



Overview of STAR Small System Correlations Results

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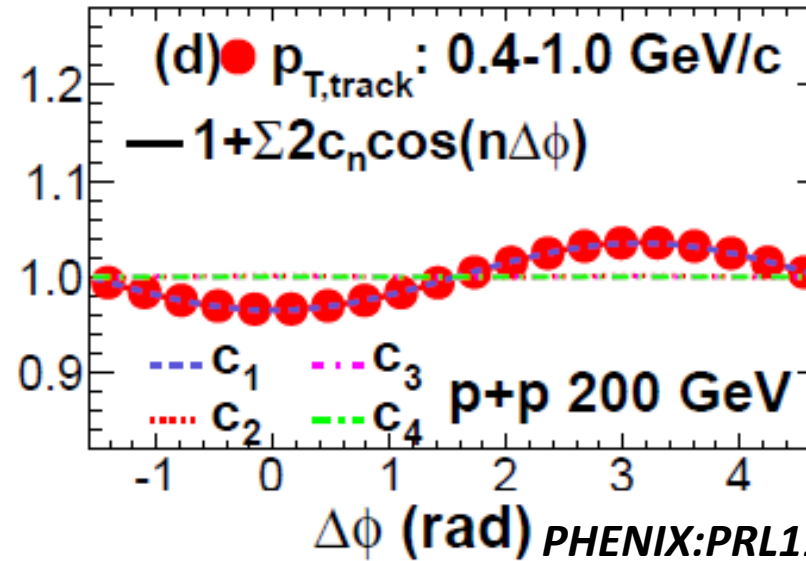
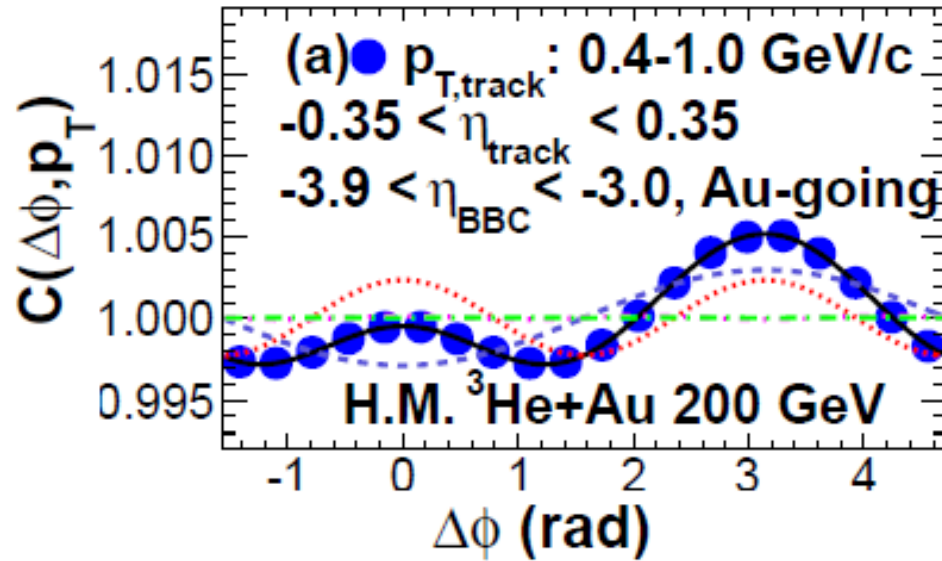




Outline

- Motivation and Analysis Methods
- Physics Results
 - ✓ The v_2 in different small systems at different collision energies
 - ✓ Model study for nonflow subtraction with AMPT
 - ✓ The $c_2\{4\}$ in d+Au collisions at different collision energies
- Summary

What is the Origin of Ridge?



PHENIX:PRL115(2015)142301

Ridge (a long-range near-side correlation) is observed in small systems at RHIC:
 Creation of a small QGP droplet or other mechanisms?

If a small QGP droplet is indeed created:
 How does the system evolve in a small QGP droplet?
 How about the dependence of multiplicity and collision energy for the flow?

