

# Workshop – 2004 RHIC & AGS Annual User's Meeting

---

Hadronic Physics with Hadronic Beams at AGS:  
Baryons, Hyperons, and Pentaquarks

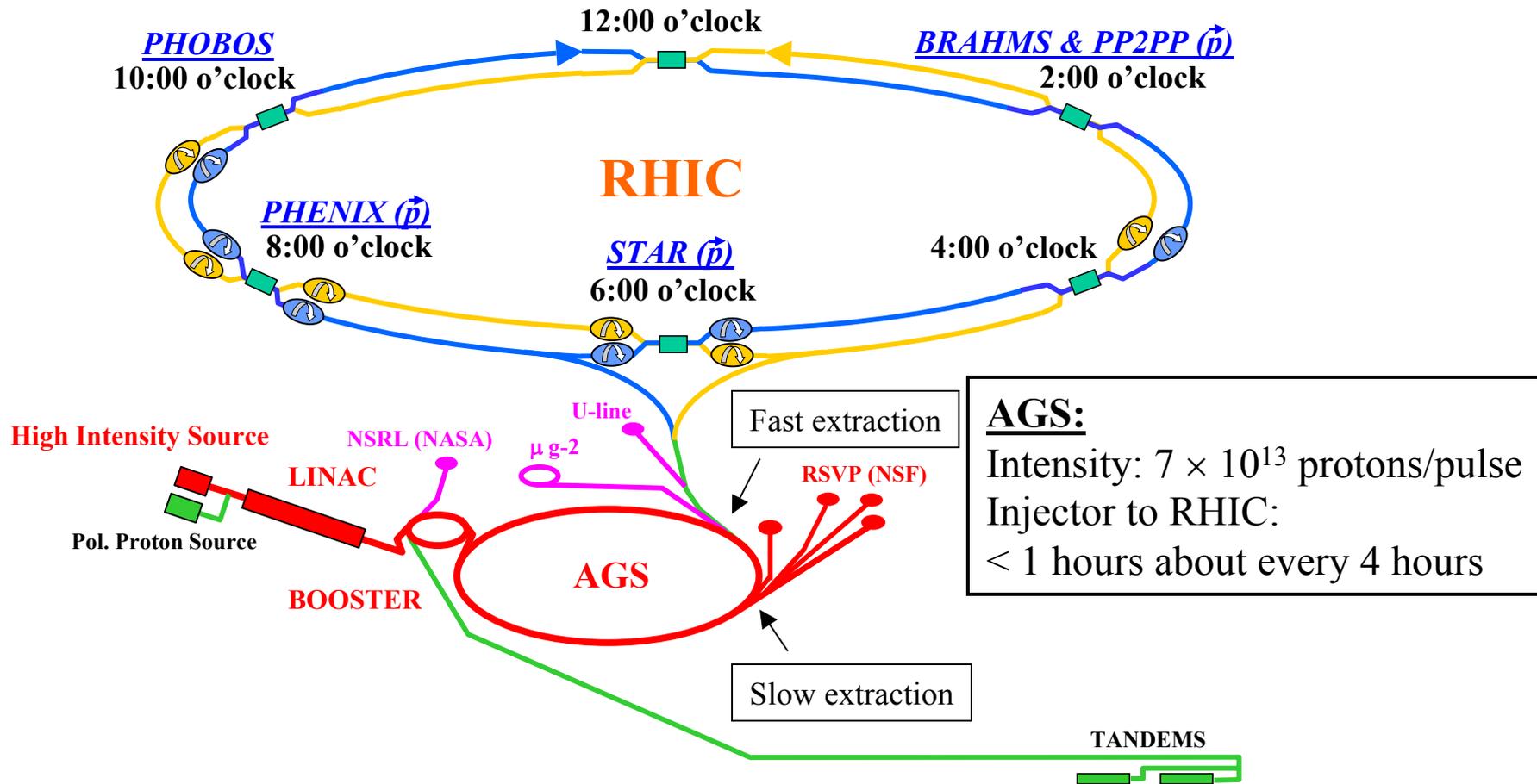
11 May 2004

## *AGS Fixed Target Future Plans*

Philip Pile  
Collider-Accelerator Department



# C-AD Accelerator Complex



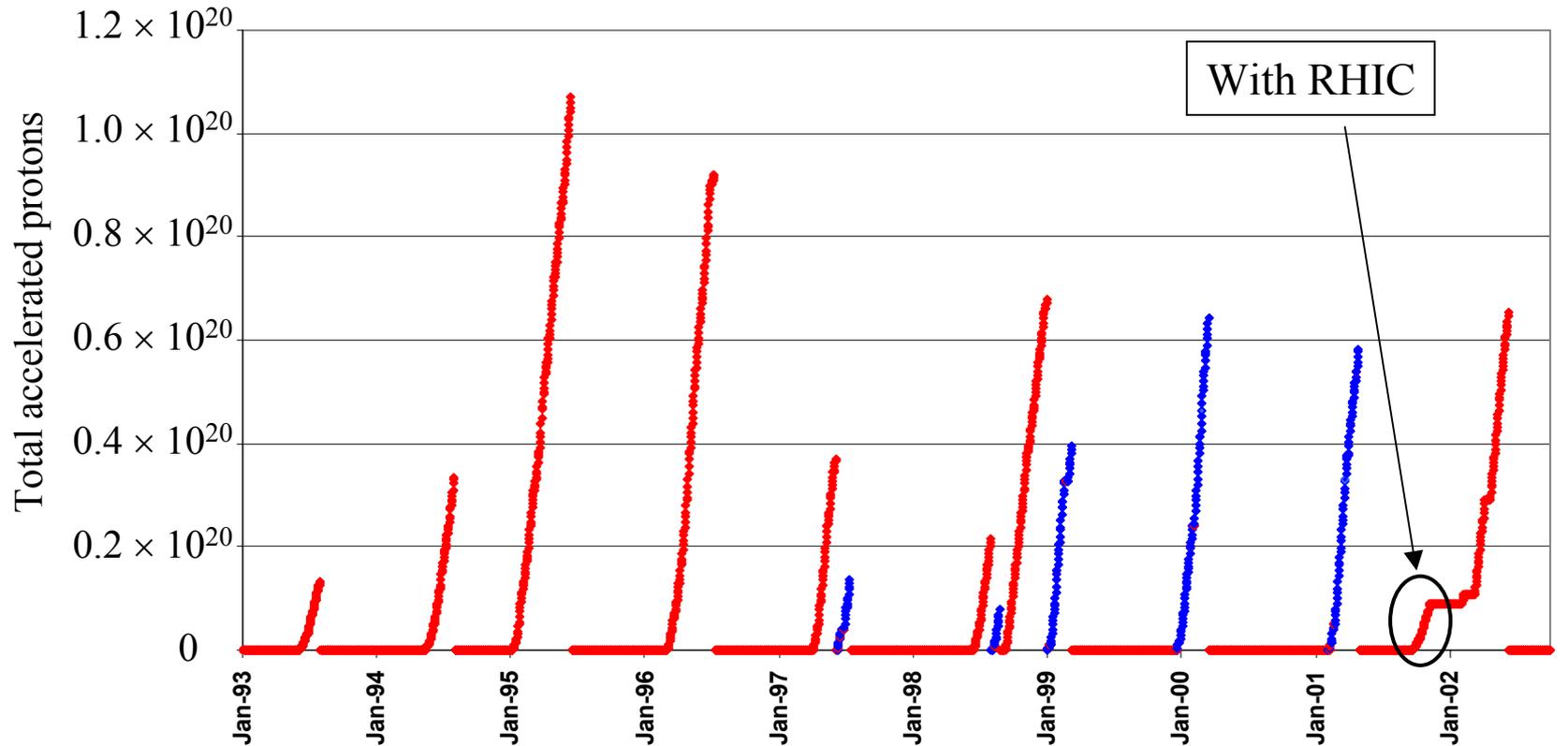
# AGS Performance

T. Roser

<b>PROTON BEAM</b>	<b>FY96</b>	<b>FY97</b>	<b>FY98/99</b>		<b>FY2000</b>	<b>FY2001</b>	<b>FY2002</b>	
	<b>SEB</b>	<b>SEB</b>	<b>FEB (g-2)</b>	<b>SEB</b>	<b>FEB (g-2)</b>	<b>FEB (g-2)</b>	<b>FEB (g-2)</b>	<b>SEB</b>
Beam Energy	24 GeV	24 GeV	24 GeV	24 GeV	24 GeV	24 GeV	24 GeV*	22 GeV*
Peak Beam Intensity	62 x 10 <sup>12</sup> ppp	62 x 10 <sup>12</sup> ppp	46 x 10 <sup>12</sup> ppp	72 x 10 <sup>12</sup> ppp	58 x 10 <sup>12</sup> ppp	61 x 10 <sup>12</sup> ppp	63 x 10 <sup>12</sup> ppp	76 x 10 <sup>12</sup> ppp
Total protons accelerated	0.9 x 10 <sup>20</sup>	0.4 x 10 <sup>20</sup>	0.1 x 10 <sup>20</sup>	0.9 x 10 <sup>20</sup>	0.4 x 10 <sup>20</sup>	0.5 x 10 <sup>20</sup>	0.6 x 10 <sup>20</sup>	0.7 x 10 <sup>20</sup>
Spill Length/Cycle Time	1.6 sec/3.6 sec	1.6 sec/3.6 sec		2.8 sec/5.1 sec				2.4 sec/5.4 sec
-> Duty Cycle	44%	44%		55%				44%
Spill Structure Modulation (peak-average) /average	20%	20%		20%				20%
Average Availability /Best Week	76% / 92%	71% / 79%	58 % / 67 %	71% / 88%	55 % / 83 %	74 % / 87 %	83 % / 88 %	85 % / 97 %
<b>HEAVY ION BEAM</b>	<b>Au</b>	<b>Au</b>	<b>Fe (NASA)</b>	<b>Au</b>	<b>Fe (NASA)</b>	<b>Fe (NASA)</b>	<b>Fe (NASA)</b>	<b>Fe (NASA)</b>
Beam Energy /nucleon	11 / 4 / 2 GeV	11 / 8 / 6 GeV	1.0 / 0.6 GeV	11 GeV	1.0 / 0.6 GeV	1.0 GeV	1.0 GeV	1.0 GeV
Peak Beam Intensity	4 x 10 <sup>8</sup> Au/p	17 x 10 <sup>8</sup> Au/p	20 x 10 <sup>8</sup> Fe/p	9 x 10 <sup>8</sup> Au/bunch	36 x 10 <sup>8</sup> Fe/p	17 x 10 <sup>8</sup> Fe/p	80 x 10 <sup>8</sup> Fe/p	49 x 10 <sup>8</sup> Fe/p
Spill Length/Cycle Time	1.4 sec/3.6 sec	1.5 sec/4.0 sec		1.2 sec/3.0 sec		0.9 sec/3.3 sec	0.9 sec/3.3 sec	0.9 sec/3.3 sec
-> Duty Cycle	39%	38%		40 %		27%	27%	27%
Spill Structure Modulation (peak-average) /average	<20%	<20%		<20%		<20%	<20%	<20%
Average Availability	80%	82 %	96 %	81 %	90 %	90 %	97 %	84 %

