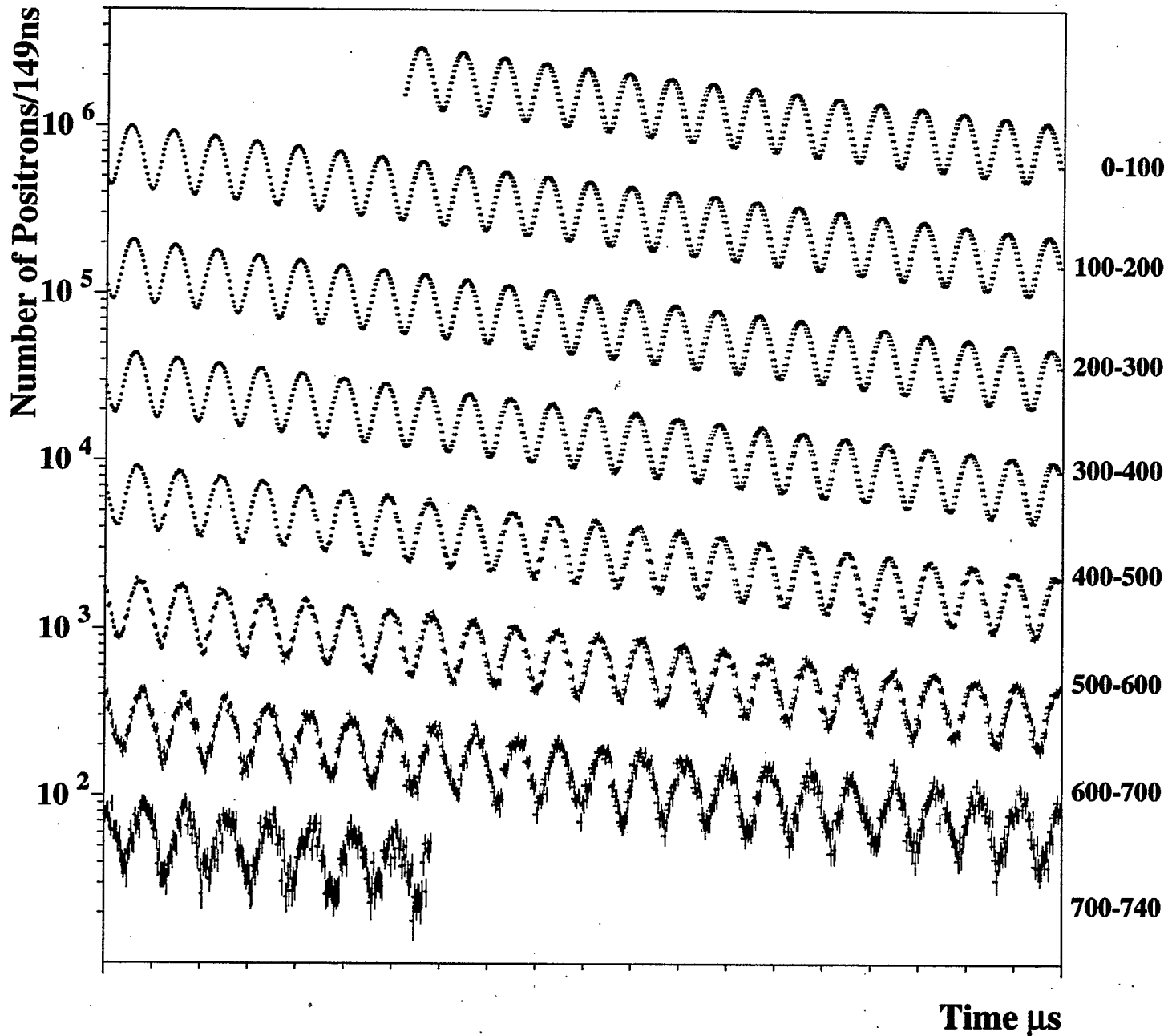




99 data



99 data - 10^9 e^+ > 2.0 GeV



Blind analysis

Separate analyses of ω_a

Values in ppm with common offset

	frequency	statistical error	systematic error
BNL group	143.37	1.24	0.17
Boston	143.25	1.24	0.25
Illinois	143.30	1.23	0.10
Ratio method 1	143.37	1.28	0.14
Ratio method 2	143.03	1.28	0.16

/

Two analyses of ω_p with common offset, weighted according to number of e^+ detected hour by hour,

agree to 0.03 ppm

Feb 2001

Values of muon anomalous moment

	$10^{-9} \times$	error (ppm)	
CERN with latest λ	11659 23	7.3	
BNL - 97	11659 25	12.9	
BNL- 98	11659 19.1	5.06	
BNL - 99	11659 20.20	1.34	
world average	11659 20.26	1.27	
Theory	11659 15.96	0.57	
difference	4.30	1.39	
difference (ppm)	3.69	1.39	2.65 σ