Participant nucleon vs. parton Glauber model

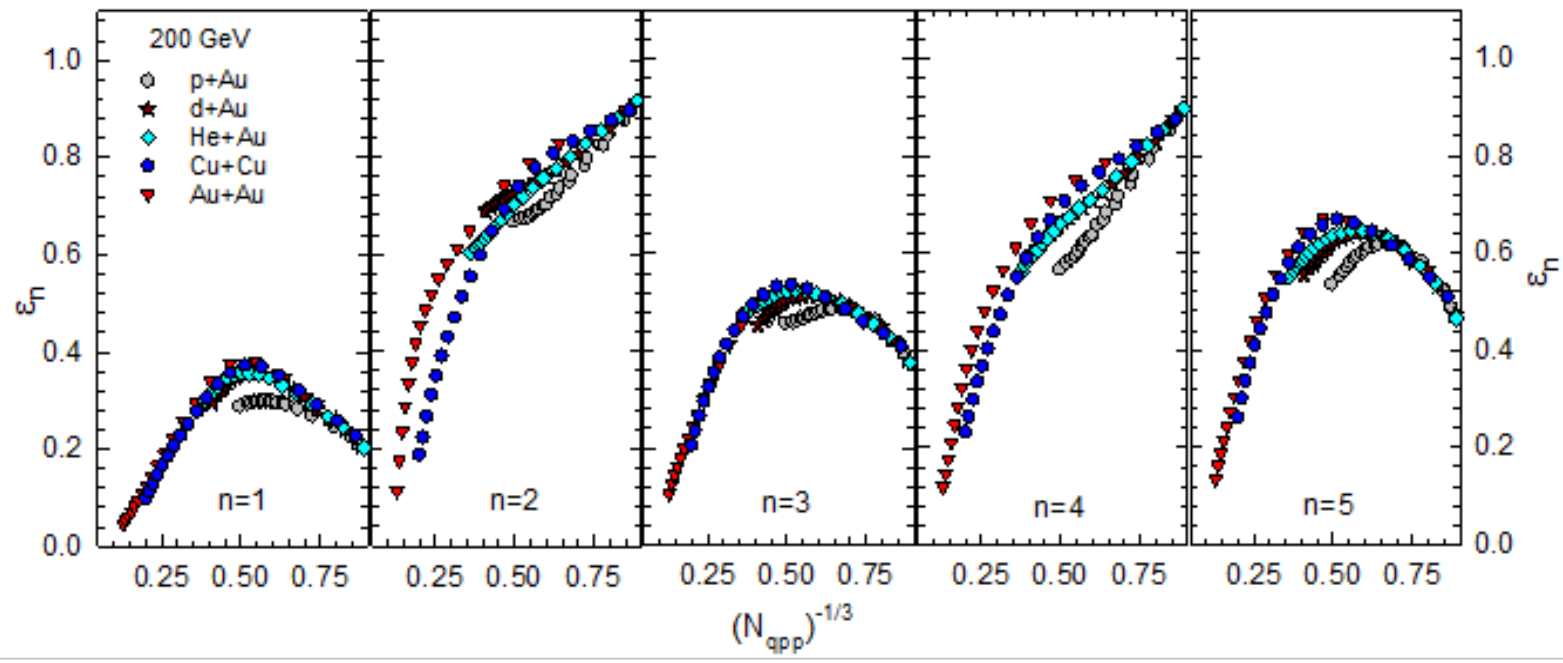


Participant nucleon

	0-5% p+Au	0-5% d+Au	0-5% He+Au
ϵ_2	0.23	0.54	0.50
ϵ_3	0.16	0.19	0.28

J. L. Nagle [PRL113\(2014\)112301](#)

Participant parton

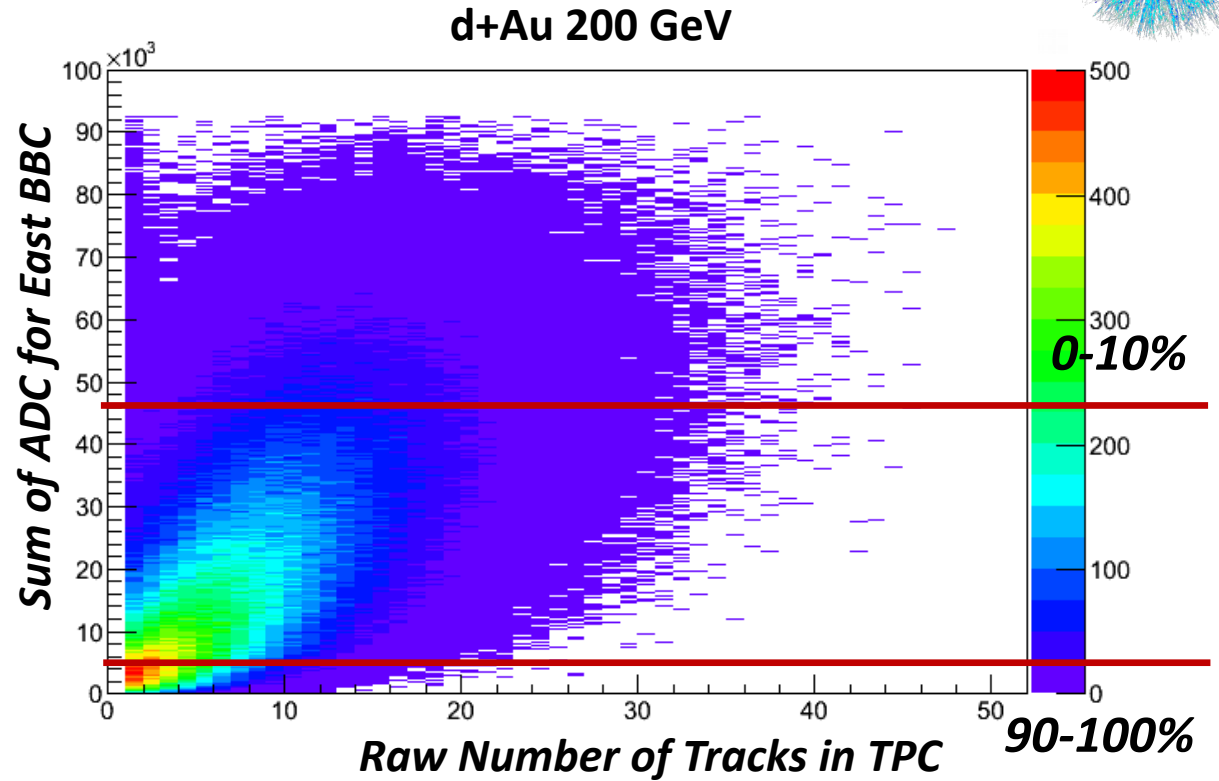
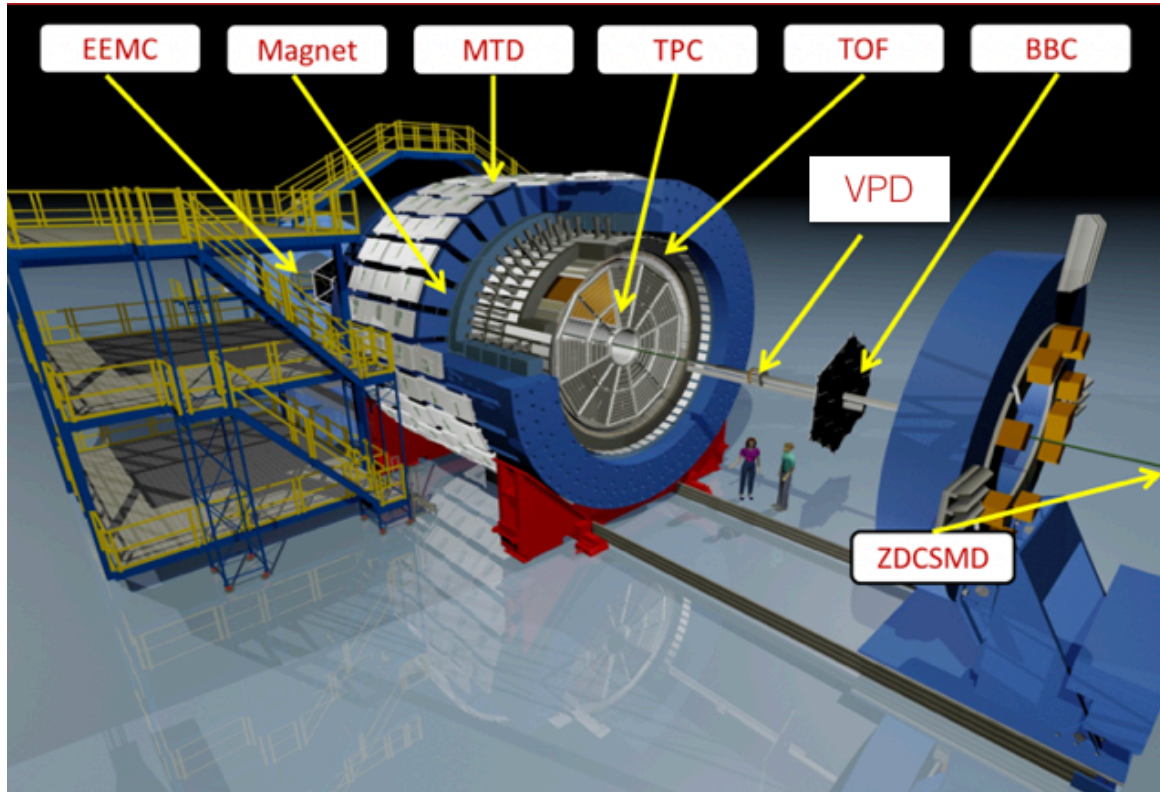


P. Liu arXiv:1804.04618

➤ Different initial geometry in small systems such as p/d/³He+Au

- **Similar eccentricity patterns for small and large systems in peripheral collisions.**
 - ✓ Trivial shape dependence for similar geometric size in central and mid-central collisions
 - ✓ Fluctuations important for small systems

Event Activity



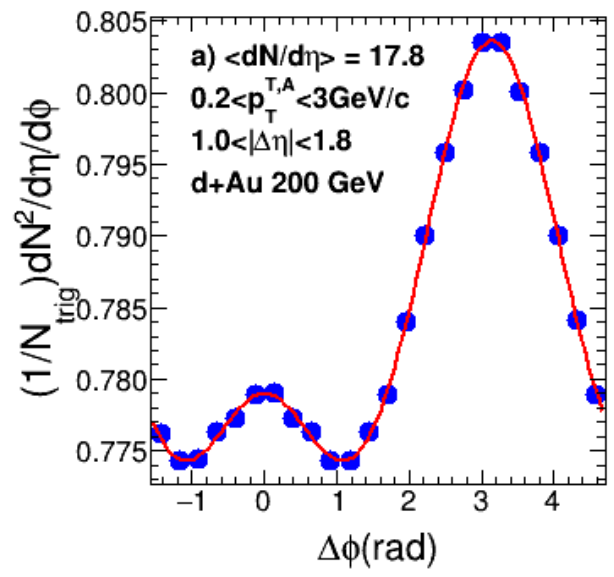
- ❑ Event classes with different activity are selected by using BBC east in the Au-going direction ($-5.0 < \eta < -3.3$)
- ❑ Long-range two-particle correlations are measured in TPC ($|\eta| < 0.9$)

- ❑ Correlation between multiplicity at backward and mid-rapidity
- ❑ 10 event classes with different TPC $\langle dN/d\eta \rangle$ are selected by sum ADC of BBC east in d+Au collisions at 200 GeV

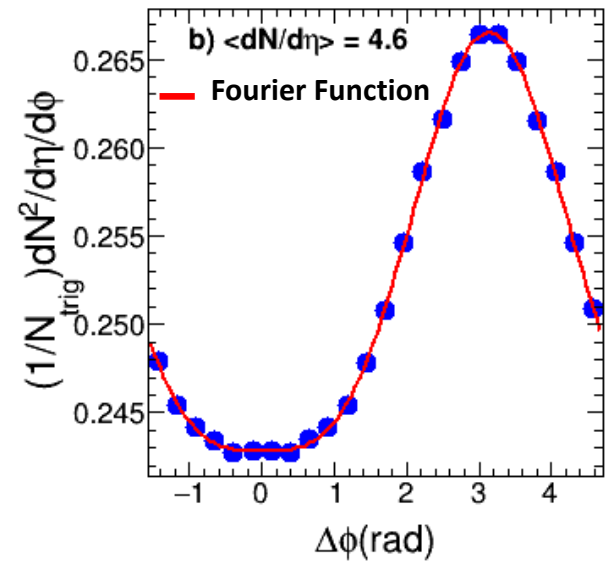
Long-range Two-particle Correlations



High Multiplicity (HM)



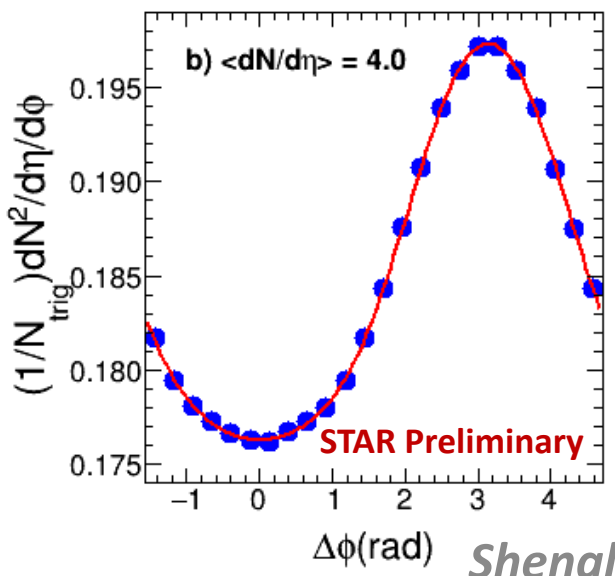
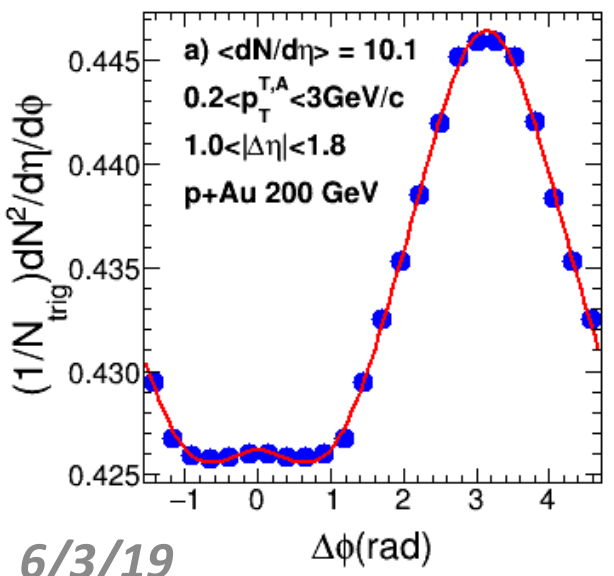
Low Multiplicity (LM)



d+Au 200 GeV

□ A near-side ridge is observed in the HM d+Au ($\langle dN/d\eta \rangle = 17.8$) and p+Au ($\langle dN/d\eta \rangle = 10.1$) collisions