\* Westinghouse  
Motor Generator

# Total Accelerated Protons at the

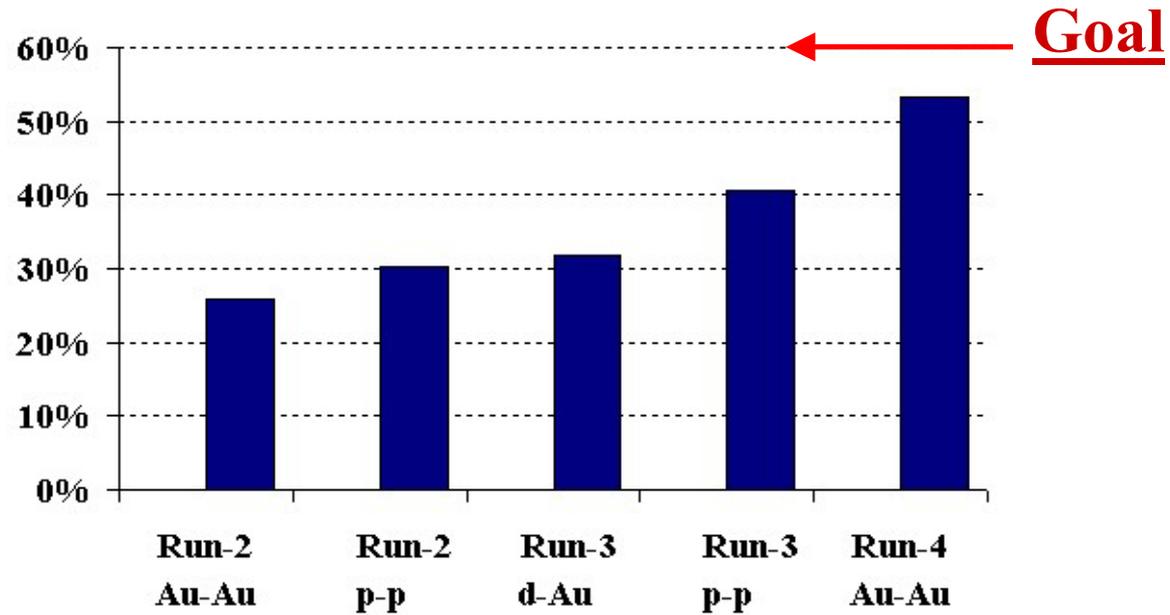


— Slow extracted beam (Kaon decay)

— Fast extracted beam (g-2)

Note: Lower total accelerated protons in later years due to much shorter running time

# RHIC Time in Store



# AGS Fixed Target Operations in RHIC Era –

## The Present Situation

- *AGS Fixed target experiments are no longer supported concurrent with RHIC operations since HEP base support was terminated at the end of 2002.*
- Two major HEP experiments submitted proposals to the NSF for construction and operations funds. These experiments, E926-K0PI0 and E940-MECO, collectively referred to as RSVP, were identified by the NSF for support and were included in the FY 2004 Presidents Budget request for a 2006 construction start.
- A third experiment, E949 ( $K^+$  rare decay) has submitted a proposal to the NSF with no action taken to date. A fourth experiment, E962 ( $\mu$  g-2) is reviewing options and may submit a new proposal to the laboratory and subsequently to the DOE and/or NSF.
- Seven other approved experiments await beam time with no clear path to success. In addition, three proposals and three LOI's are on the books with more to come.

# AGS Fixed Target Operations in RHIC Era –

## The Present Situation (cont')

- AGS fixed target experiments are still supported but are only scheduled outside RHIC operations and only if this can be done with minimal impact to RHIC activities. Full operations costs are recovered (roughly 0.5M\$/week). Two AGS experiments ran in this mode in FY2003 – Radiobiology (NASA-E966) and Proton Radiography (NNSA-E963), one is planned in 2004, K0PI0 (NSF-E926) test.

# AGS Fixed Target Operations in RHIC Era – The Future - RSVP Expectation

- The NSF plans to provide base support (\$'s for personnel and materials) for AGS operations for the RSVP experiments. With this support, concurrent operation with RHIC is planned as well as limited running outside RHIC operations.
- The present plan for the future is to simplify the AGS switchyard to help insure efficient running of the RSVP experiments. This will result in a more robust, low beam loss switchyard *at the expense of beam sharing*. The RSVP experiments are not impacted since they cannot share the beam due to different beam conditions required by each. Furthermore, the RSVP experiments are designed to accept the full AGS beam intensity so beam sharing is not planned.

# RSVP requirements

## KOPIO:

- $100 \times 10^{12}$  ppp; 5 s cycle; extraction @ 24 GeV [ $20 \times 10^{12}$  p/s ]
  - Proton throughput achieved
  - AGS injection energy upgrade to 2 GeV to reach  $100 \times 10^{12}$  ppp (R&D)
- Slow extraction with micro-bunching (200ps ( $\sigma$ ) every 40 ns)
  - Extraction with high frequency cavity (25 MHz, 100 MHz) (R&D)  
< 250 ps every 10 ns demonstrated; no change in losses
  - Extinction of  $10^{-3}$  (R&D)

## MECO:

- $40 \times 10^{12}$  ppp; 1 s cycle; extraction @ 8 GeV [ $40 \times 10^{12}$  p/s ]
  - No transition losses and reduced AGS injection losses (R&D)
- Slow extraction of 2 bunches with  $10^{-9}$  extinction during 1.35  $\mu$ s gap.
  - h=2 Booster operation; achieved  $22 \times 10^{12}$  protons per Booster cycle
  - Two Booster cycles per AGS cycle
  - Bunch merging to 2 bunches or develop h=1 Booster ops. (R&D)
  - Gap cleaning with resonant transverse kicker plus rf kicker in beam line (R&D)

*$\Rightarrow$  No protons left to share with other experiments*



May 04

# AGS Experimental Area

*AGS Experiments- as of FY 2004  
Approved/Proposed*

E962,  $\mu$  g-2  
E952,  $\nu$  mass **'OR'**

U- E963, Proton Rad; P945C (NNSA)

V1,  $\pi$   $\mu$  Beam Line

U Line

RHIC Transfer Line

D6-E930,  $\Lambda$  Hypernuclei - Ge Ball  
E961,  $\Lambda\Lambda$  Hypernuclei (CDS)  
E964,  $\Lambda\Lambda$  Systems (Emulsion-Counters) **'OR'**

A3 - *RSVP* MECO E940,  $\mu$  N  $\rightarrow$  eN  
(NSF  $\geq$  2005 construction start)

A3 - NASA Radiobiology (HI)

A3-E951,  $\mu\mu$  Collider Targetry  
P965, detector test

B5 - *RSVP* K0PI0  
E926,  $K_L^0 \rightarrow \pi^0 \nu \bar{\nu}$   
(NSF  $\geq$  2005 construction start)

**RHIC**  
eCooling Test Facility (under construction)

**C6/8-LESBII**  
(decommissioned - FY2003)

**C4-LESBIII**

E949,  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$   
E927,  $K_{e3}$ ; E953, hyperon spect;

**'OR'**

D-Target

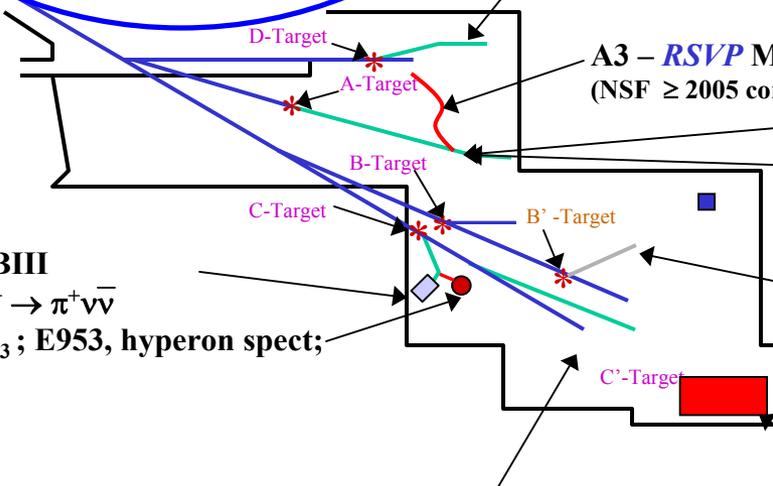
A-Target

B-Target

C-Target

B'-Target

C'-Target

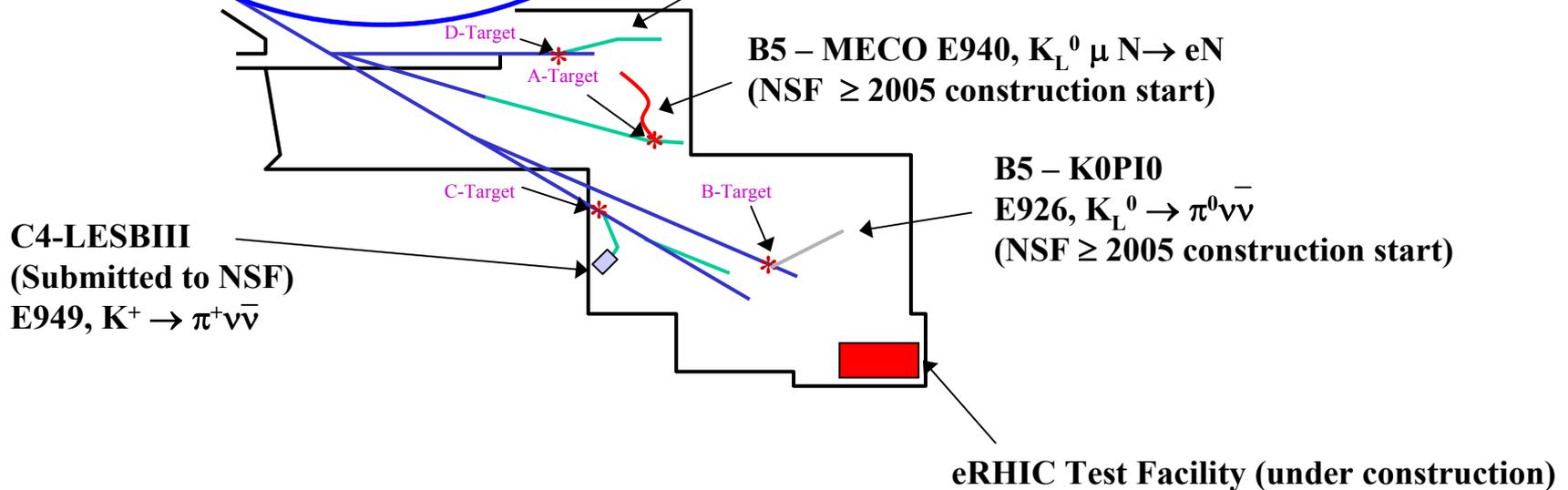


7 May 04

# AGS Experimental Area

*FY2004 Experiments  
What's happening this year*

•K0PI0, 4 day test



## The Plan

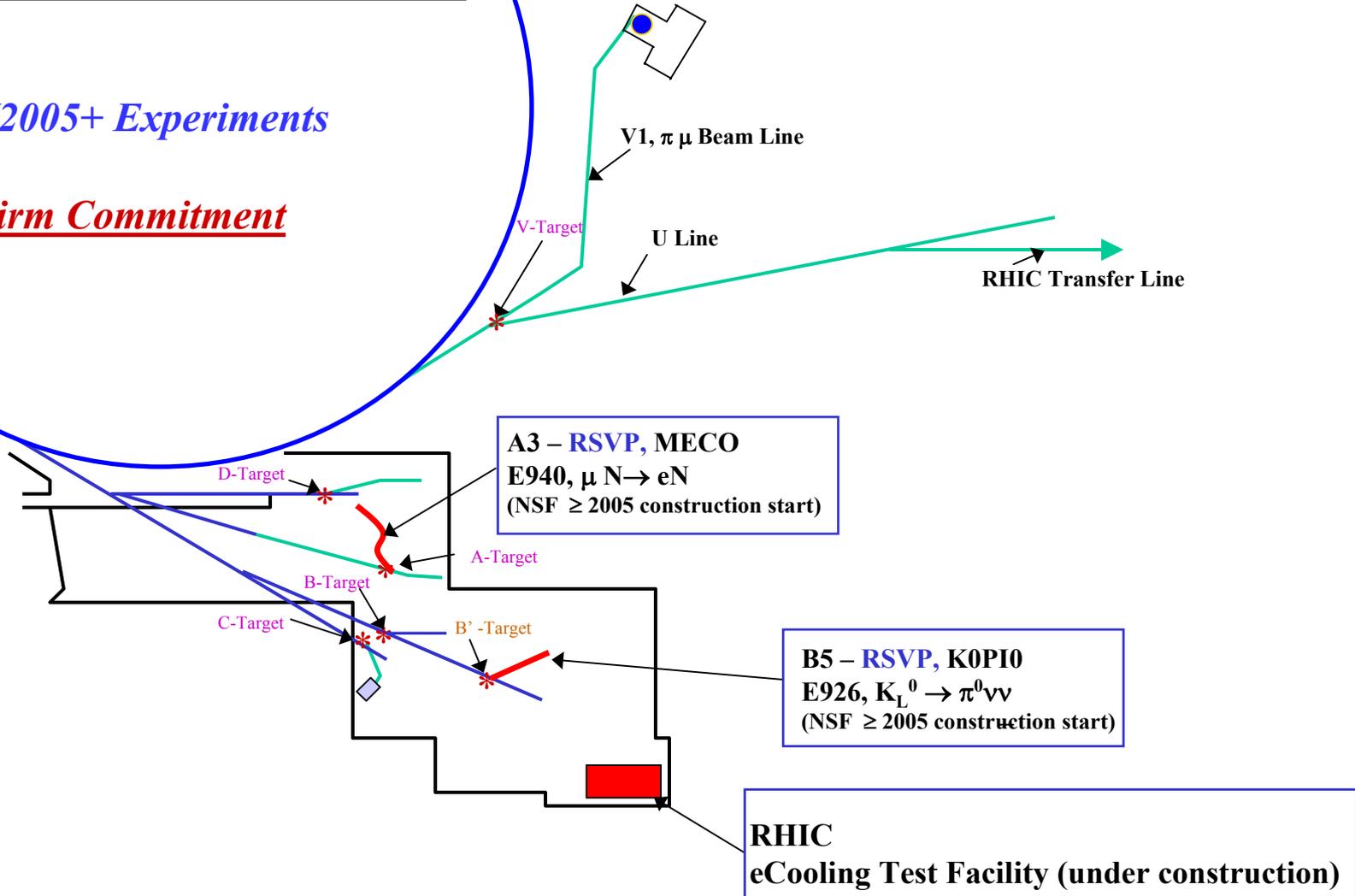
- D-Line Decommissioned
- A, B, C Lines run as “OR’s”

May 04

### AGS Experimental Area

*FY2005+ Experiments*

*~Firm Commitment*



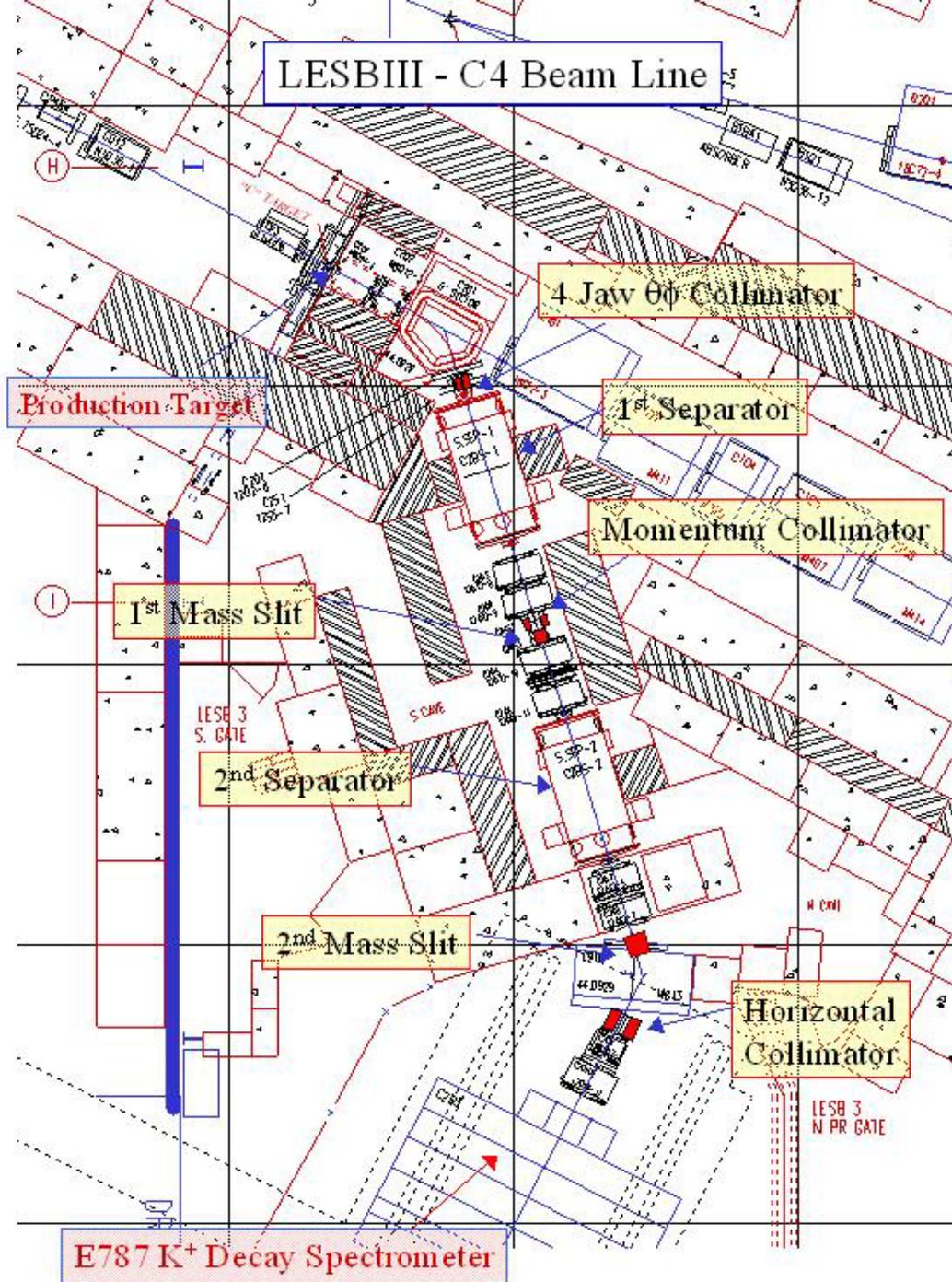
# AGS Fixed Target Operations in RHIC Era – The Future – RSVP Impact on other AGS Experiments, an example

- RSVP construction is expected to take about 5 years. Beam development for these experiments is planned between now and the end of construction.
- Once RSVP begins operation there will be no beam time available for other experiments for ~5-10 years.
- E949 has proposed to run in the C-line for about 3 years during RSVP construction.
- An example schedule follows – assumes we follow the PHENIX RHIC beam use proposal and maintain a B-C beam splitting capability during this period.