□ A Fourier function is employed to extract the $V_{n,n}$



p+Au 200 GeV

$$dN/d\Delta\phi \sim 1 + \sum_{n=1}^4 2V_{n,n} \times \cos(n\Delta\phi)$$

$$\text{Integral } v_n = \text{sqrt}(V_{n,n}); v_n(p_T) = V_{n,n}(p_T)/v_n$$

Two Jet Subtraction Methods



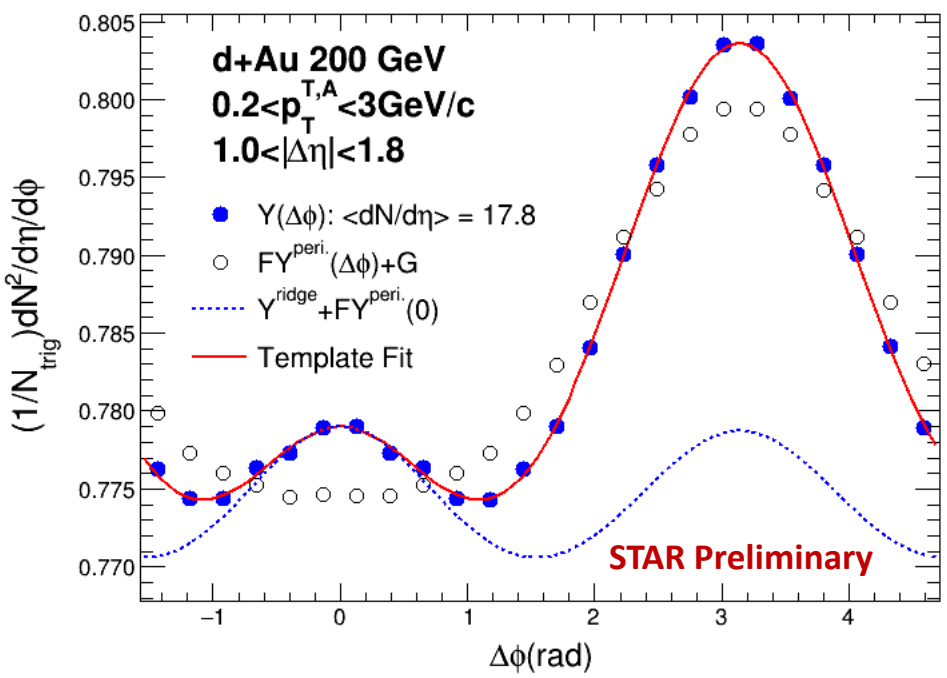
1. Low multiplicity subtraction scaled by short-range ($|\Delta\eta| < 0.5$) near-side jet yield

$$V_{n,n}^{HM}(\text{subtracted}) = V_{n,n}^{HM} - V_{n,n}^{LM} \times \frac{N_{asso}^{LM}}{N_{asso}^{HM}} \times \frac{Y_{jet,near-side}^{HM}}{Y_{jet,near-side}^{LM}}$$

ATLAS: PRC90(2014)044906
 CMS: PLB765(2017)193
 STAR: PLB743(2015)333

✓ Assumption: short-range near-side jet modification = long-range away-side jet modification

2. Template Fit



✓ A new developed method to subtract away-side jet contribution by ATLAS:

$$Y_{templ.}(\Delta\phi) = F \times Y_{LM}(\Delta\phi) + Y_{ridge}(\Delta\phi)$$

where

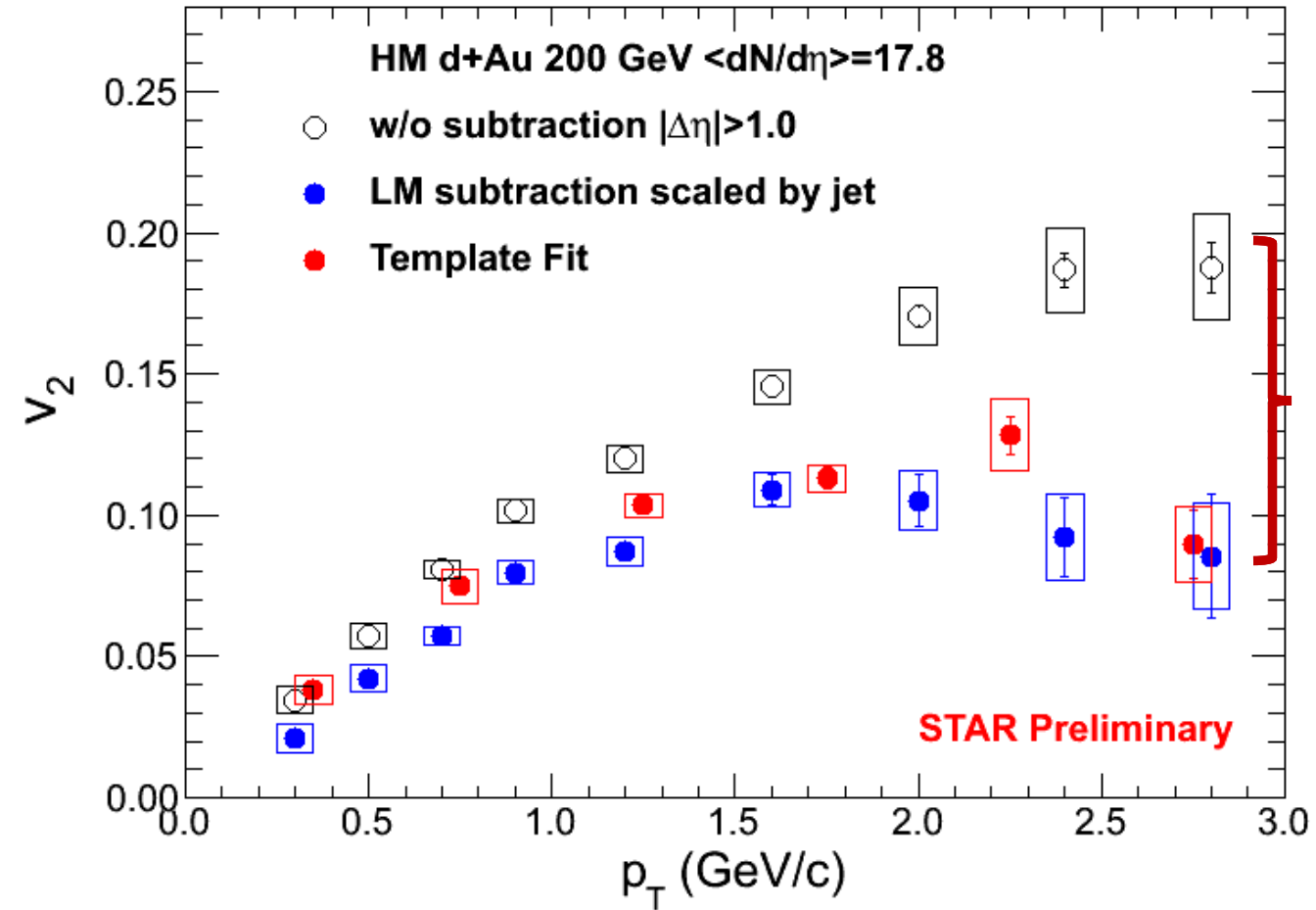
$$Y_{ridge}(\Delta\phi) = G \times (1 + 2 \times \sum_{n=2}^4 V_{n,n} \times \cos(n\Delta\phi))$$

ATLAS: PRL(116)172301

✓ Assumption: away-side jet shape can be measured in LM events and scaled by fit parameter "F" due to jet modification

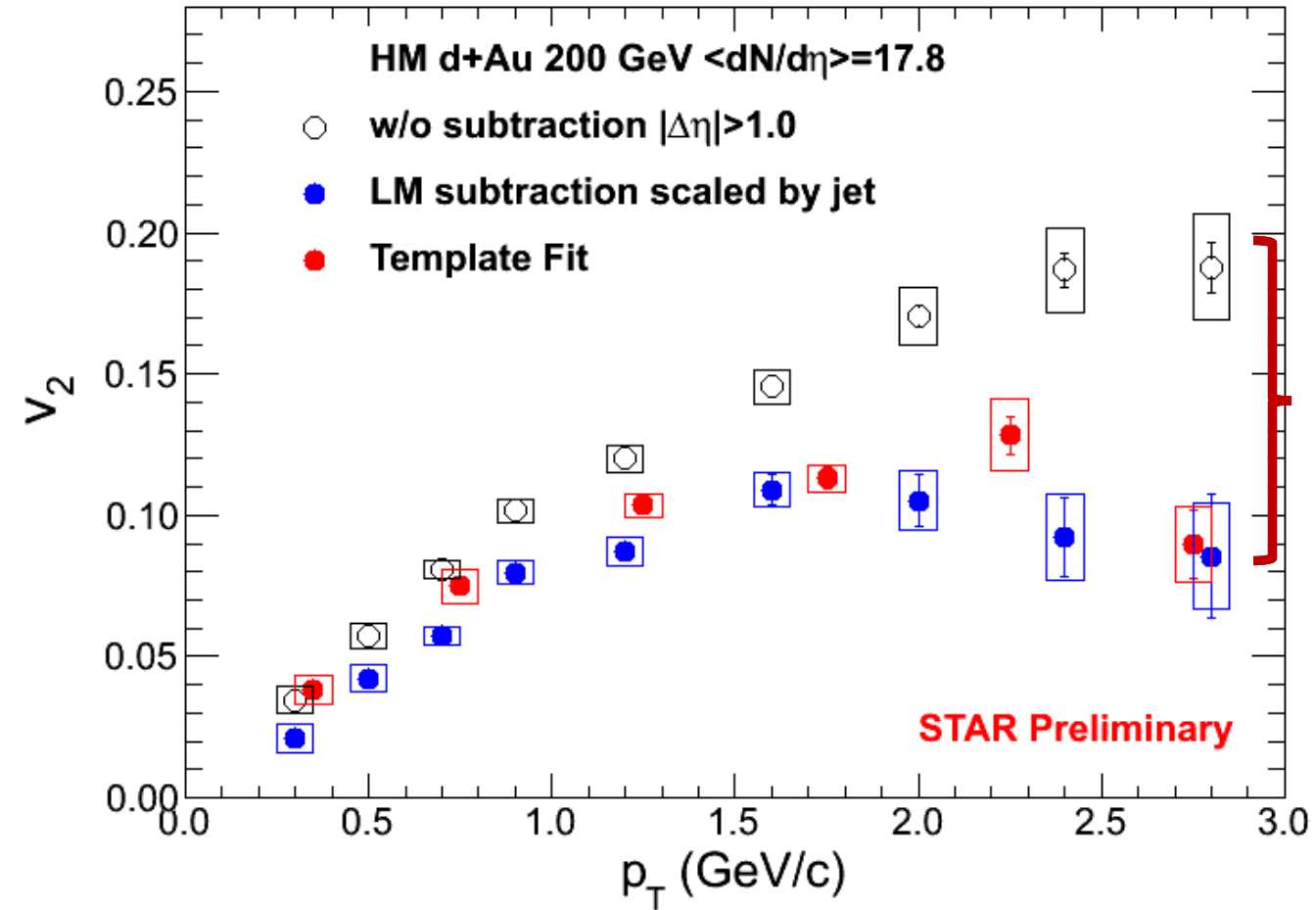
It will cause a bias if assumptions are not correct