<b>RSVP and E949 Example Run Plan</b>												
<b>Assumptions</b>		1/26/2004 ppile										
FY 2003 Dollars, full NSF Indirects												
KOPIO total hours	8000	80 weeks		Beginning in 5th year of construction (last year)								
MECO total hours	4000	40 weeks		Beginning in 5th year of construction (last year)								
Power cost	\$ 85.00 per MWhr		FY 2006 and beyond									
		FY04	FY05	FY06#	FY07#	FY08#	FY09	FY10	FY11	FY12	FY13	FY14
<b>RHIC Weeks Total</b>		27	27	27	27	27	27	27	27	27	27	27
Available w ith pp*		0	5	0	19	0	19	0	19	0	19	0
Available w ith HI*		14	9	19	0	19	0	19	0	19	0	19
<b>With RHIC HI</b>		← MRE Construction Years →										
KOPIO		0	0	7	0	6	0	9.5	0	9.5	0	19
MECO		0	0	4	0	7	0	9.5	0	9.5	0	0
E949 w ith KOPIO		0	0	6	0	5	0	0	0	0	0	0
E949 alone		0	0	8	0	6	0	0	0	0	0	0
<b>With RHIC pp</b>												
KOPIO		0	0	0	9	0	7	0	9.5	0	19	0
MECO		0	0	0	6	0	7	0	9.5	0	0	0
E949 w ith KOPIO		0	0	0	8	0	0	0	0	0	0	0
E949 alone		0	0	0	4	0	0	0	0	0	0	0
<b>Without RHIC</b>												
KOPIO		1.4	4.3	0	0	0	0	0	4	5.5	8	8
MECO		1.2	2.6	0	0	0	0	0	4	2.5	0	0
E949 w ith KOPIO		0	0	0	0	0	0	0	0	0	0	0
E949 alone		0	0	0	3	1	0	0	0	0	0	0
<b>Fixed Costs (labor)</b>												
RSVP				\$ 2,421,818	\$ 2,421,818	\$ 2,421,818	\$ 2,421,818	\$ 2,421,818	\$ 2,421,818	\$ 2,421,818	\$ 2,421,818	\$ 2,421,818
KOPIO				\$ 379,875	\$ 379,875	\$ 379,875	\$ 379,875	\$ 455,850	\$ 455,850	\$ 455,850	\$ 759,751	\$ 759,751
MECO				\$ 379,875	\$ 379,875	\$ 379,875	\$ 379,875	\$ 683,776	\$ 683,776	\$ 683,776	0	0
<b>Other Costs (scales)</b>				<b>RSVP Commissioning</b>				<b>RSVP Running Years</b>				
RSVP				\$ 1,530,772	\$ 1,496,388	\$ 1,809,094	\$ 1,396,629	\$ 2,644,061	\$ 3,748,838	\$ 4,497,475	\$ 3,748,838	\$ 4,497,475
KOPIO		\$ 820,558	\$ 2,520,284	\$ 992,997	\$ 1,011,960	\$ 851,140	\$ 787,080	\$ 1,347,639	\$ 2,018,025	\$ 2,653,676	\$ 4,036,050	\$ 4,594,969
MECO		\$ 646,232	\$ 1,400,169	\$ 486,053	\$ 552,579	\$ 850,594	\$ 644,676	\$ 1,154,377	\$ 1,690,439	\$ 1,664,078	\$ -	\$ -
<b>TOTAL RSVP Operations Cost</b>		<b>\$ 1,466,789</b>	<b>\$ 3,920,453</b>	<b>\$ 6,191,391</b>	<b>\$ 6,242,495</b>	<b>\$ 6,692,397</b>	<b>\$ 6,009,952</b>	<b>\$ 8,707,521</b>	<b>\$ 11,018,745</b>	<b>\$ 12,376,673</b>	<b>\$ 10,966,456</b>	<b>\$ 12,274,012</b>
<b>E949 Operations Cost</b>				<b>E949 Running Years</b>								
Fixed (labor)				\$ 190,255	\$ 190,255	\$ 190,255						
Other (scales)				\$ 2,871,960	\$ 3,016,385	\$ 2,686,467						
<b>Total E949 Operations Cost</b>				<b>\$ 3,062,215</b>	<b>\$ 3,206,640</b>	<b>\$ 2,876,722</b>						
<b>Weeks per year</b>												
KOPIO		1.4	4.3	7	9	6	7	9.5	13.5	15	27	27
MECO		1.2	2.6	4	6	7	7	9.5	13.5	12	0	0
E949		0	0	14	15	12						
<b>Total Weeks of Physics Running</b>				<b>E949 Integrated Weeks</b>			<b>RSVP Integrated Weeks</b>					
KOPIO				9.5	23	38	65	92				
MECO				9.5	23	35						
E949		0	0	14	29	41						
								KOPIO Plan	12000 hrs	150	80 hr w eeks	
								MECO Plan	2777 hrs	35	80 hr w eeks	

\*Based on PHENIX Decadal plan and constant effort scenario through FY10

# assmes experiments pay for 12 FTE base manpower plus an additional 4 FTE (shared) during first 3 years of beam development



# Properties of C4 LESB-III Beam Line

- Maximum Momentum: 820 MeV/c
- 2 Stage Separation, 3<sup>rd</sup> order optics, special collimators
- Length: 19.6 meters
- ~ 30 TP/sec target limit
- Particle flux\* per  $1 \times 10^{13}$ , 25 GeV/c protons on 9 cm Pt target

<u>Particle</u>	<u>Momentum</u>	<u>Flux</u>	<u>Beam Purity</u>
K <sup>+</sup>	800 MeV/c	$4.8 \times 10^6$	71%
K <sup>+</sup>	Stopped	$1.0 \times 10^6$	



# Properties of D6 2GeV Beam Line

- Maximum Momentum: 1900 MeV/c
- 2 Stage Separation, 3<sup>rd</sup> order optics, special collimators
- Length: 31.6 meters
- Particle flux\* per  $3 \times 10^{13}$ , 25 GeV/c protons on 9 cm Pt target

<u>Particle</u>	<u>Momentum</u>	<u>Flux</u>	<u>Beam Purity</u>
K <sup>-</sup>	1800 MeV/c	$8.0 \times 10^6$	36%
K <sup>-</sup>	1800 MeV/c	$3.0 \times 10^6$	70%
K <sup>-</sup>	1200 MeV/c	$1.1 \times 10^6$	94%
K <sup>-</sup>	930 MeV/c	$3.8 \times 10^5$	97%
p	1800 MeV/c	$3.6 \times 10^6$	47%
p	1400 MeV/c	$1.2 \times 10^6$	99%

\* Proton flux in D-line limited by beam loss considerations, up to  $7.6 \times 10^{13}$  protons/AGS cycle have been extracted.

# AGS High Energy Physics Experiments - FY 2005+

---

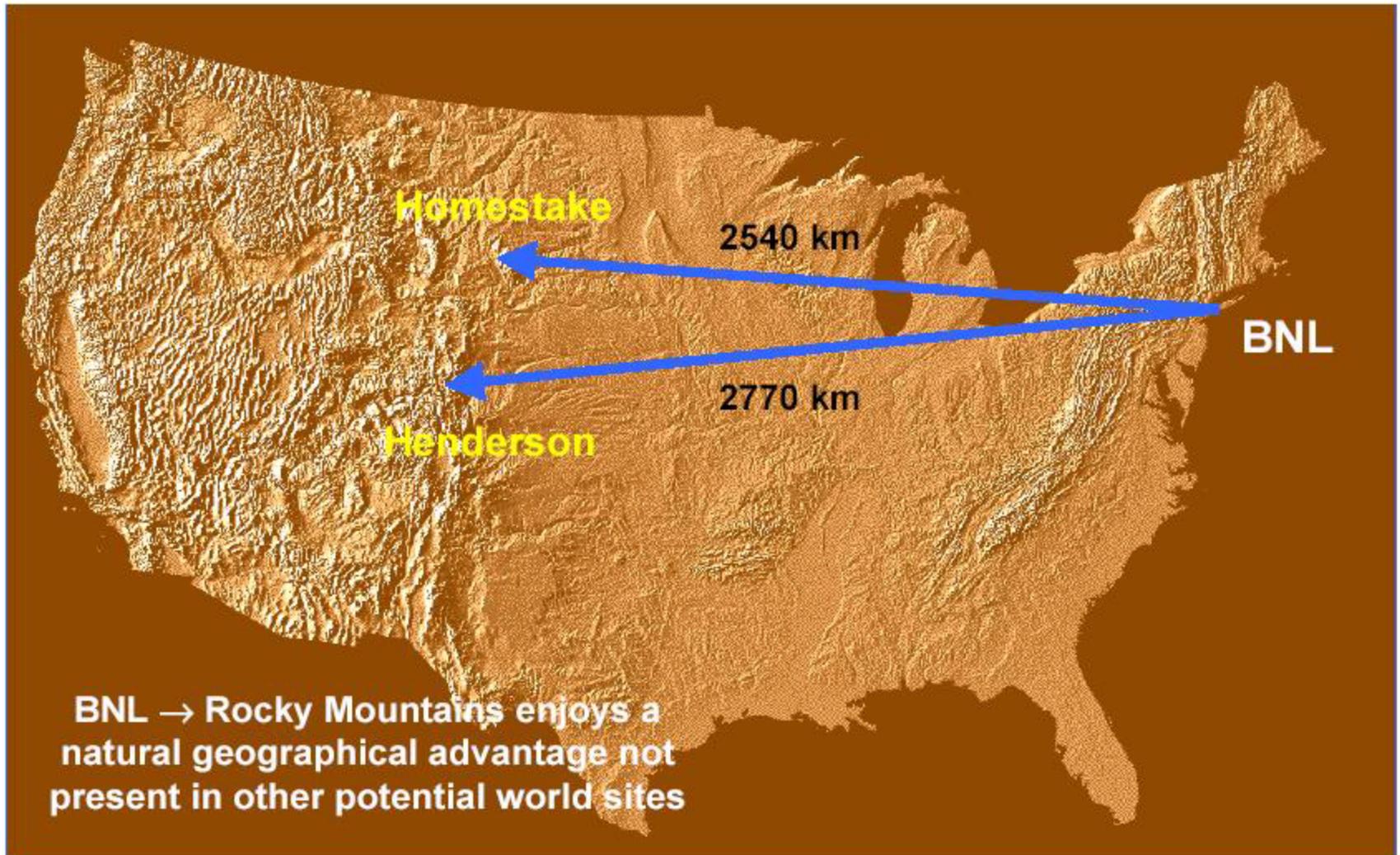
- **E949 – Rare K decay** (2338/6000 hrs)
  - Was approved (DOE) to run through FY 2003, however no funds for AGS HEP operations were identified in the FY2003 presidential budget request
  - Experiment needs two more long running years to complete goals
  - Proposal to NSF to run concurrent with RSVP beam studies
  
- **E951 - Muon Collider Target Studies**
  - Will continue subject to MCOG guidance and restoration of DOE HEP funding – However, once E940 beam line construction begins (2006?) the A3 beam will no longer be available
  
- **E952 - Muon Neutrino Mass (g-2 storage ring)**
  - Scientific approval from BNL, submitted to DOE for approval to run, will not be considered unless if DOE HEP funding for the AGS is restored. Could possibly run parasitic to E949 once E962 is complete.
  
- **E962 – muon g-2** (Continuation of E821) (0/2000 hrs)
  - A request has been submitted to DOE to run this experiment, however the FY 2003 funding problem lead to no action at the DOE level. Planning to submit new proposal (upgraded from E962) to BNL and if again judged “Must Do” seek funding from DOE and/or NSF.
  
- **RSVP - Rare Symmetry Violating Processes**
  - NSF Major Research Equipment (MRE) proposal (\$100M+); ≥ FY 2005 construction start
  - MECO, E940,  $\mu^- \rightarrow e^-$  conversion
  - K0P10, E926, Rare K decay  $K_L^0 \rightarrow \pi^0 \nu \bar{\nu}$

# AGS High Energy Experiments - FY 2005+

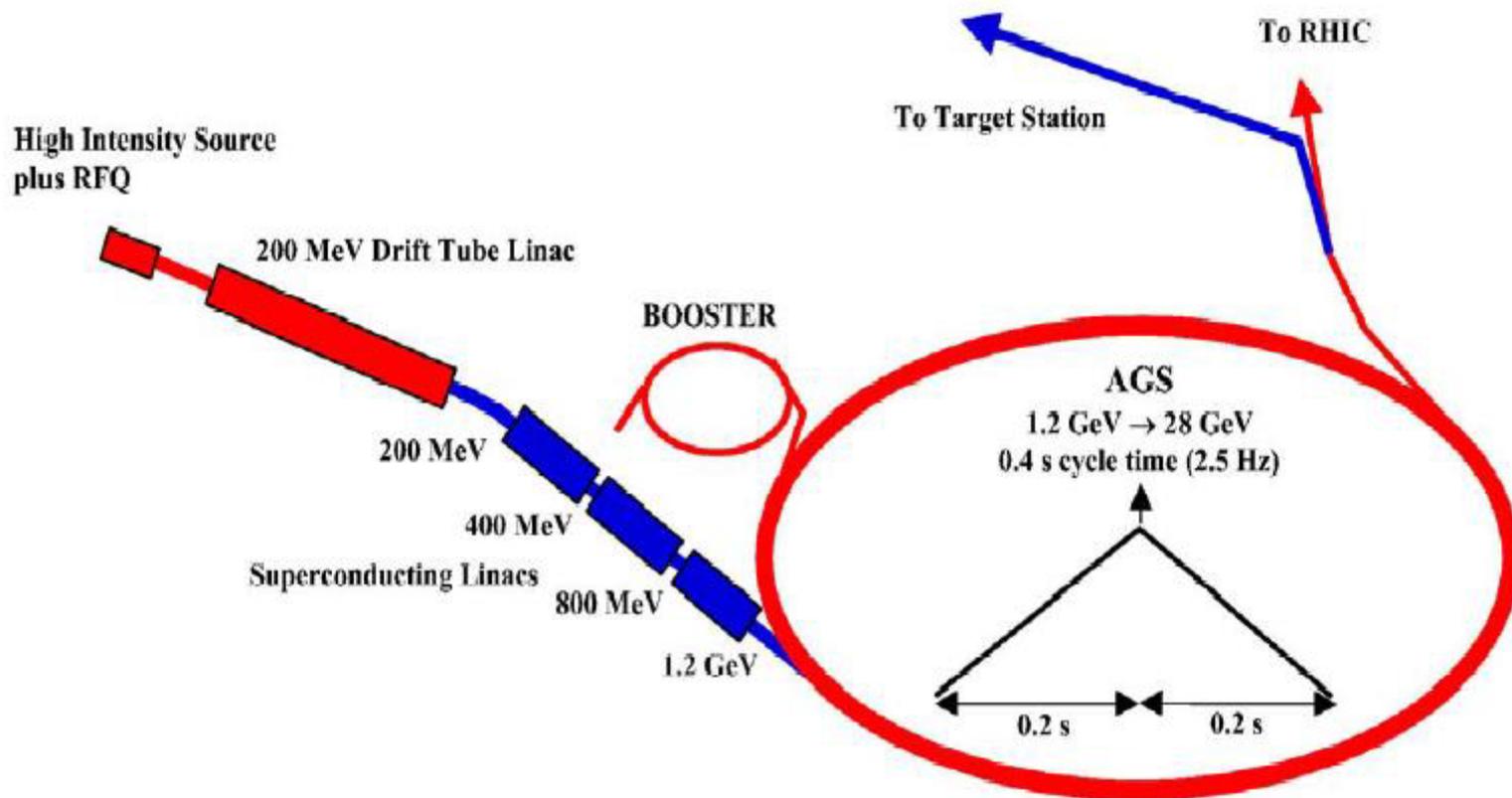
---

- **P965 – A3**, [1 week detector test (Submitted Aug 2002, deferred)]: *Proposal to Measure the Efficiency of Electron Charge Sign Determination up to 10 GeV in a Magnetized Liquid Argon Detector ( $\mu$ LANNDD)*  
BNL, UCLA, Texas, Zurich, Hawaii, Napoli, Pisa, Princeton  
McDonald (Princeton)
- **LOI – V1**, (g-2 storage ring) *Muon Electric Dipole Moment Experiment*, Boston, BNL, BINP, Cornell, Heidelberg, Illinois, Indiana, Yale  
Semertzidis (BNL), Miller (Boston)
- **LOI – New Beam**, [neutrino oscillation experiment]: *Neutrino Physics with Detectors at Baselines of 100-1000 km from BNL*, BNL, UCLA, Texas, Zurich, Hawaii, Napoli, SUNY-SB, Pisa, Princeton  
Diwan (BNL)