v_2 in HM d+Au (0-10%) at 200 GeV



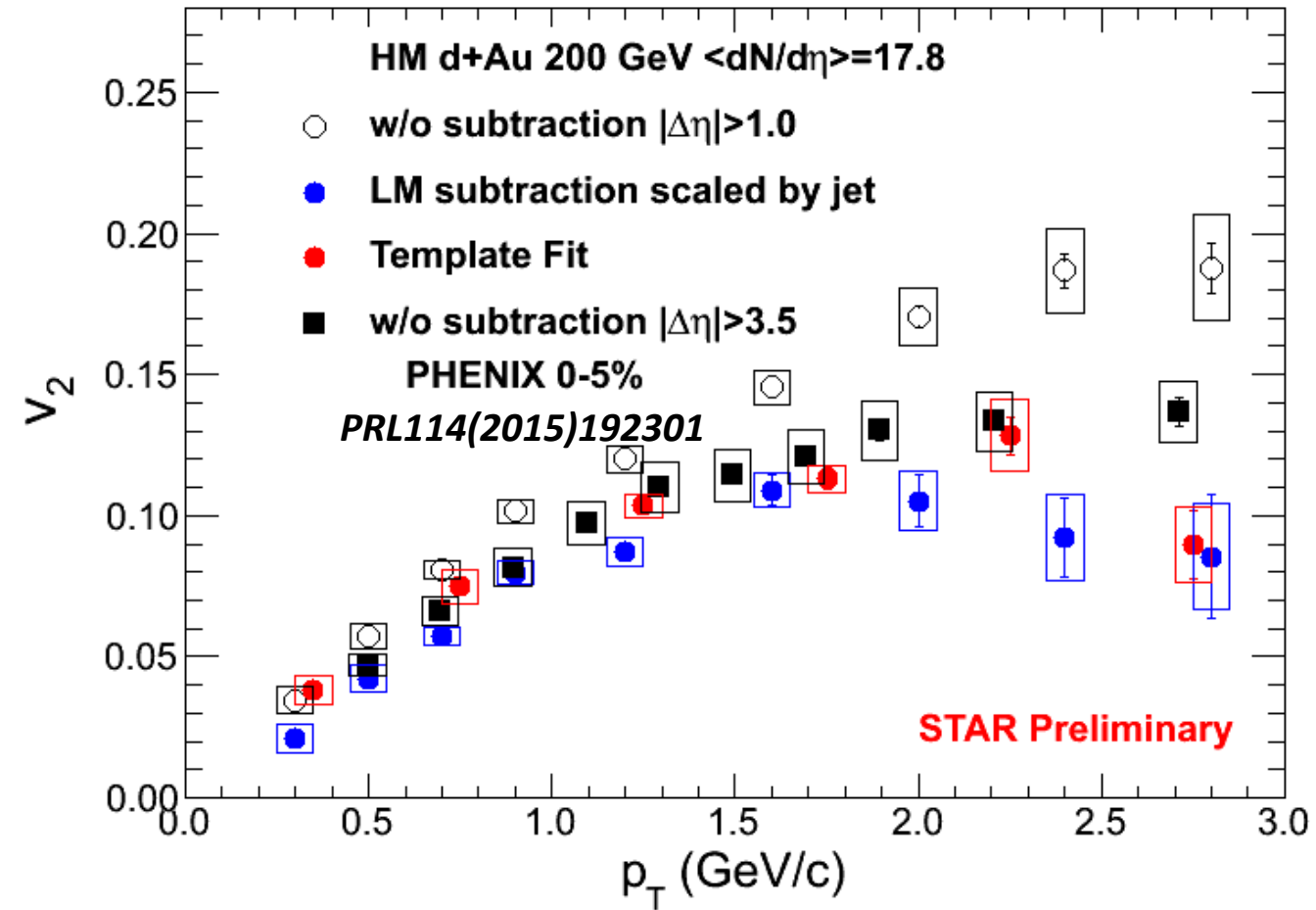
□ v_2 without subtraction is larger than that with subtraction for both methods. ***The subtraction of non-flow is crucial in small system!***

v_2 in HM d+Au (0-10%) at 200 GeV



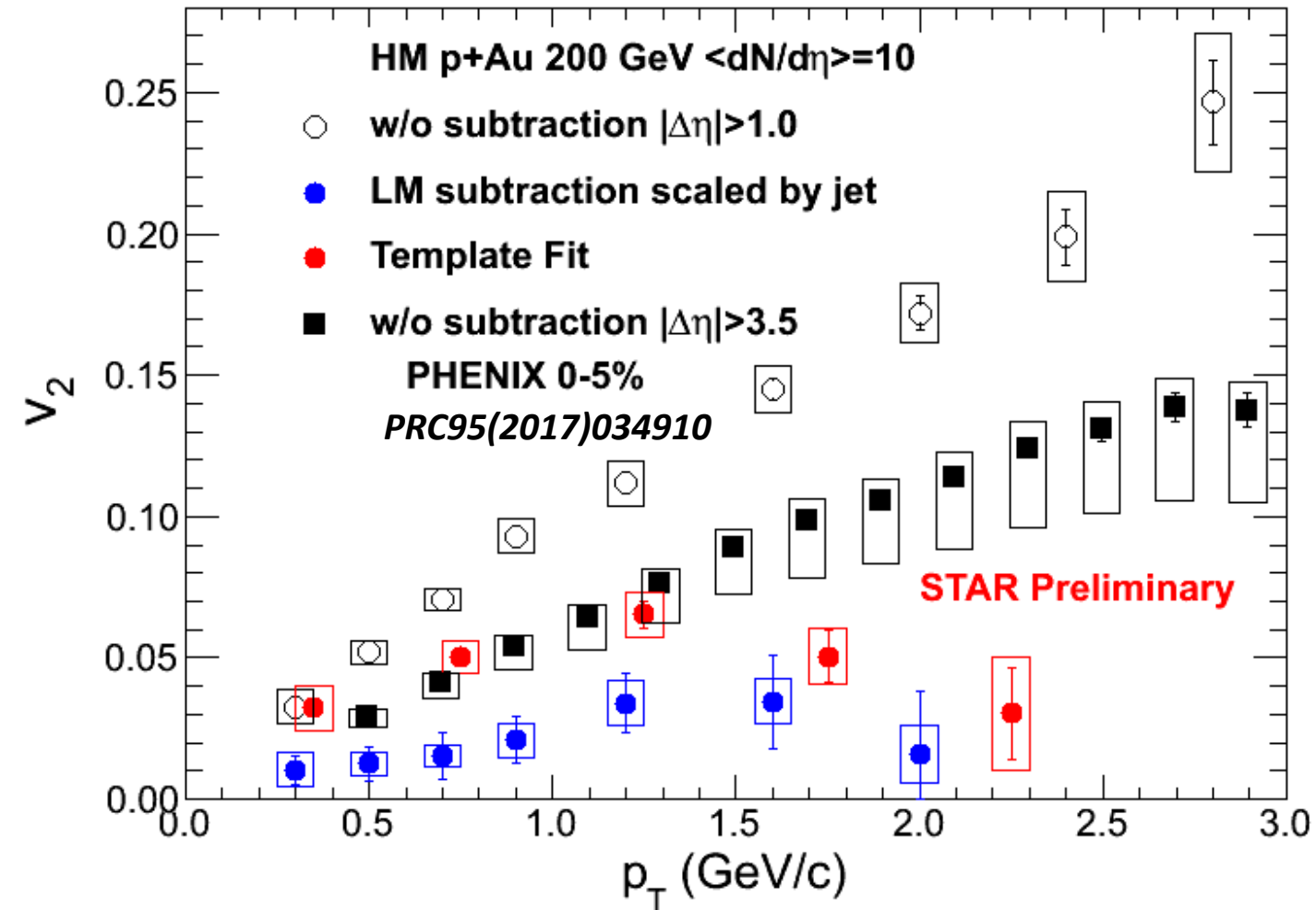
- v_2 without subtraction is larger than that with subtraction for both methods. ***The subtraction of non-flow is crucial in small system!***
- At lower p_T , the v_2 from LM subtraction is around 35% lower than that from template fit, while they are quite similar at intermediate p_T