# BNL → Rocky Mountains Super Neutrino Beam

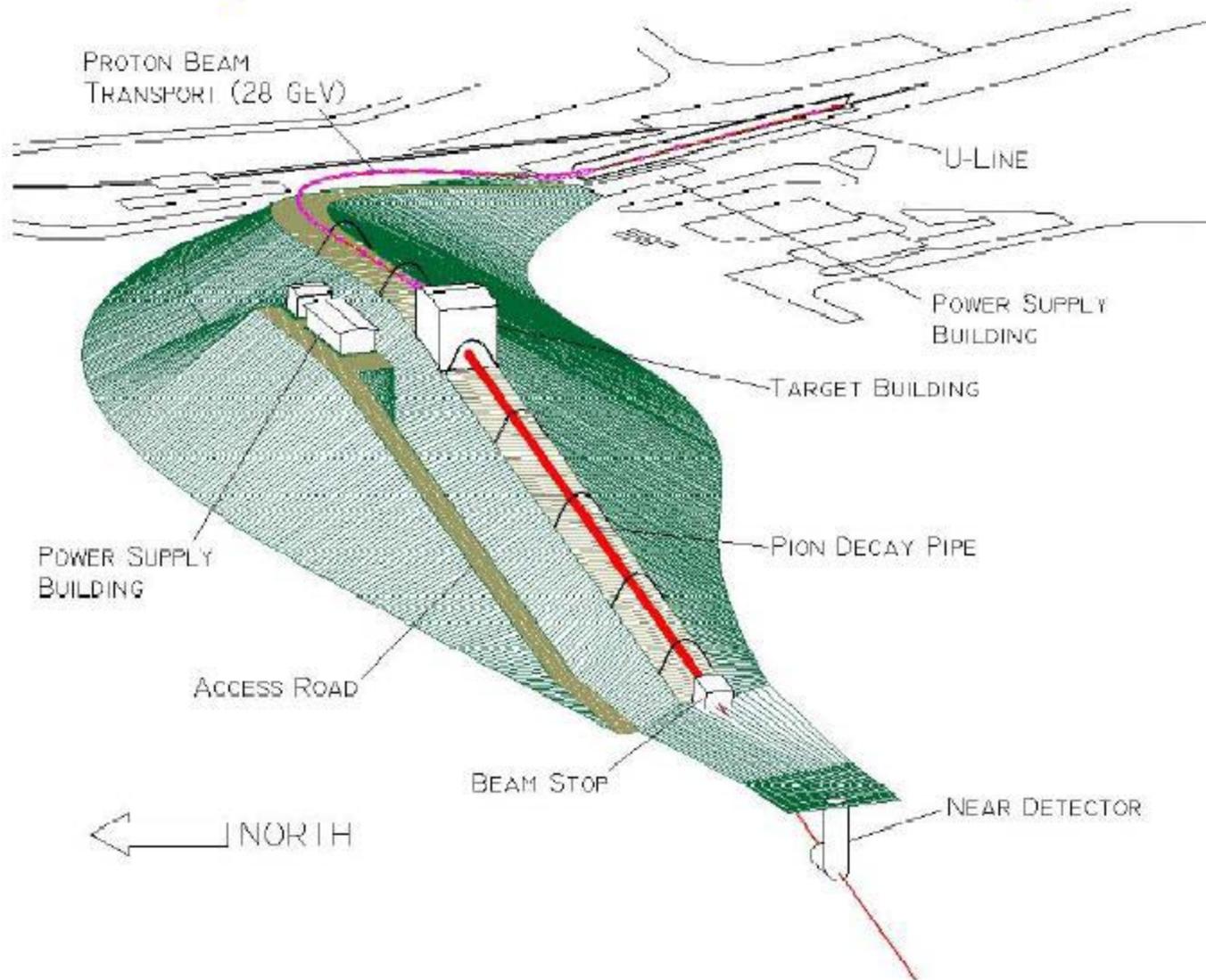


# 1-2 MW Super Neutrino Beam at AGS

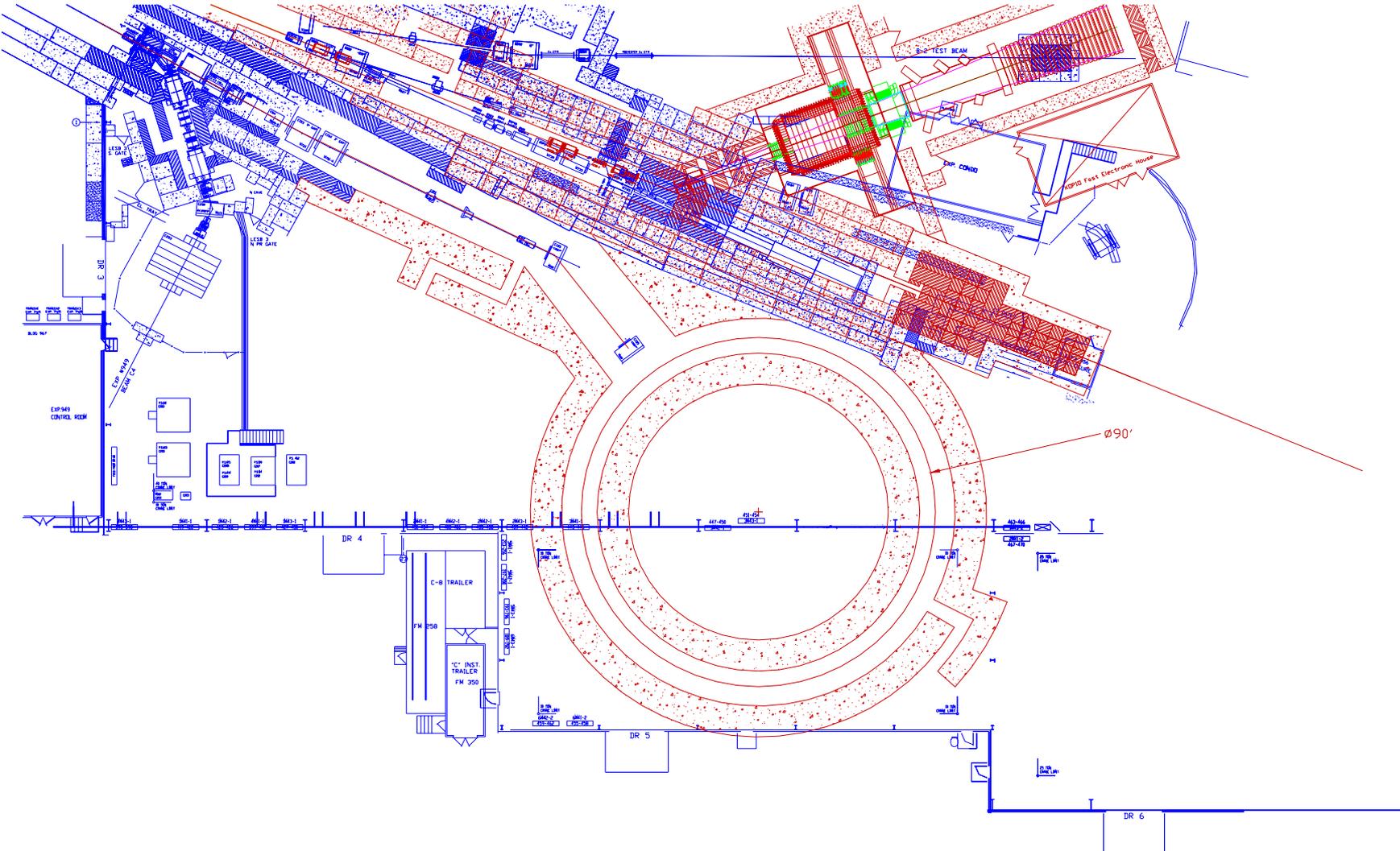


- BNL will complete in June 2004, a Conceptual Design in preparation for a proposal to DOE to upgrade the AGS to 1-2 MW target power and construct the wide-band **Super Neutrino Beam** as listed in the DOE's "Facilities for the Future of Science" plan of November 2003

# 3-D Super Neutrino Beam Perspective



# deuteron EDM experiment (Semertzidis proposal in preparation) – C3 Beam Location



# AGS Medium Energy Experiments - FY 2005+

---

## Crystal Ball (shipped to Mainz in 2003, possible return to BNL in 3-4 years)

- E927 - C2/D6, [233 (tests)/2280 hrs]: *Measurement of the  $K_{e3}^+$  decay rate and spectrum*, UCLA, JINR, ACU, ANL, Az.U, Boskov, U.Colo, GWU, U.Karlsruhe, Kent, PNPI, Regina, Valparaiso – **(subject to final approval - technical issues)**  
Nefkens (UCLA), Comfort (Arizona)
- E953 - C2/D6, [0/450 hrs]: *Neutral Hyperon Spectroscopy with the Crystal Ball*, UCLA, JINR, ACU, ANL, Az.U, Rudjer Boskovic, U.Colo, GWU, Karlsruhe, Kent State, PNPI, Regina, Maryland, Uppsala, Valparaiso  
Manley (Kent State), Spinka (ANL), Nefkens (UCLA)
- EOI – D6, (Xstal Ball in D6) *Search for Lambda Radial Trajectory 1 (1/2-,3/2-)*,  
Peaslee (Maryland)

# AGS Medium Energy Experiments - FY 2005+

---

## D6 (2GeV) Beam Line

- \* E930 – D6, [1002/1428 hrs]: *High resolution spectroscopy of hypernuclei using large acceptance germanium detector*, Tokyo, Tohoku, Osaka, BNL, CMU, China Inst. Of Atomic Energy, Kyoto, Hampton, North Carolina A&T, Freiburg  
Tamura (Tohoku)
- \* E961 – D6, (continuation of E906 - CDS) [1200 hrs] *Search for Double- $\Lambda$  Hypernuclei by Sequential Pionic Decays*, Osaka Electro-Communication U, KEK, Tokyo, BNL, CMU, Freiburg, Hampton, INR-Moscow, Kyoto, TRIUMF, Manitoba, New Mexico, Osaka, Pusan National, Temple  
Fukuda (Osaka Electro-Communications), Rusek (BNL), Chrien (BNL)
- \* E964 – D6, [0/1400 hrs] Hybrid Emulsion, : *Systematic Study of Double Strangeness System by an Emulsion-Counter Hybrid Method*, BNL, CMU, Gifu, China Inst. Of Atomic Energy, Kyoto, UNM, Osaka City, Pusan, Tokoyu, Tokyo  
Imai (Kyoto), Tamura (Tokoyu), Nakazawa (Gifu)
- LOI – D6, (LANL MRS spectrometer for D6) *Double Strange Nuclei*, BNL, CMU, Houston, KEK, Tohoku  
May (BNL)