v_2 in HM d+Au (0-10%) at 200 GeV



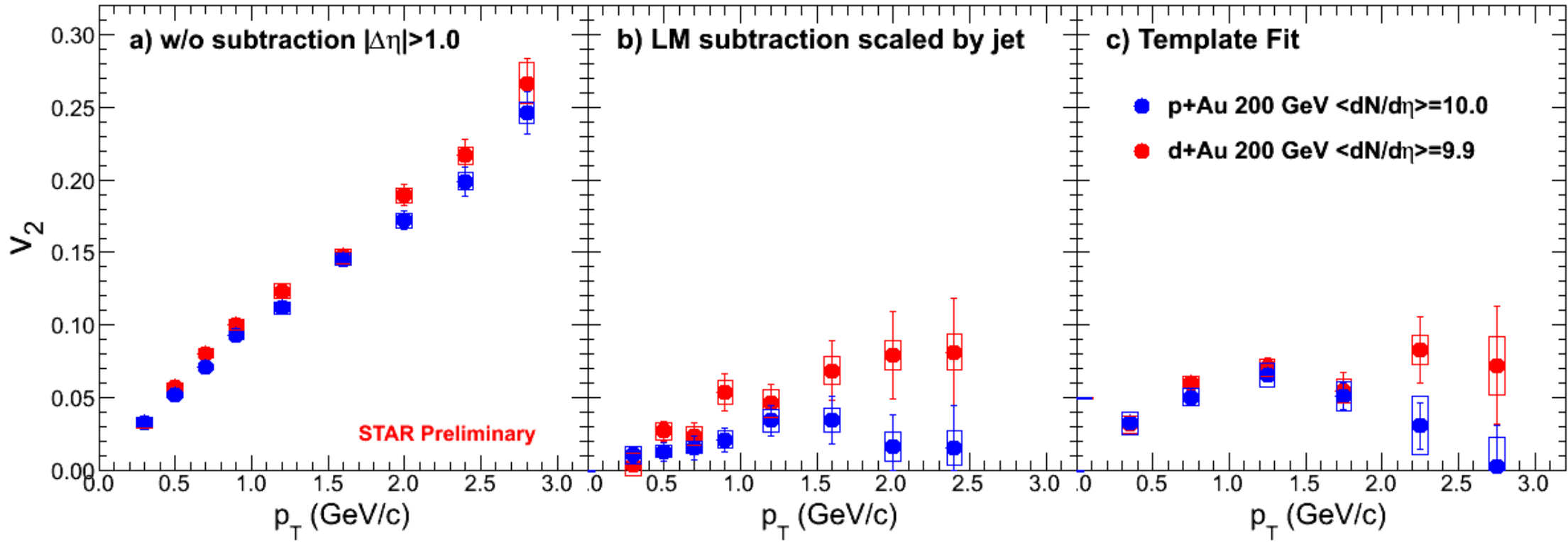
- v_2 without subtraction is larger than that with subtraction for both methods. **The subtraction of non-flow is crucial in small system!**
- At lower p_T , the v_2 from LM subtraction is around 35% lower than that from template fit. While they are quite similar at intermediate p_T
- The subtracted v_2 measured by STAR is similar to PHENIX measurement, which has at least 10% non-flow

v_2 in HM p+Au (0-10%) at 200 GeV



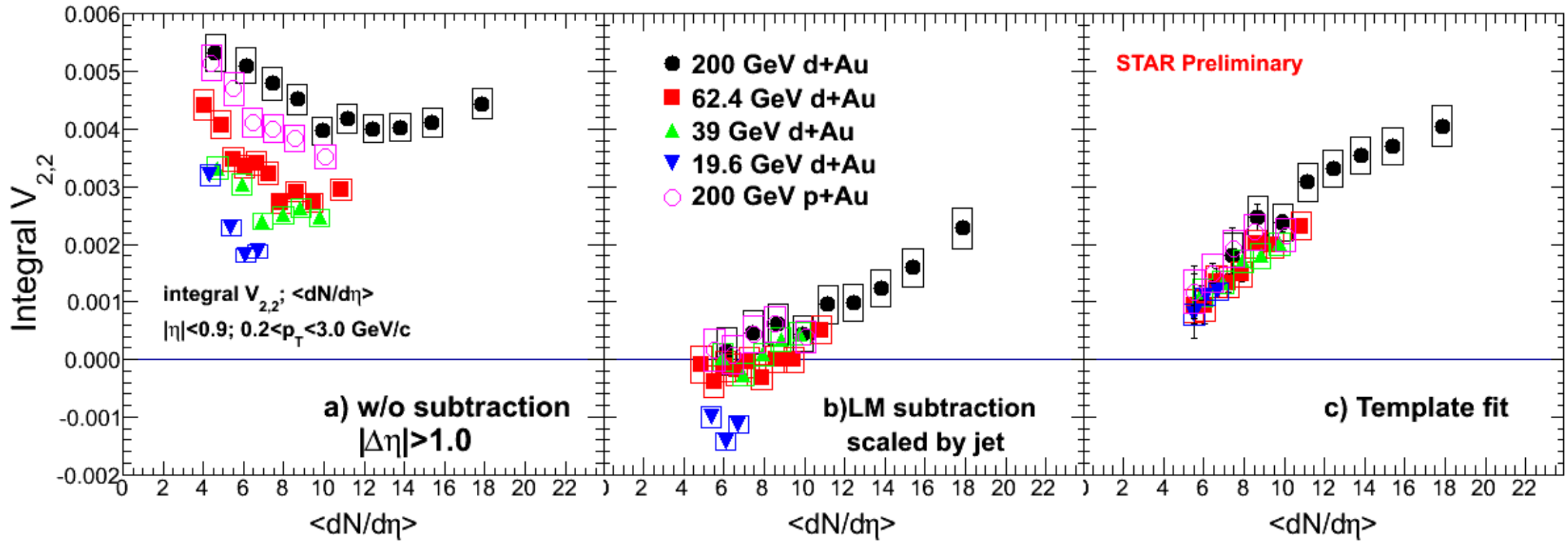
- Compared to d+Au results, v_2 in p+Au without subtraction is much larger than that with subtraction for two methods
- In p+Au collision, the v_2 from LM subtraction is much lower than that from template fit.
- v_2 from template fit method is similar to PHENIX measurement at low p_T

p/d+Au v_2 with same $\langle dN/d\eta \rangle$



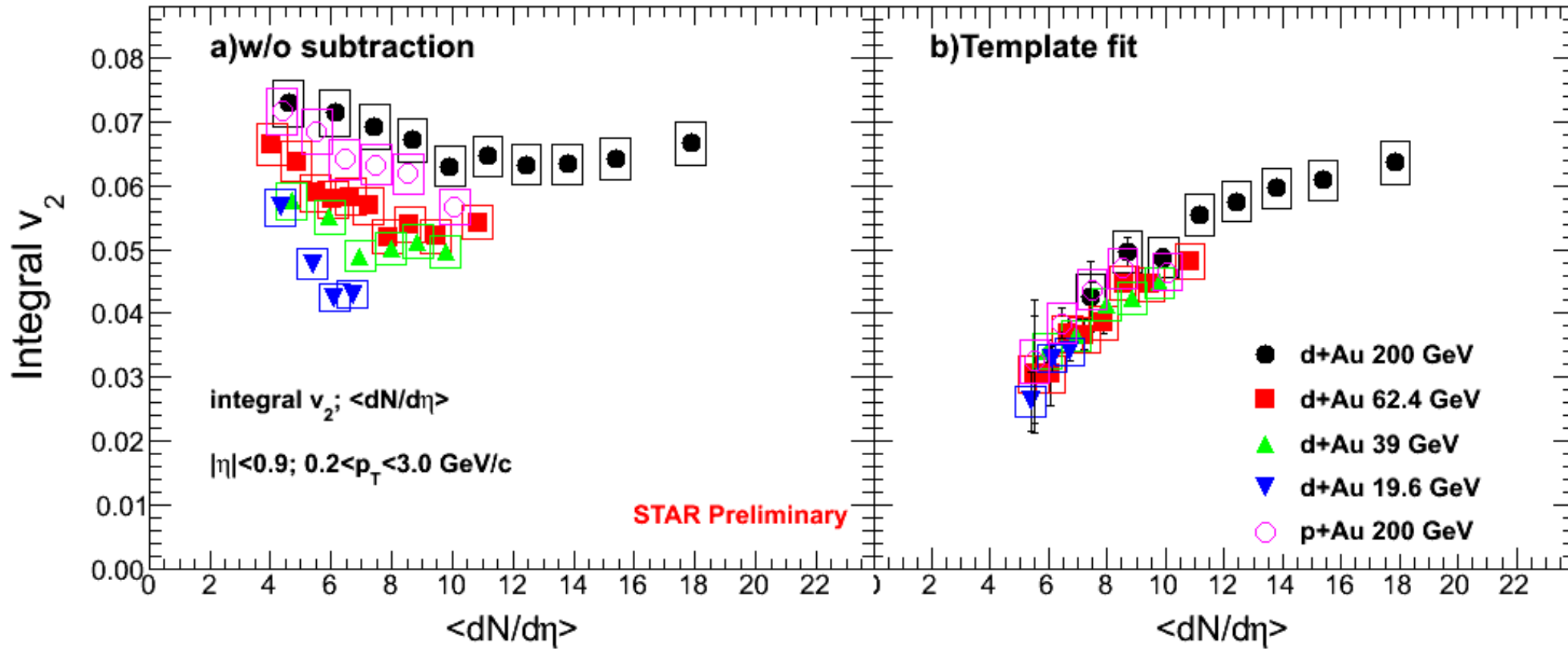
- ❑ By LM subtraction method, v_2 in d+Au is a little bit larger than that of p+Au collisions
- ❑ v_2 between p+Au and d+Au collisions from template fit is similar, while the initial eccentricities are different by a factor of two

Integral $V_{2,2}$ vs. $\langle dN/d\eta \rangle$



- There is large difference between two methods
- LM subtraction leads to a negative $V_{2,2}$ at low energy
 - ✓ Different kinematics between near- and away-side jet-like correlations?
- $V_{2,2}$ from template fit increases as a function of $\langle dN/d\eta \rangle$

Integral v_2 from Template Fit



- ❑ The unsubtracted integral v_2 as a function of $\langle dN/d\eta \rangle$ is different in different systems at different collision energies
- ❑ The integral v_2 from template fit shows a universal trend as a function of $\langle dN/d\eta \rangle$

Modified Template Fit



Template Fit:

$$Y_{templ.}(\Delta\phi) = F \times Y_{peri.}(\Delta\phi) + Y_{ridge}(\Delta\phi)$$

where

$$Y_{ridge}(\Delta\phi) = G \times (1 + 2 \times \sum_{n=2}^4 V_{n,n}^{templ.} \times \cos(n\Delta\phi))$$

Modified Template Fit



Template Fit:

$$Y_{templ.}(\Delta\phi) = F \times Y_{peri.}(\Delta\phi) + Y_{ridge}(\Delta\phi)$$

where

$$Y_{ridge}(\Delta\phi) = G \times (1 - 2 \times \sum_{n=2}^4 V_{n,n}^{templ.} \times \cos(n\Delta\phi))$$

The associate particles under the pedestal can also evolve into the flow

The yield of such part of particles will be $2\pi F Y_{peri.}(0)$

It will give large contribution if G is not much larger than $F Y_{peri.}(0)$

Modified Template Fit



Template Fit:

$$Y_{templ.}(\Delta\phi) = F \times Y_{peri.}(\Delta\phi) + Y_{ridge}(\Delta\phi)$$

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$$Y_{ridge}(\Delta\phi) = G \times (1 + 2 \times \sum_{n=2}^4 V_{n,n}^{templ.} \times \cos(n\Delta\phi)) + F \times Y_{peri.}(0)$$

Modified Template Fit



Template Fit:

$$Y_{templ.}(\Delta\phi) = F \times Y_{peri.}(\Delta\phi) + Y_{ridge}(\Delta\phi)$$

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The associate particles under the pedestal can also evolve into the flow

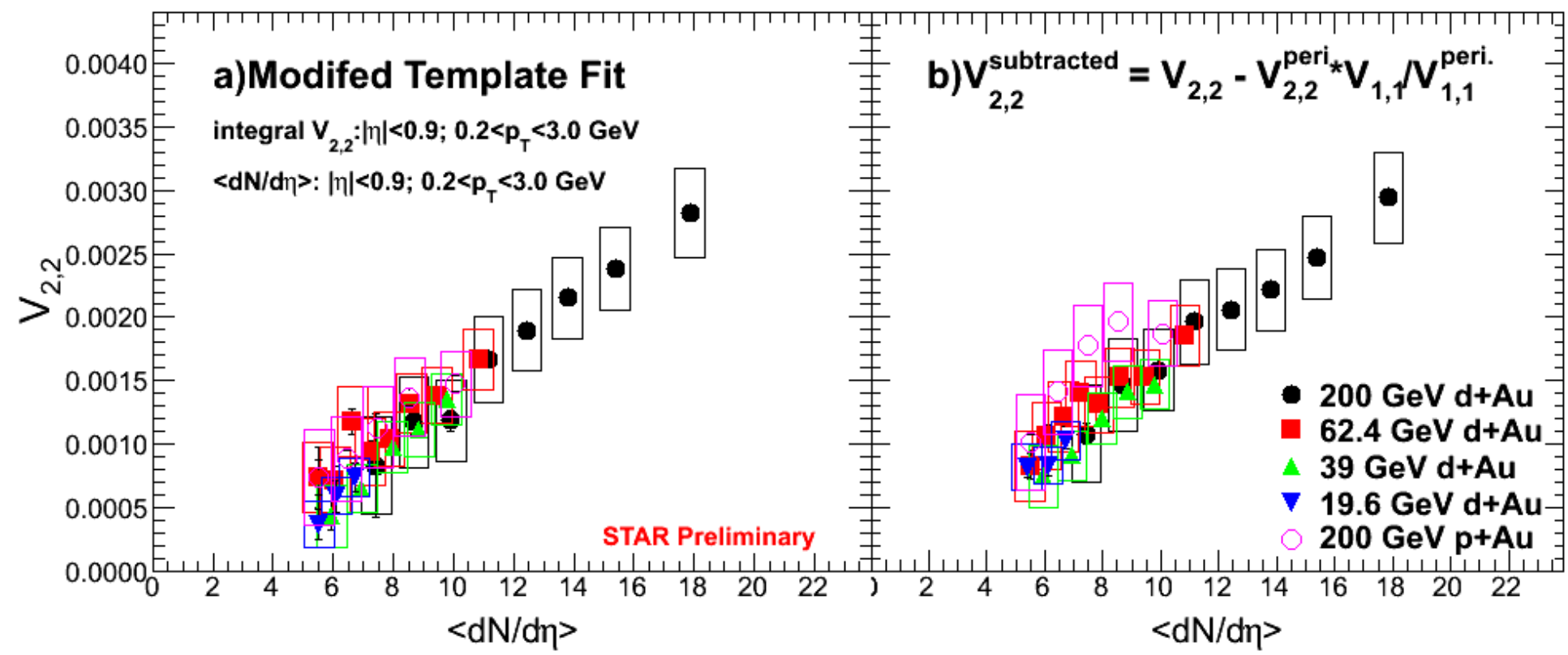
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$$Y_{ridge}(\Delta\phi) = G \times (1 + 2 \times \sum_{n=2}^4 V_{n,n}^{templ.} \times \cos(n\Delta\phi)) + F \times Y_{peri.}(0)$$

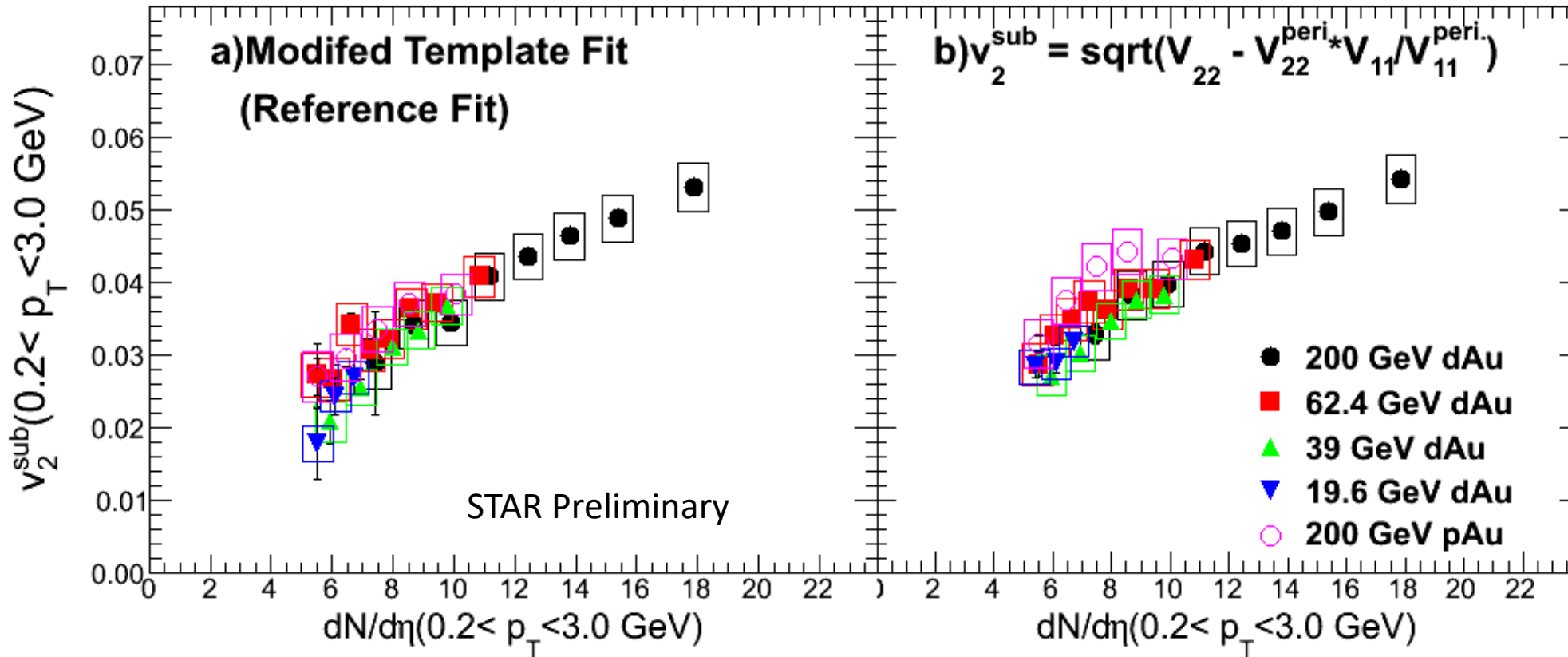
$$V_{22}^{Modified} = V_{22}^{templ.} \times G / (G + F Y_{peri.}(0))$$