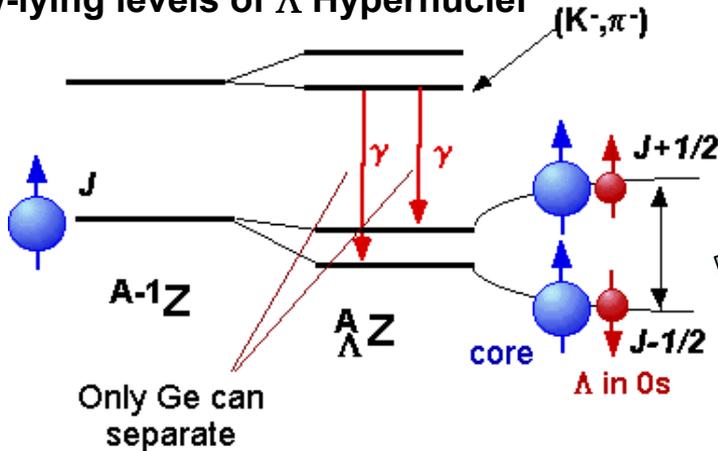
\* These experiments were proposed by physicists from Japan interested in bridging the gap between now and the time the Japanese Hadron Facility (a.k.a. JKJ) comes on line. Equipment funds for these experiments come from Japanese sources. 27

# E930 ('98,'01)

## Precision $\gamma$ Spectroscopy of $\Lambda$ Hypernuclei for $\Lambda N$ Interaction

### Motivation

Low-lying levels of  $\Lambda$  Hypernuclei



2-body  $\Lambda N$  effective interaction

$$V_{\Lambda N}^{\text{eff}} = V_0(r) + \underbrace{V_\sigma(r) \vec{s}_\Lambda \vec{s}_N + V_\Lambda(r) \vec{l}_{\Lambda N} \vec{s}_\Lambda + V_N(r) \vec{l}_{\Lambda N} \vec{s}_N + V_T(r) S_{12}}_{\Lambda N \text{ spin-dependent interaction strengths}}$$

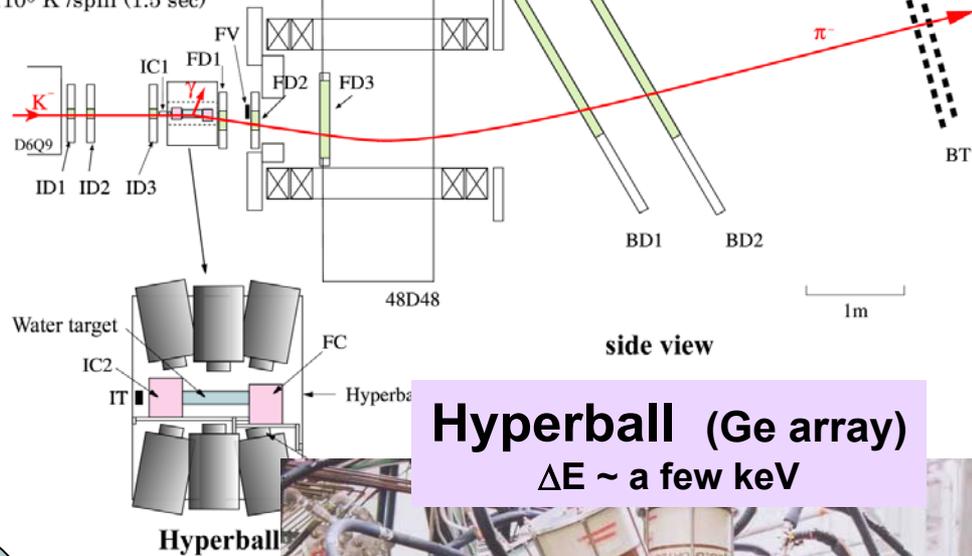
### Results

$\gamma$  rays from  ${}^9_\Lambda\text{Be}$ ,  ${}^{16}_\Lambda\text{O}$ ,  ${}^{15}_\Lambda\text{N}$ ,  ${}^7_\Lambda\text{Li}$  (8 new  $\gamma$  rays) were observed.  
 -> All of  $\Delta$ ,  $S_A$ ,  $S_N$ ,  $T$  were determined and cross-checked.

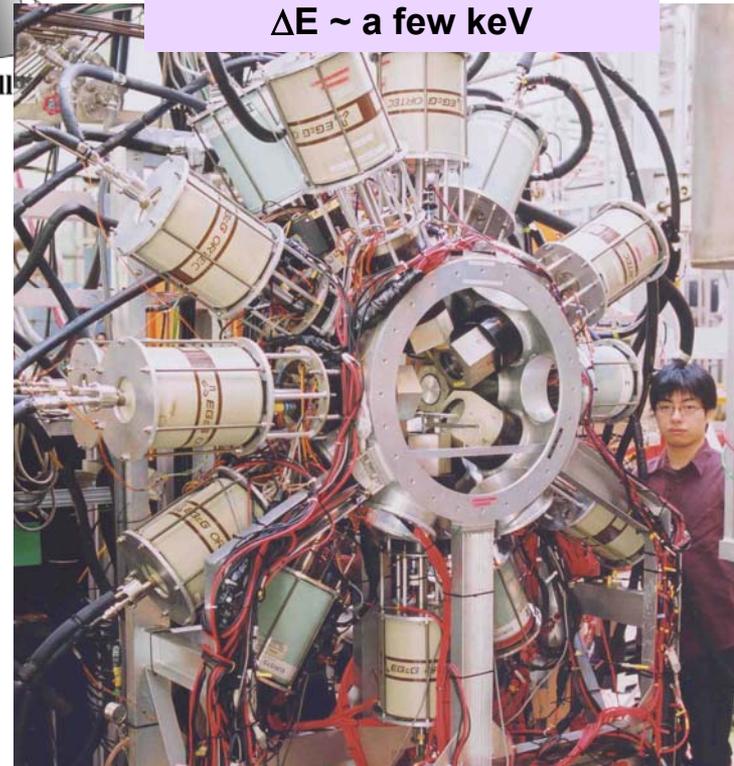
### Setup

BNL-AGS D6 line

930 MeV/c,  $K^- : \pi^- = 3:1$   
 $2.3 \times 10^5$   $K^-/\text{spill}$  (1.5 sec)

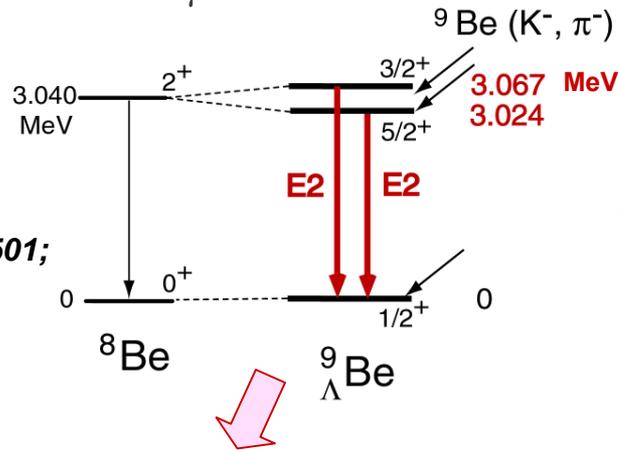
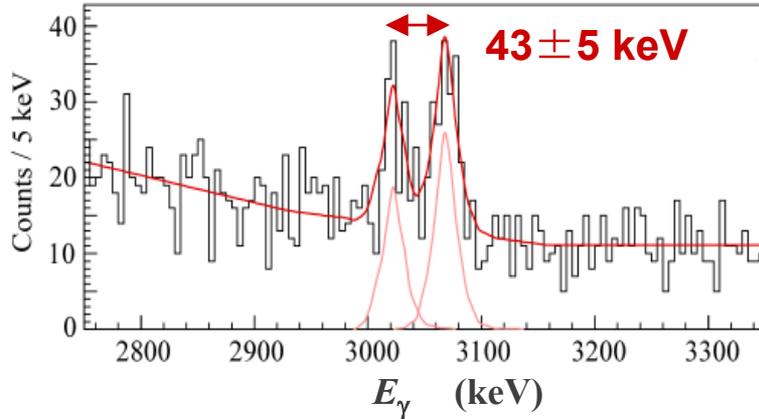


**Hyperball (Ge array)**  
 $\Delta E \sim$  a few keV



# Highlights of E930 results: Observation of “Hypernuclear Fine Structure”

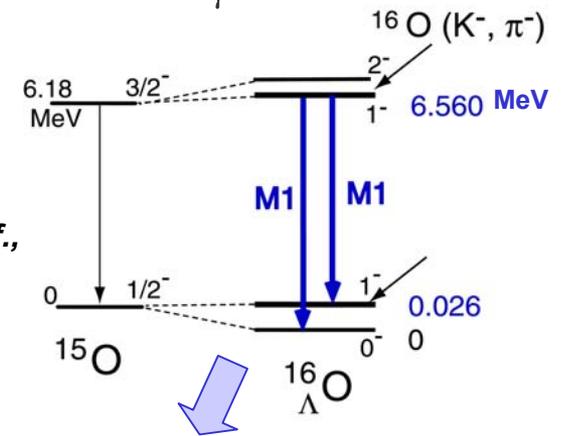
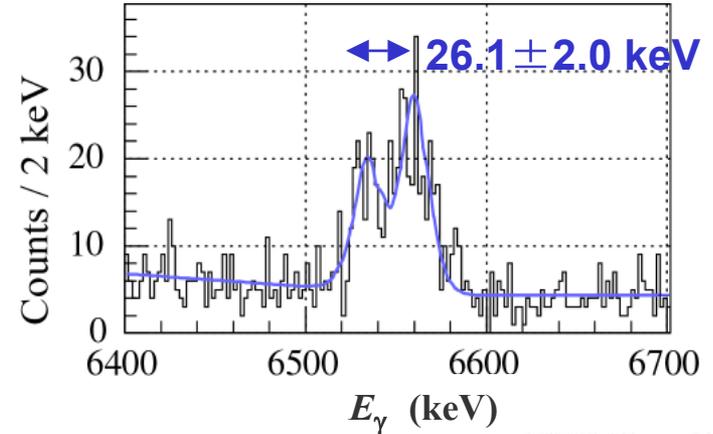
${}^9\text{Be} (K^-, \pi^- \gamma) {}^9_{\Lambda}\text{Be}$



Akikawa et al.,  
PRL 88 (2002) 082501;  
Tamura et al.,  
proc. HYP03 conf.