Modified Template Fit



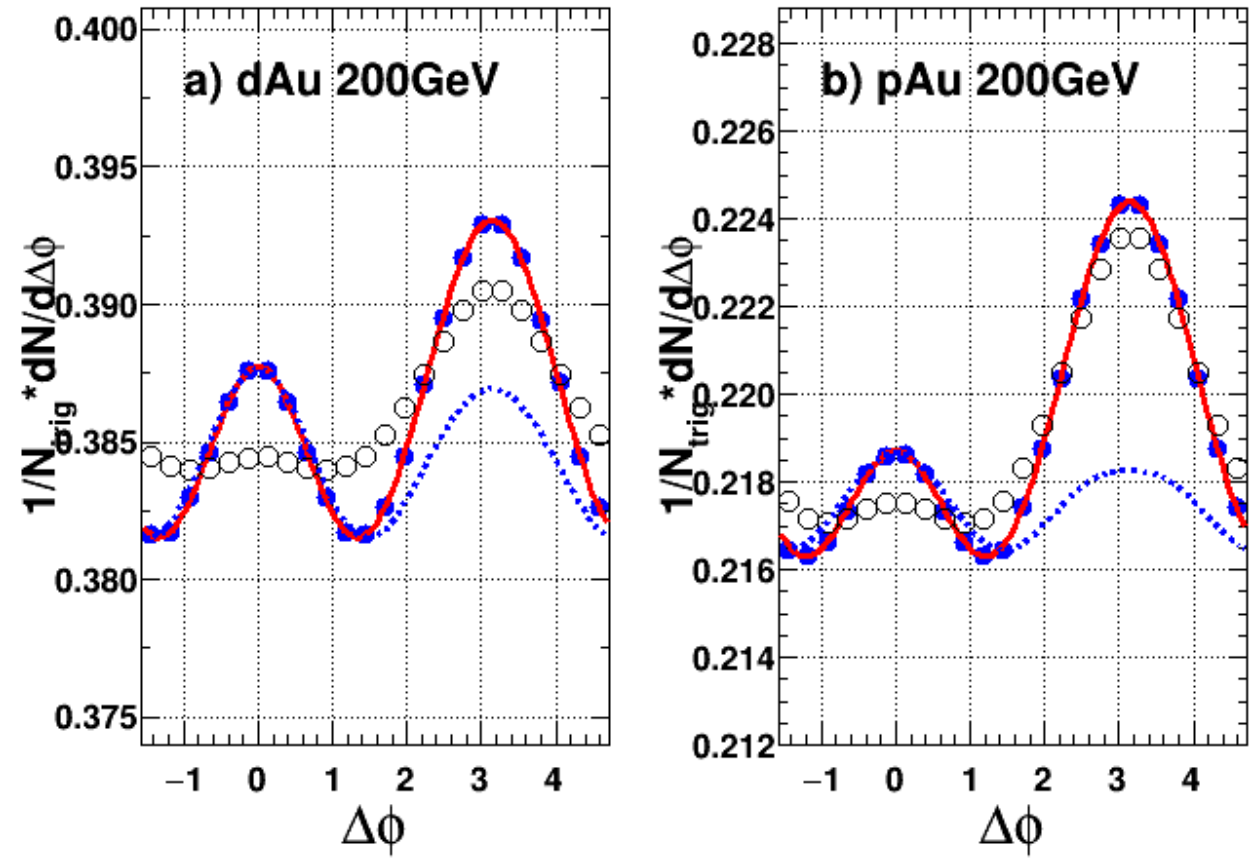
- Using the modified template fit, $V_{2,2}$ shows a universal linear trend as a function of $\langle dN/d\eta \rangle$ for different systems and collision energies
- The results are similar to the $V_{2,2}$ after peripheral subtraction scaled by the $V_{1,1}$

Integral v_2 from modified template fit



- The integral v_2 from two subtraction methods show a universal trend as a function of $\langle dN/d\eta \rangle$

Nonflow subtraction in AMPT



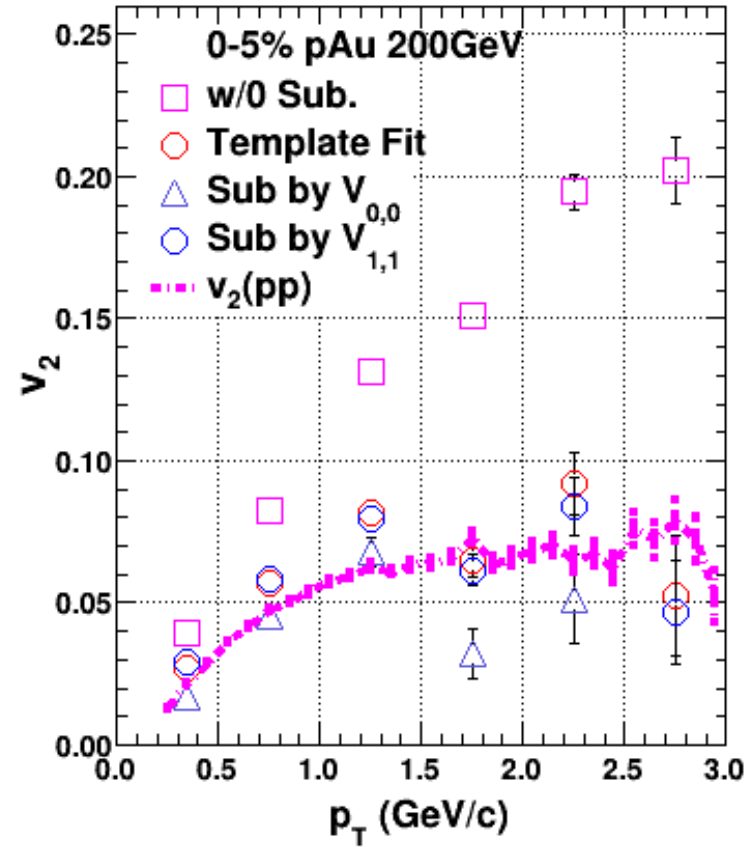
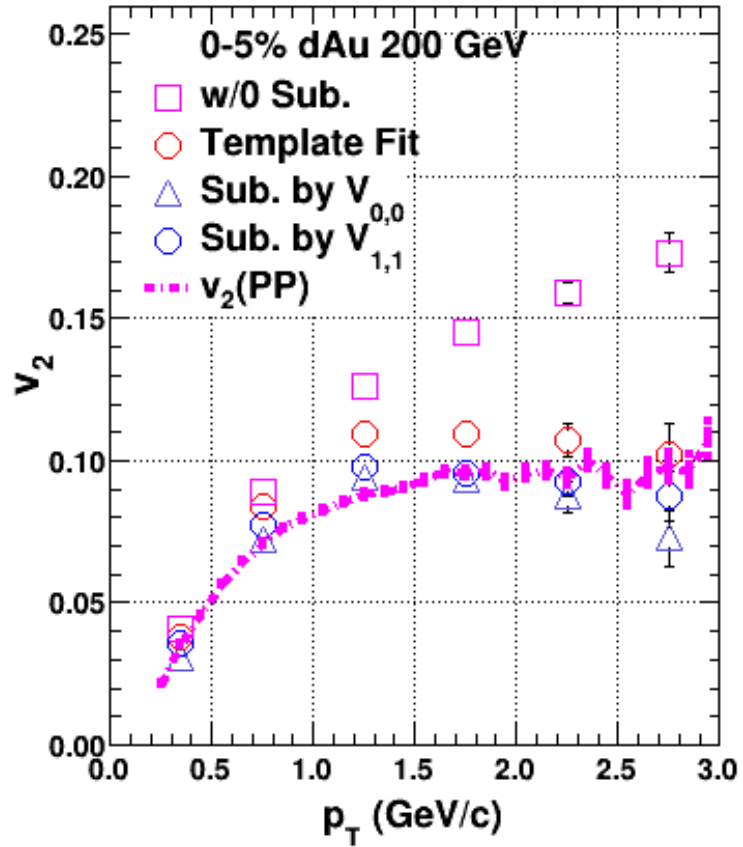
$|\Delta\eta| > 1.0, 0.2 < p_T(\text{trig,asso.}) < 3.0$

- 0-5% p/d+Au
- $FY^{pp}+G$
- $Y^{\text{ridge}}+FY^{pp}(0)$
- Template Fit

pp as reference, a near side peak is shown even $|\Delta\eta| > 1.0$

0-5% centrality is selected with particles $-5.0 < \eta < -3.0$

$v_2(p_T)$ in 0-5% pAu, dAu at 200GeV



$$V_2(pp) = \langle \cos 2(\phi - \psi_{PP}) \rangle / F$$

$$F = \sqrt{\frac{\langle \cos 2(\psi_{PP} - \psi_{EP,A}) \rangle \langle \cos 2(\psi_{PP} - \psi_{EP,B}) \rangle}{\langle \cos 2(\psi_{EP,A} - \psi_{EP,B}) \rangle}}$$

ψ_{PP} : participant plane

$\psi_{EP,A}$: event plane from particles $-4.5 < \eta < -2.5$

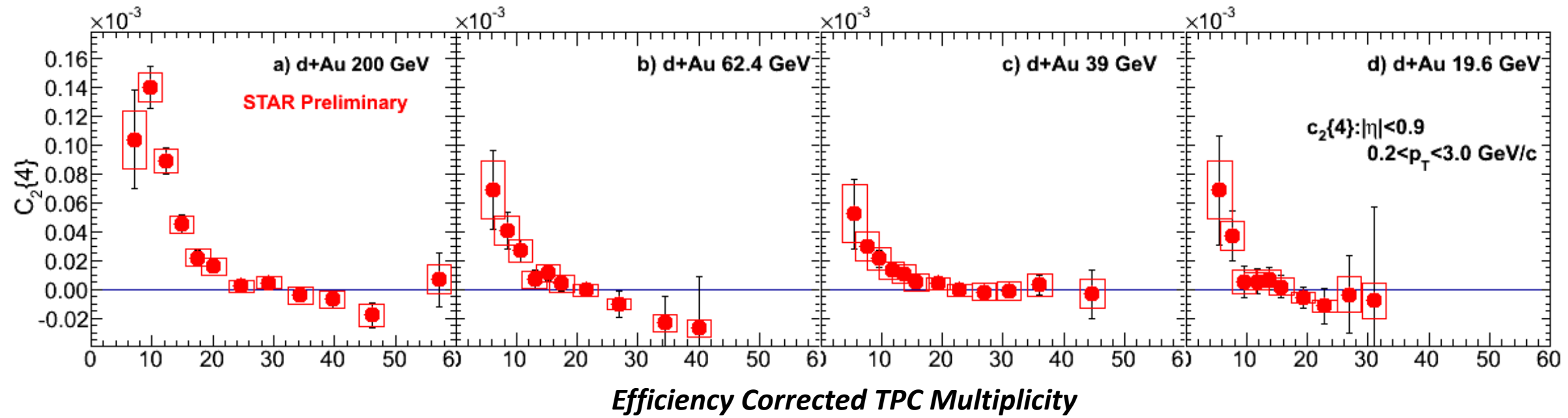
$\psi_{EP,B}$: event plane from particles $0 < \eta < 2.5$

Results from different subtraction are similar

The subtracted $v_2(p_T)$ have same trend as $v_2(PP)$ as a function of p_T

The subtraction methods works well even with $|\Delta\eta| > 1.0$

$c_2\{4\}$ vs. $\langle dN/d\eta \rangle$



Four-Particle Cumulant

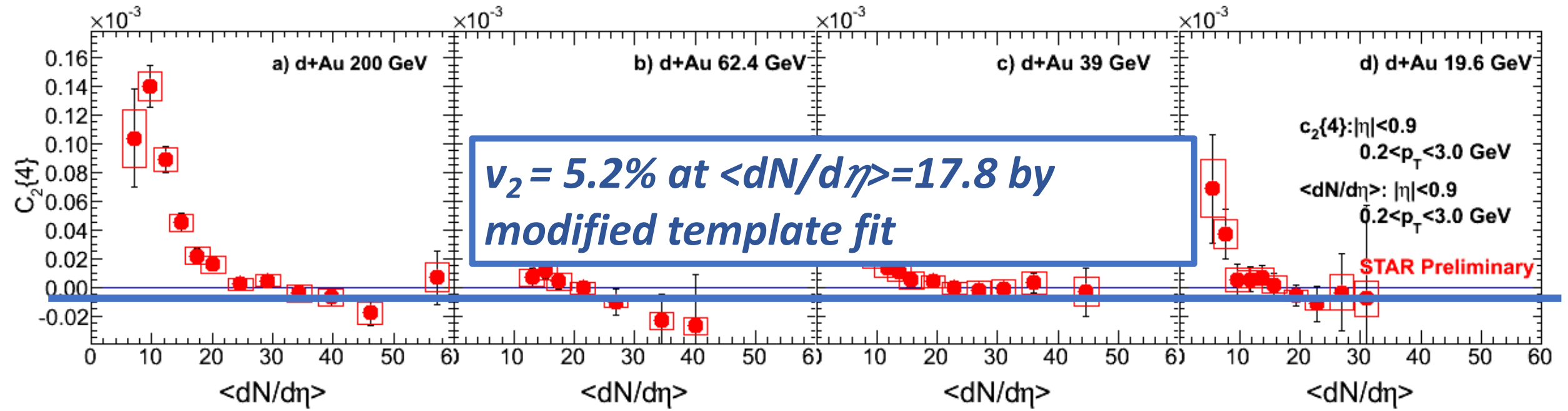
$$c_2\{4\} = \langle\langle e^{-i2(\phi_i + \phi_j - \phi_k - \phi_l)} \rangle\rangle - 2 \langle\langle e^{-i2(\phi_i - \phi_j)} \rangle\rangle$$

$\phi_i, \phi_j, \phi_k, \phi_l$ are the azimuthal angles of four different particles in an event ; $\langle\langle \rangle\rangle$ represents the average over all particles from all events within a given multiplicity range

$$v_2\{4\} = \sqrt[4]{-c_2\{4\}}$$

An indication that $c_2\{4\}$ is negative for high multiplicity d+Au collisions at 200 and 62.4 GeV, while the statistical uncertainties are large

$c_2\{4\}$ vs $\langle dN/d\eta \rangle$

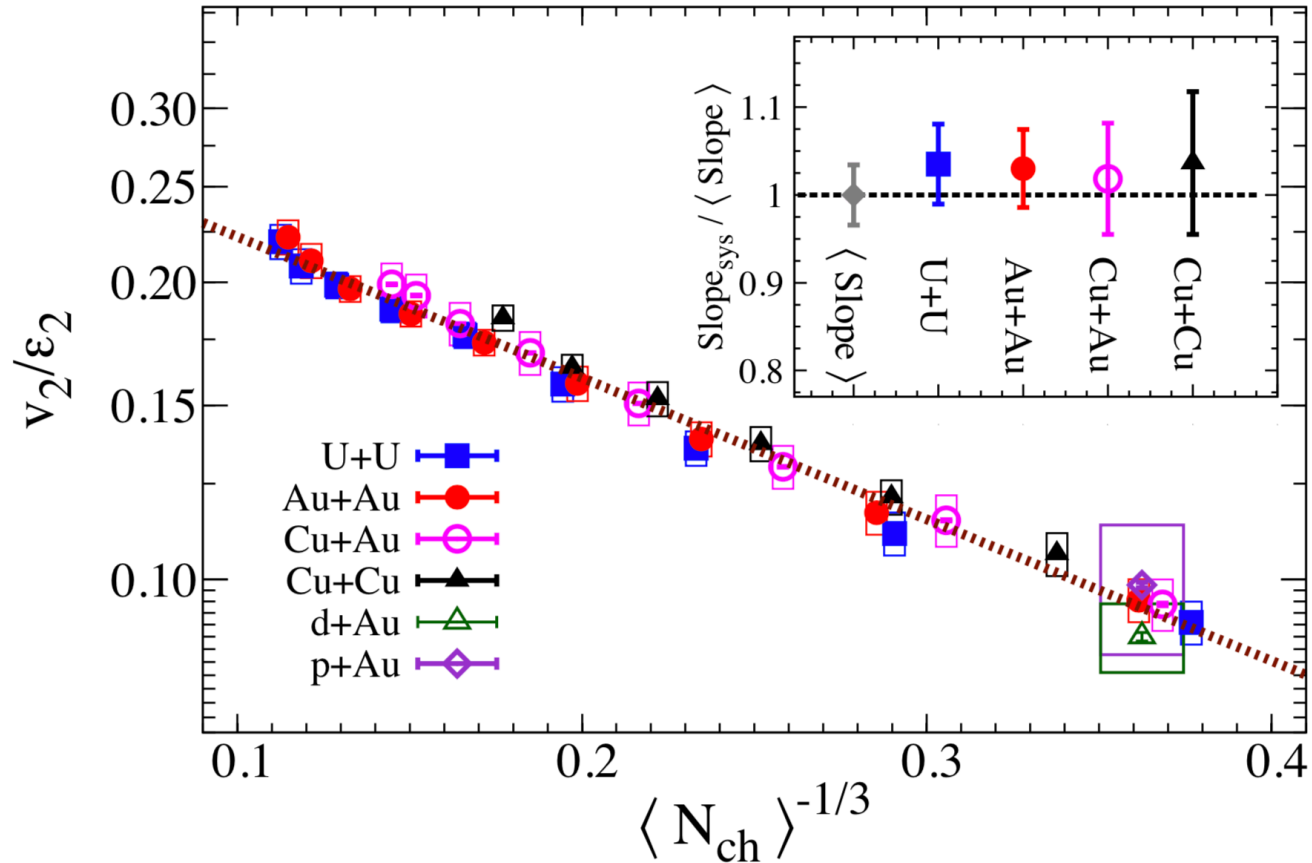


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An indication that $c_2\{4\}$ is negative for high multiplicity in d+Au collisions at 200 and 62.4 GeV

$$v_2\{4\} = \sqrt[1/4]{-c_2\{4\}}$$

From large to small system



A universal scaling from large to small system. Driven by same physics?

Summary

We see similar v_2 between p/d+Au collisions for same multiplicity.

✓ v_2 is not only driven by initial geometry

The integral v_2 extracted by template fit shows a universal trend as a function of $\langle dN/d\eta \rangle$ for different small systems at different energies. v_2 in large and small systems follow an universal trend

✓ Multiplicity plays an important role for the flow in small systems!

$c_2\{4\}$ is negative at high multiplicity at 62.4 and 200 GeV, but the measurements are limited by statistics.

Comparison of v_2 between two and four particles correlation, and testing of nonflow subtraction in AMPT, both indicate that nonflow is well controlled in STAR