$\Lambda\text{N}$  spin-orbit force:  $S_{\Lambda} = -0.01 \text{ MeV}$   
= Consistent with a quark model

${}^{16}\text{O} (K^-, \pi^- \gamma) {}^{16}_{\Lambda}\text{O}$



Ukai et al.,  
proc. HYP03 conf.,  
to be published.

$\Lambda\text{N}$  tensor force:  $T = 0.03 \text{ MeV}$   
= Consistent with meson-exchange models

# ***BNL-E964: (Approved by AGS-PAC Oct. '01)***

## **International collaboration AGS-E964(BNL)**

### **Systematic Study of Double Strangeness System with an Emulsion-Counter Hybrid Method**

#### **AGS-E964 collaborators (now)**

**USA:** BNL R. E. Chrien, M. May, P. Pile, A. Rusek  
 CMU G. B. Franklin  
 Houston Ed. Hungerfold, K. J. Lan, Y. Cui, Song  
 New Mexico B. Bassalleck

**CHINA:** CIAE H. Guo, Z. Liu, S. Lu, J. Zhou

**INDIA:** AMU R. Hasan

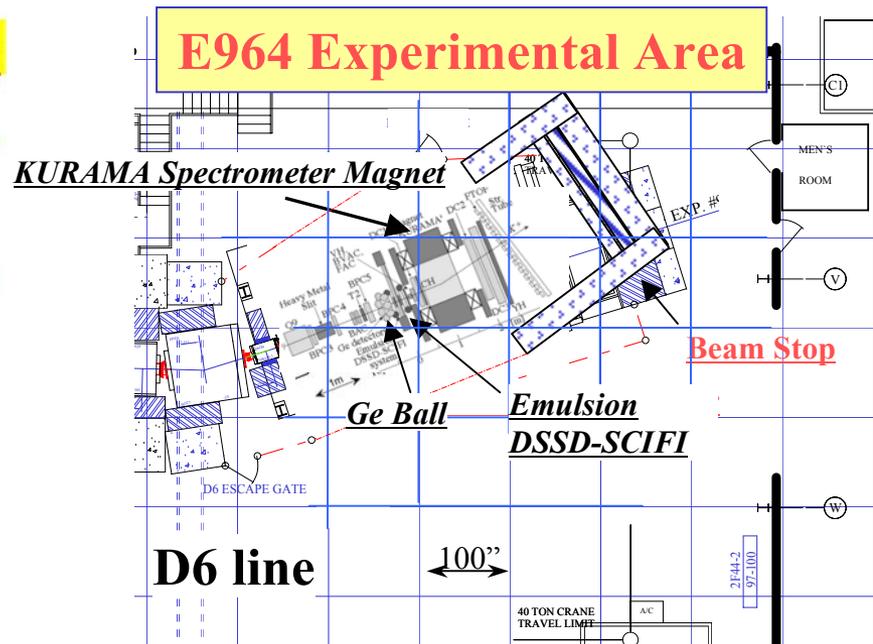
**JPN:** Gifu K. Nakazawa  
 Kyoto H. Funahashi, K. Imai, N. Saito, M. Yosoi  
 OCU K. Yamamoto, T. Yoshida  
 Osaka E-C H. Hotchi  
 Riken K. Tanida  
 Toho S. Ogawa, H. Shibuya  
 Tohoku Y. Miura, H. Tamura, M. Ukai

**KOREA:** GNU J. S. Song, C. S. Yoon  
 PUSAN J. K. Ahn, S. J. Kim

**MM:** MANDALAY M. Thein, K. S. Myint

**UK:** UCL D. H. Davis, D. Tovee

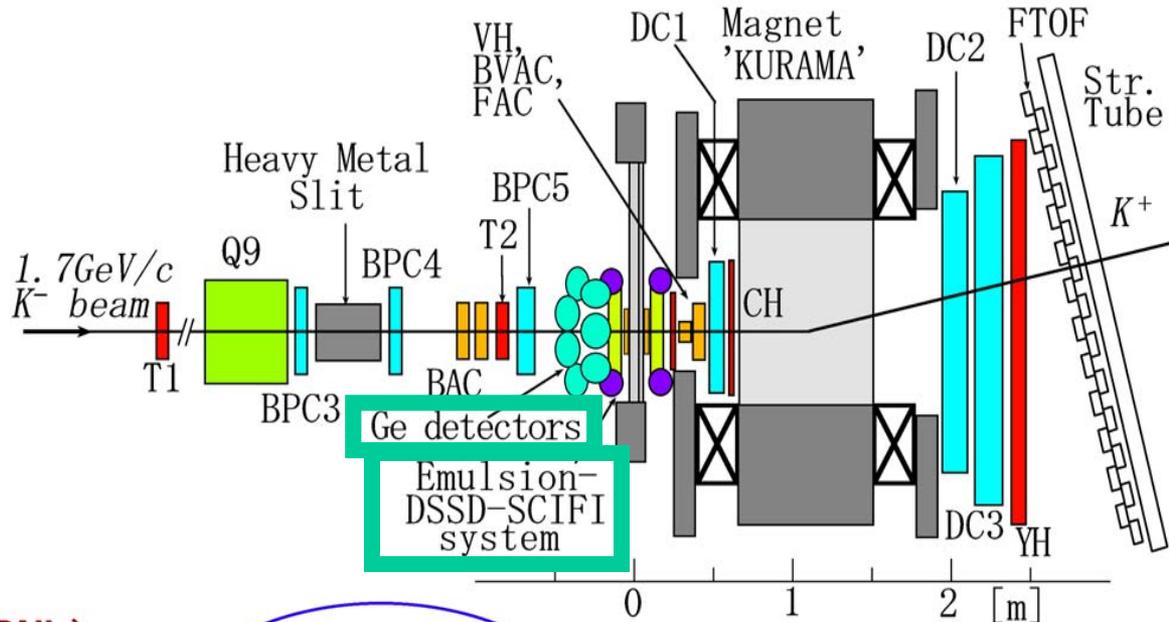
**+ Graduate students**



Beam :  $K^-$  (1.7GeV/c),  
 $2 \times 10^5 K^-/\text{spill}$  (4.1sec) with  $K^-/\pi > 9$   
Time requested : **100 hours** for detector tune  
**100 hours** for beam tune  
**1100 hours** for beam exposure  
 Beam time in **FY2006** is required.

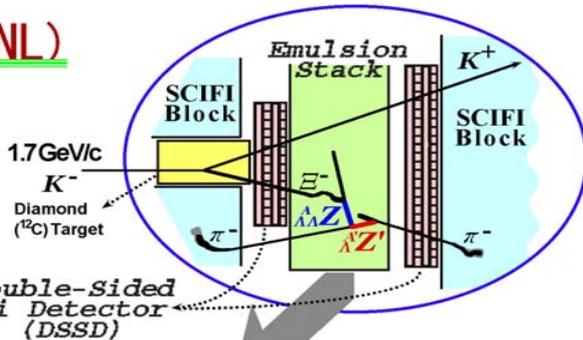


# Setup around the target (E964)

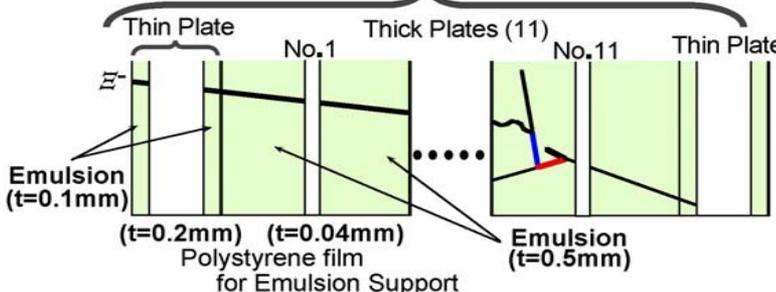


## AGS-E964 (BNL)

Apparatus around the Target



### Emulsion Stack



## DSSD (Double-Sided Si Detector)

Detector  
Position (n,l)  
Ag > (7,6)  
Br > (6,5)

Precise prediction of  $E^-$  tracks on the first emulsion plate.  
=> Reduction in background tracks which are picked up in the first plate.

## Volume of nuclear emulsion

70 liters (E373) => **210 liters**

## Track scanning method

Automated + Interactive (E373)  
=> **Fully Automated (speed x10)**

## (As for X-ray measurement)

X-ray data are only analyzed for **clear  $E^-$  stopping events** which are identified by nuclear emulsion.  
=> **quite low background**

# Other AGS Experiments - FY2005+

---

- **AGS Radiobiology** (*NASA*)
  - Completed last run in Nov 2002
  - Expect limited AGS use since NSRL is now operational
  - Facility will be displaced (move to Switchyard?) by E940 ( $\geq 2005$ )
  
- **P963 – Proton Radiography** (*National Nuclear Security Administration - NNSA*)
  - Multiple bunch extraction studies in FY 2002-3 (successful)
  - Proposal for dynamic testing under discussion with BNL management



# Final Comments

- The BNL priority for AGS Fixed target experiments is:
  - 1) RSVP (and finish E949 and E962 if possible)
  - 2) Long base line neutrino experiment to follow RSVP
- Experiments must make their case (again!) and devise a strategy to secure operations funds to be considered.
- Once RSVP construction plans are set (towards the end of this year) they will be difficult to make changes to accommodate other experiments.